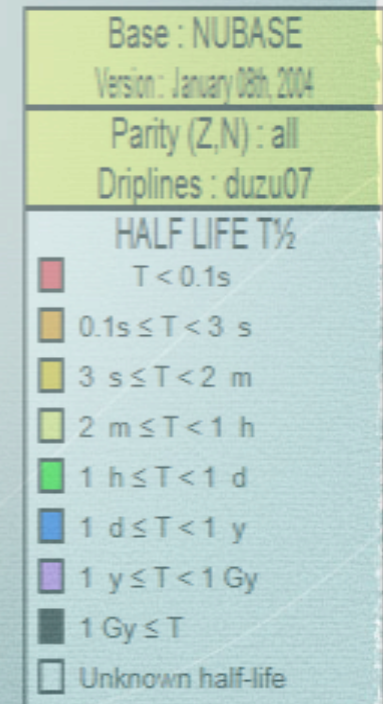


Beta Delayed Particle Emission

Ismael Martel
Department of Applied Physics
University of Huelva
Huelva (Spain)



Huelva



Cristobal Columbus
1492



Recreativo
Football Club
1889



1960's



1992

History:

→1492, the discovery of America: Shipbuilders, Caravels and the crew of Cristobal Columbus were from Huelva. Depart from a small port located at the village of Palos de la Frontera (Huelva).

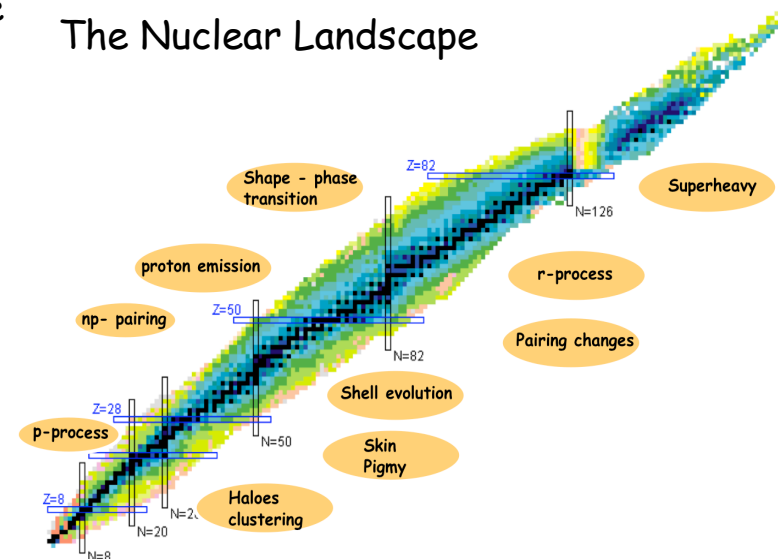
→1889, first football team in Spain (soccer) founded by British workers at "Rio Tinto" mines (Rio Tinto Company, London, 1873).

→1960's, one of largest industrial sites in Spain (Chemicals, Petrol & Mining industry)

→1992, University of Huelva was born, one of the youngest Universities of Spain. (15.000 students/ 150.000 habitants of Huelva)

→1999, the Nuclear & Particle Physics group (Estructura de la Materia)

The Nuclear Landscape



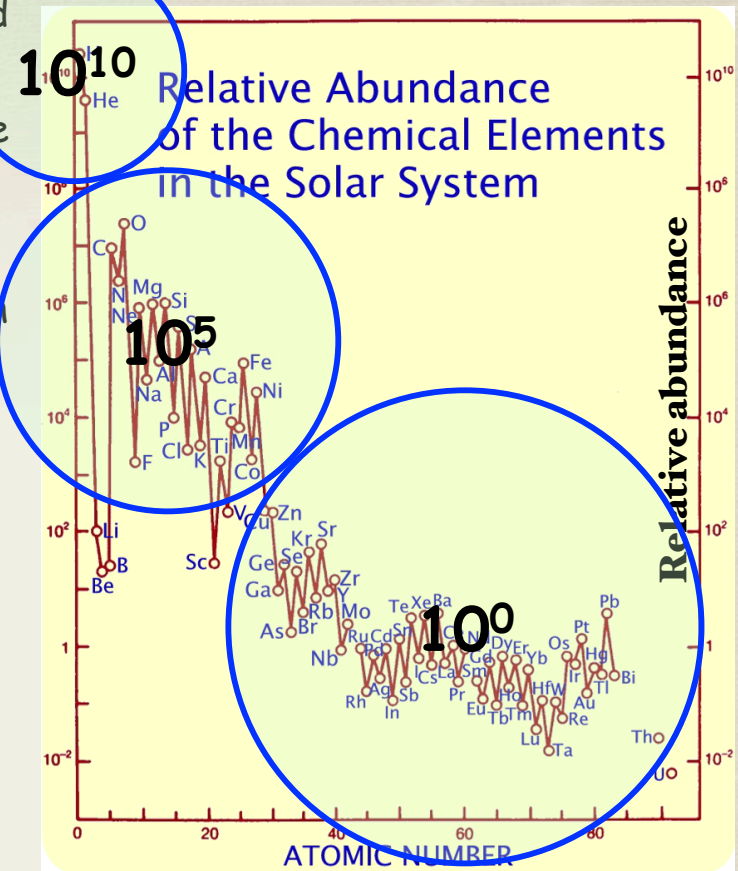
Introduction

Along history, there has been a constant effort to understand the structure and mechanism of the nature that surround us:

- Why the Universe and the Nature have the structure we observe?
- Which are the basic constituents of matter?
- How the different building blocks of matter interact with each other?
- Where, when and how the Universe has been originated?



“Creation pillars”, nucleosynthesis of stars at Eagle’s Nebulae

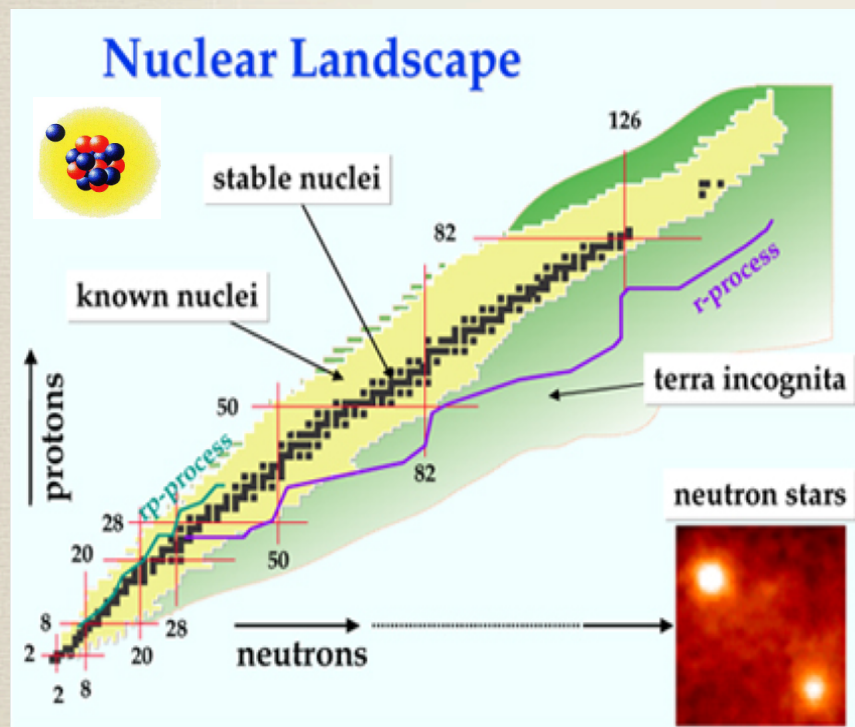


The research efforts carried out in basic nuclear physics (and Science in a wide sense) along last century (XX) has provided an un-precedent knowledge of the subatomic structure of matter and its constituents, its dynamics and the Origin of the Universe itself.

From a historical point of view, the major steps in the understanding of the Universe have taken place in **particle accelerators**.

At present **Radioactive Beam Facilities** we can customize our nuclear system (N,Z), "fabricate" any nucleus controlling the number of constituent protons and neutrons.

Proton Rich Nuclei $\leftarrow \rightarrow$ **Neutron Rich Nuclei** $\leftarrow \rightarrow$ **Light unbound systems** $\leftarrow \rightarrow$ **Super-heavy's**



\rightarrow **Evolution** of nuclear structure and nuclear dynamics,

\rightarrow **Exotic (N,Z) combinations** \rightarrow isospin degree of freedom

- **Evolution of shell structure**, phase shape transitions, nucleon-nucleon pairing, spin-orbit interaction
- **Halo**, skin, cluster nuclear structures
- **Beyond** the drip lines \rightarrow unbound nuclei & resonances
- **Exotic decay modes** and Reaction dynamics of exotic systems
- **Test of astrophysical** scenarios \rightarrow nuclear astrophysics

Spectroscopic tools \rightarrow Particle Detectors + Accelerators

Theoretical tools: Precise knowledge of theoretical framework **well tested** with stable nuclei \rightarrow
Example: **DIRECT NUCLEAR REACTIONS** (FRESCO, ECIS, etc,...)

The nuclear landscape

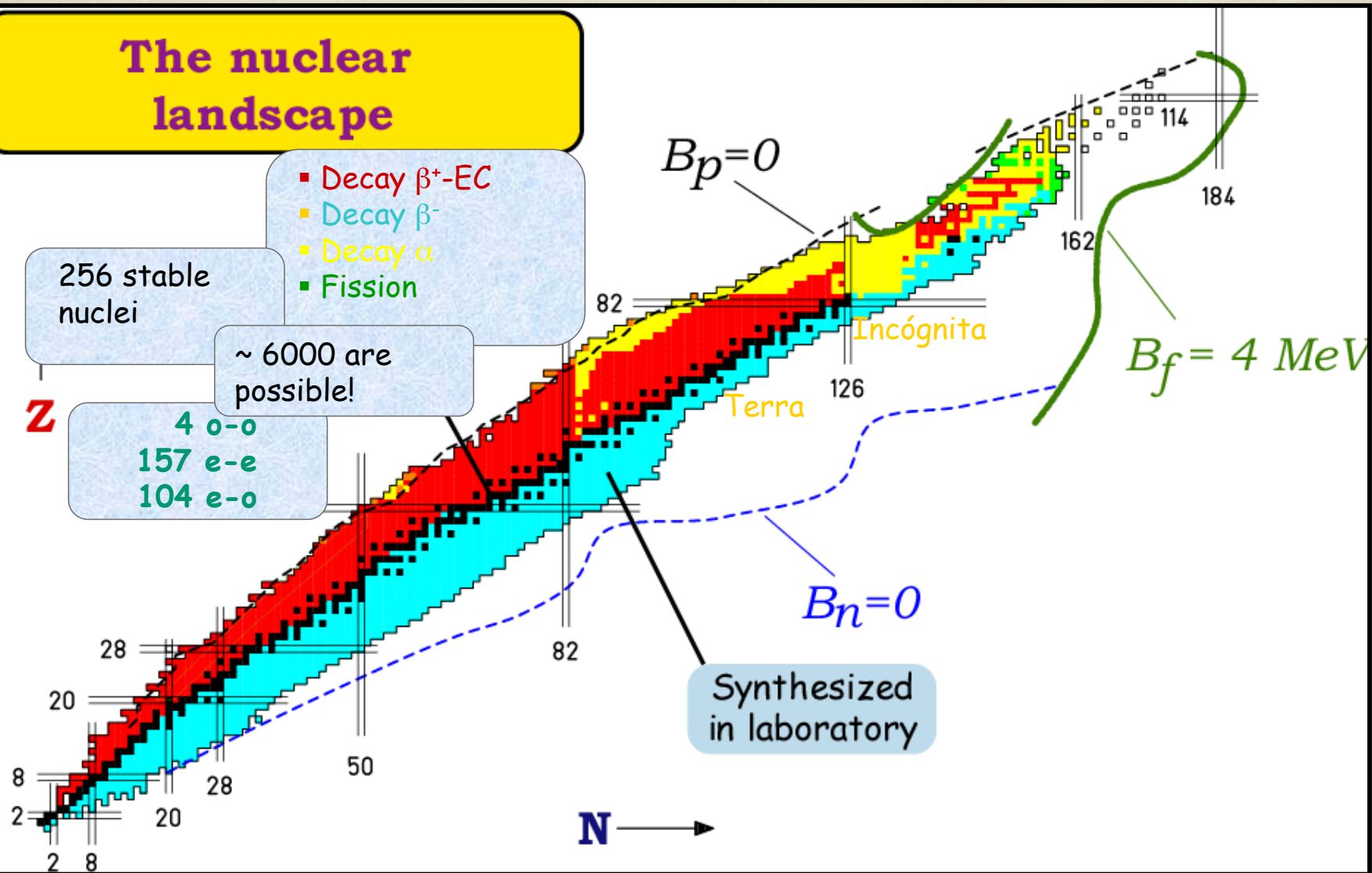
- Decay β^+ -EC
- Decay β^-
- Decay α
- Fission

256 stable nuclei

~ 6000 are possible!

Z

4 o-o
157 e-e
104 e-o



Nuclear stability and radioactivity

Atomic nuclei are very "particular" systems → only "magic" combinations of (Z,N) are possible → stable nuclei
 → nuclear interaction/ nuclear Structure

Far from "stable" configurations → excess of energy →

nucleons tends to reorganize → **particle emission**

→ **weak nuclear force** ($p \leftrightarrow n$) + **strong + Coulomb** → radioactive decay or **radioactivity**.

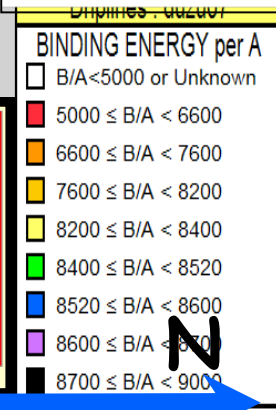
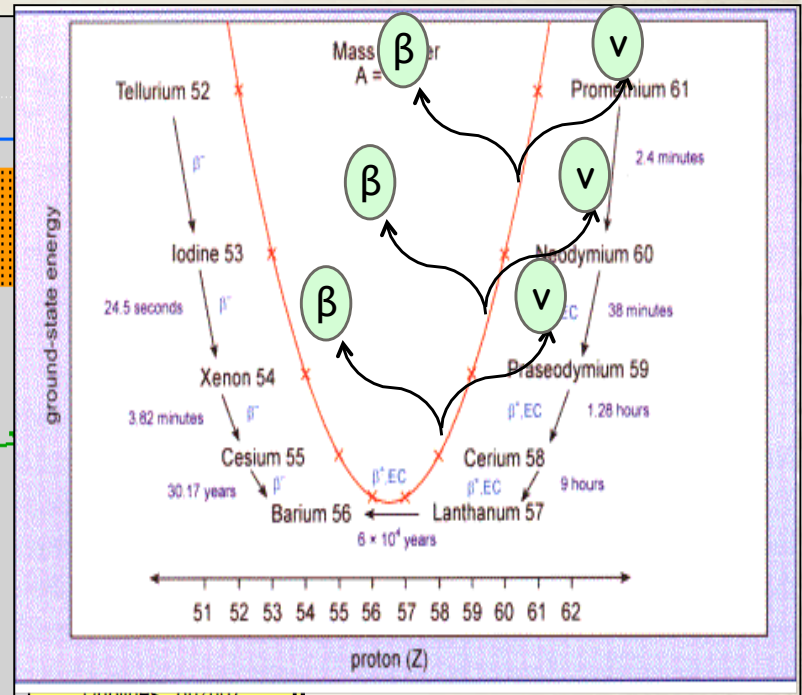
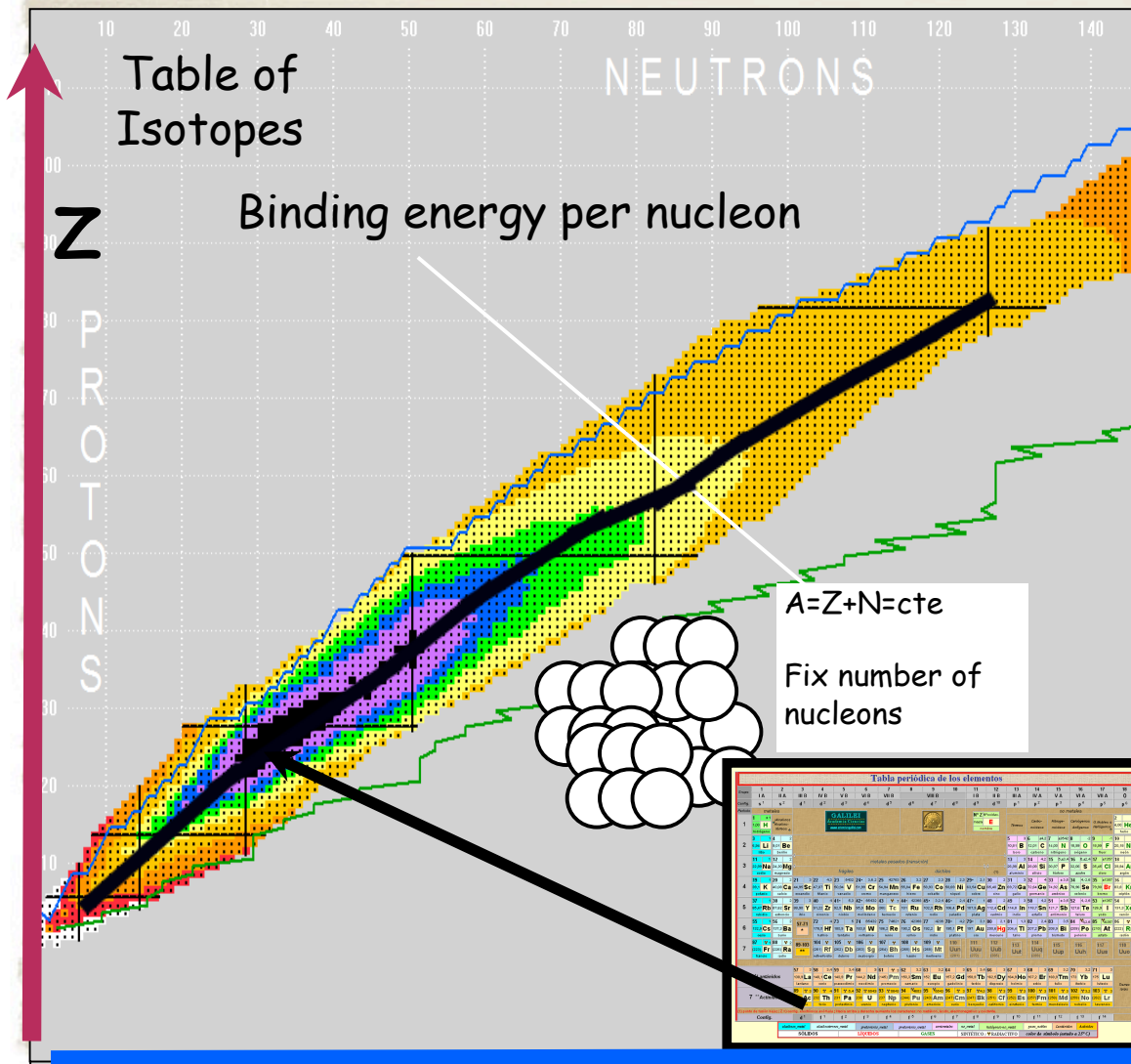
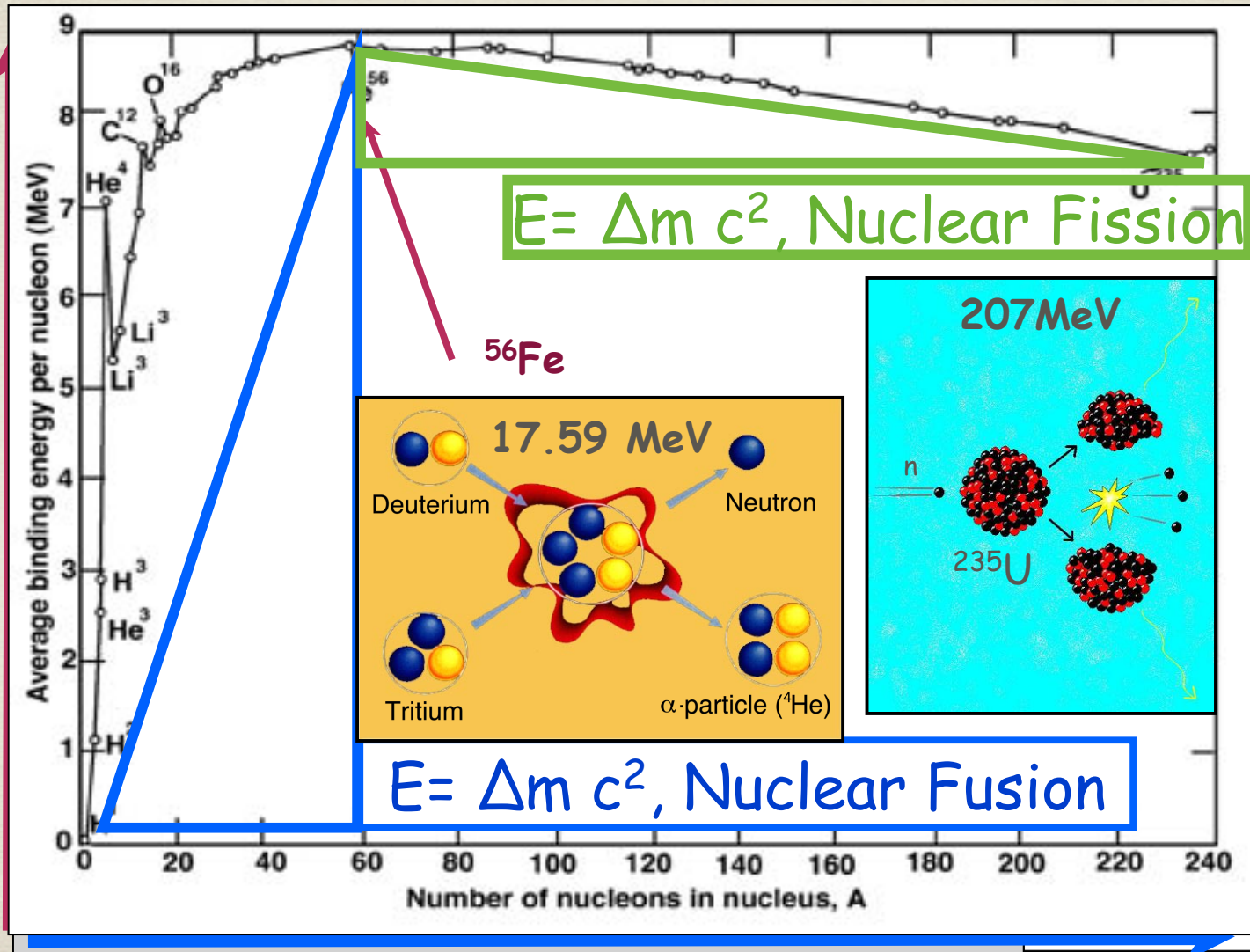
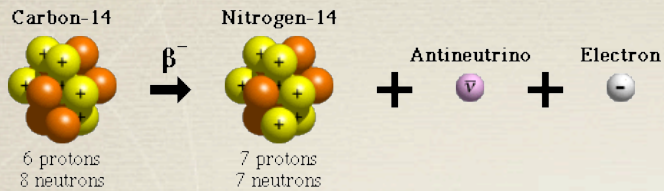


Tabla periódica de los elementos



Some common types of radioactivity

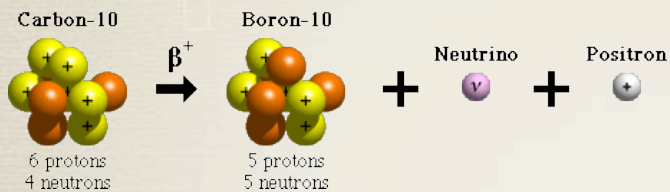
Beta -



A neutron is transformed in a proton, with the emission of one electron and one antineutrino;
 $Z \Rightarrow Z+1$
 $N \Rightarrow N-1$

A = constant

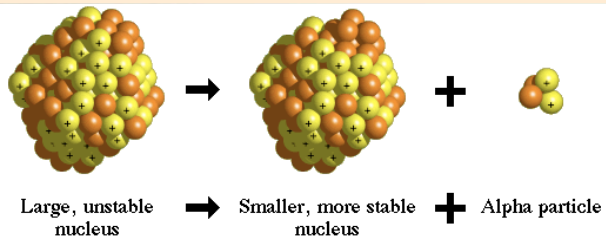
Beta +



A proton is transformed in a neutron, with the emission of a positron and one neutrino;
 $Z \Rightarrow Z - 1$
 $N \Rightarrow N+1$

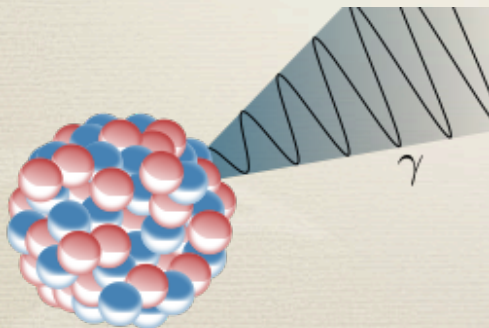
A = constant

Particle emission: Alpha



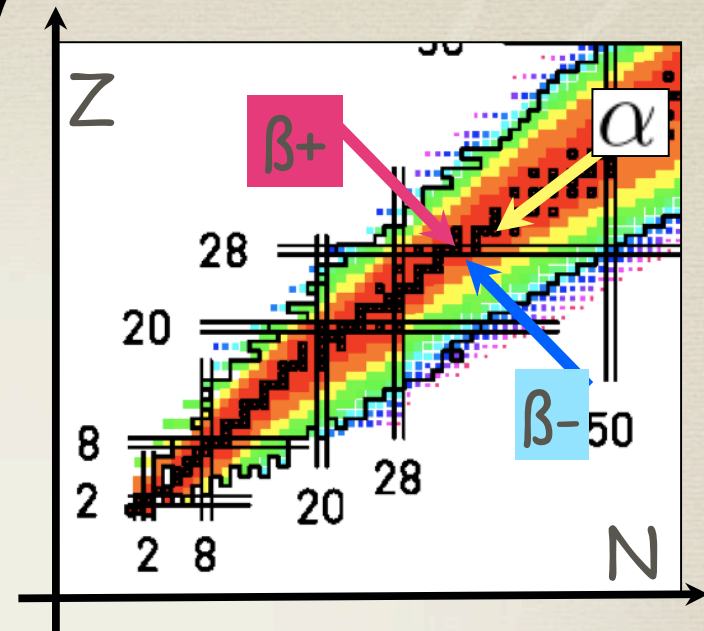
High mass nuclei can decay by emission of a helium nucleus;
 $Z \Rightarrow Z-2$
 $N \Rightarrow N - 2$
A \Rightarrow A-4

Gamma decay



Emission of electromagnetic radiation (photons) occurs during transitions between nuclear states of higher to lower energies.

\rightarrow **NO change in (N,Z) or A values**



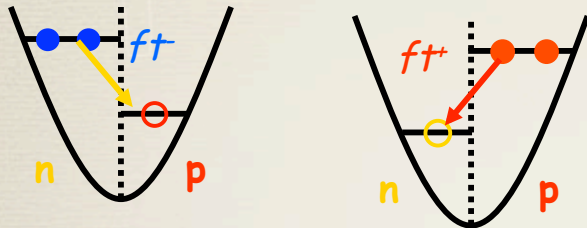
Beta decay

As it was previously discussed, **weak interaction** is one of the vehicles used for nuclear systems to release the excess of energy and travel from drip-lines to the Valley of Stability.

$$\beta^- : n \rightarrow p + e^- + \bar{\nu} \quad \beta^+ : p \rightarrow n + e^+ + \nu$$

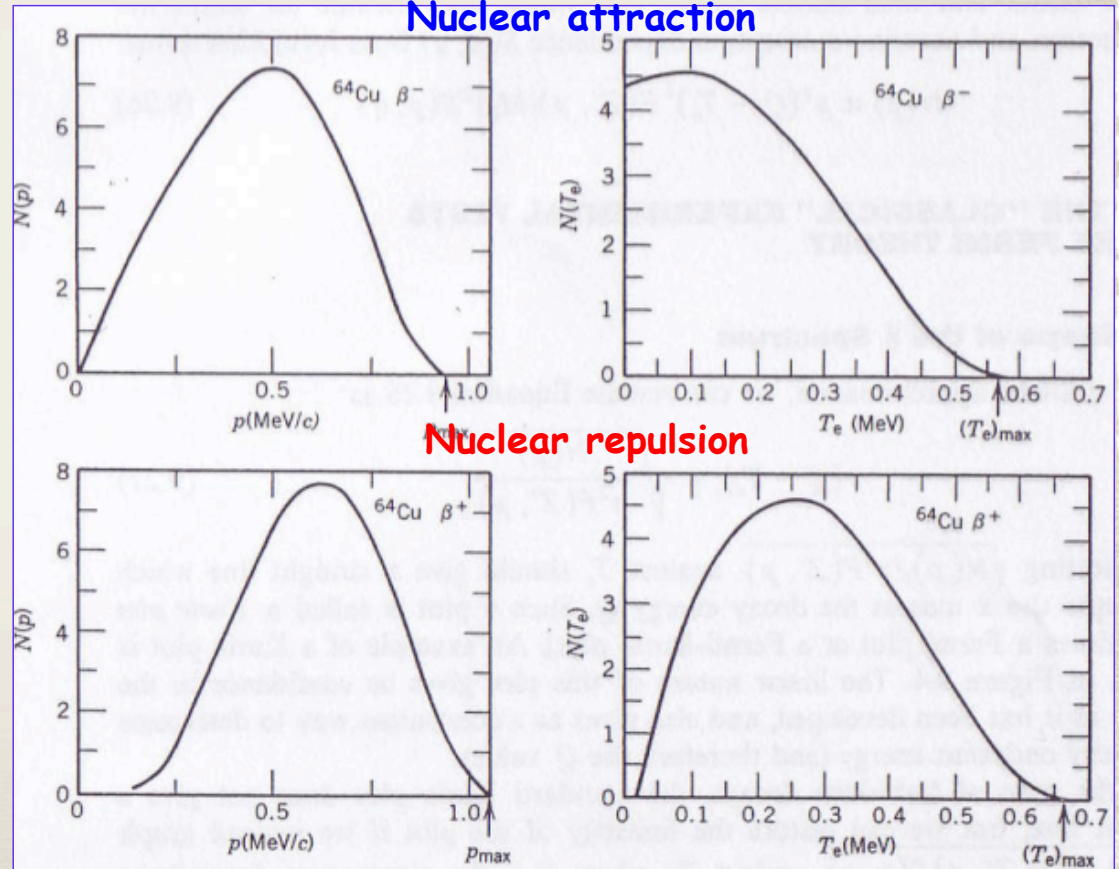
$$\text{E.C.} : p + e^- \rightarrow n + \nu$$

Competing process: **E.C.** Electron Capture, (Alvarez 1938)



Momentum & energy distributions

Nuclear attraction



The energy spectrum of beta particles is continuous: **three body process**

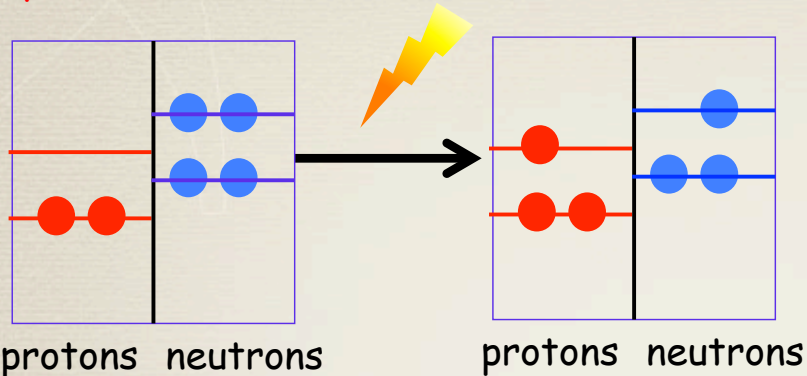
Pauli 1931

- Neutrino
- Beta
- Residual nucleus

Neutrino → Reines & Cowan 1950

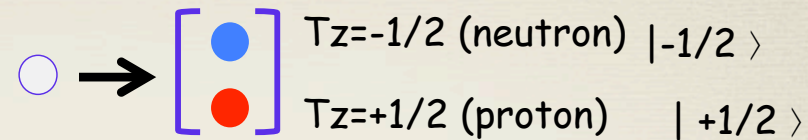
Beta decay and isospin

Beta decay: transformation of
 proton \leftrightarrow neutron

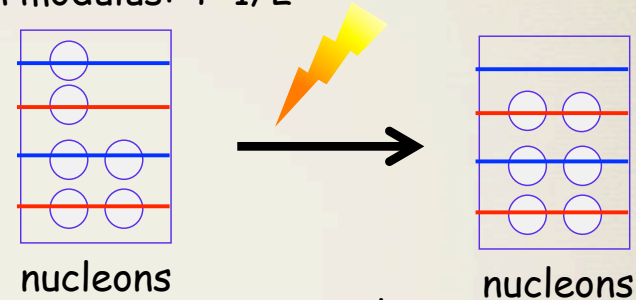


Concept of nucleon: particle that can be proton/
 neutron:
 \rightarrow new quantum number ISOSPIN (T) describes
 "the charge state of the nucleon"

Dirac ket



Isospin modulus: $T=1/2$



eigenvalues

$$T_z | +1/2 \rangle = +1/2 | T = +1/2 \rangle$$

$$T_z | -1/2 \rangle = -1/2 | T = +1/2 \rangle$$

The isospin operator

OPERATIONS & OPERATORS

$| -1/2 \rangle$ neutron $| +1/2 \rangle$ proton

$$I = | -1/2 \rangle \langle -1/2 | + | +1/2 \rangle \langle +1/2 |, \text{ identity}$$

$$\langle -1/2 | +1/2 \rangle = \langle -1/2 | +1/2 \rangle = 0, \text{ orthogonal}$$

$$\langle -1/2 | -1/2 \rangle = \langle +1/2 | +1/2 \rangle = 1, \text{ normalization}$$

$$T_z = (-1/2) | -1/2 \rangle \langle -1/2 | + (+1/2) | +1/2 \rangle \langle +1/2 |$$

$$T_+ = | +1/2 \rangle \langle -1/2 |$$

Isospin flip

$$T_+ | -1/2 \rangle = | +1/2 \rangle$$

$$T_- = | -1/2 \rangle \langle +1/2 |$$

operators

$$T_- | +1/2 \rangle = | -1/2 \rangle$$

$$Q = (2 T_z + 1)/2 = \text{charge operator}$$

$T_{\pm} \rightarrow$ beta decay operators!

$$M(A, T, T_z) = a(A, T) + b(A, T) T_z + c(A, T) T_z^2$$

The isobaric multiplet mass equation (IMME), Wigner 1957
 \rightarrow drip lines, exotic radioactivity, etc

For a system nucleons $T = \sum T(i)$ $T_z = \sum T_z(i)$ $T^{+/-} = \sum T^{+/-}(i)$ $i=1...A$ nucleons

Beta interaction

$$V_b^{(+/-)} = g_v T^{(+/-)} + g_A S T^{(+/-)}$$

g_v = Fermi constant
 g_A = Gamow-Teller constant

Experimentally, beta decay can change spin of final nuclei (S operator)

Beta transition prob. \rightarrow Fermi Golden Rule

$$\lambda(i, f; E_n, E_b) = 2\pi/\hbar |\langle i | V_b | f \rangle|^2 \rho(Q - E_n, E_b)$$

Transition probability

matrix element

Density of final states β, ν
 recoil excit: E_n

$$|M(F)|^2 = |\langle i | T^{(+/-)} | f \rangle|^2$$

$$|M(GT)|^2 = |\langle i | S T^{(+/-)} | f \rangle|^2$$

$$\rho(Q - E_n, E_b) \sim p_b(Q - E_b)^2 F(Z_f, E_b)$$

$|i\rangle = |A; i\rangle$ = initial nuclear state

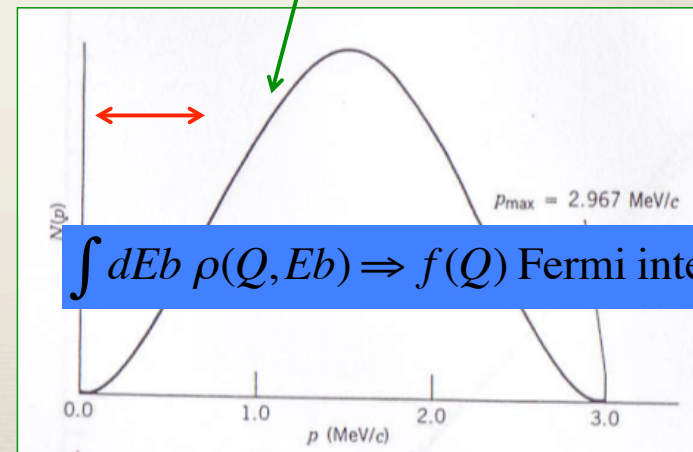
$|A; f\rangle = |A; f\rangle \times |\beta\rangle \times |\nu\rangle$ final state

Phase space
 nucl rep./atrac.
 (Fermi function)

1st Forbidden decays, etc

$$|\beta\rangle \langle \nu| \sim e^{ik_\nu r} e^{ik_\beta r} \sim 1 - (k_\beta + k_\nu)r + \dots \sim 1$$

Allowed approximation
 Allowed decays
 $L=0$ ($r=0$)



$\int dE_b \rho(Q, E_b) \Rightarrow f(Q)$ Fermi integral

Finally

$$\lambda(i, f; En) = \frac{m_e^5 c^4}{2\pi \hbar^7} \left[g_V^2 |M(F)|^2 + g_A^2 |M(GT)|^2 \right] f(Z_f, Q - En)$$

$$f(Z, E) = \frac{1}{m_e^5 c^7} \int_0^{p^{\max}} F(Z_f, p) p^2 (E - Eb) dp \quad \text{Fermi integral}$$

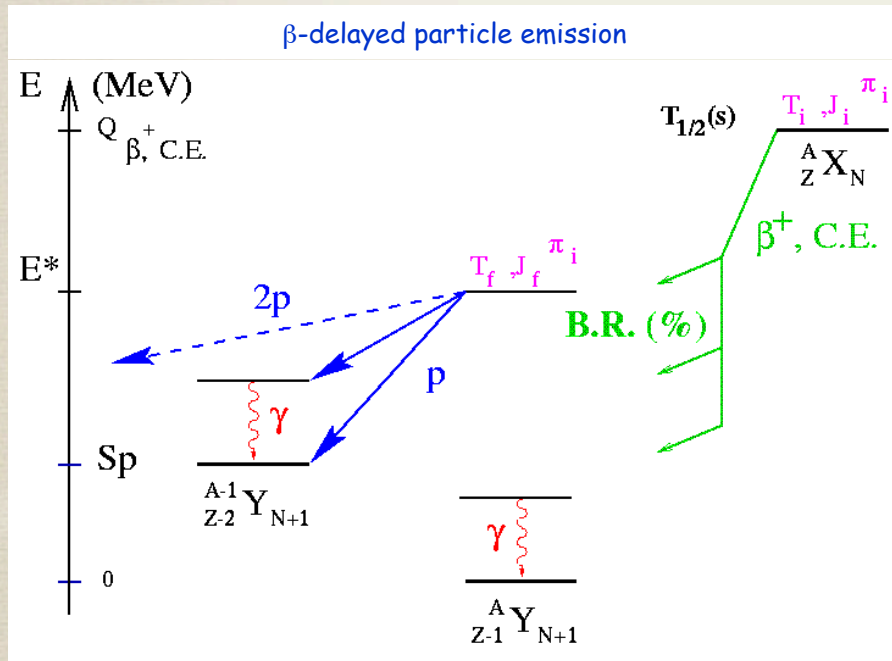
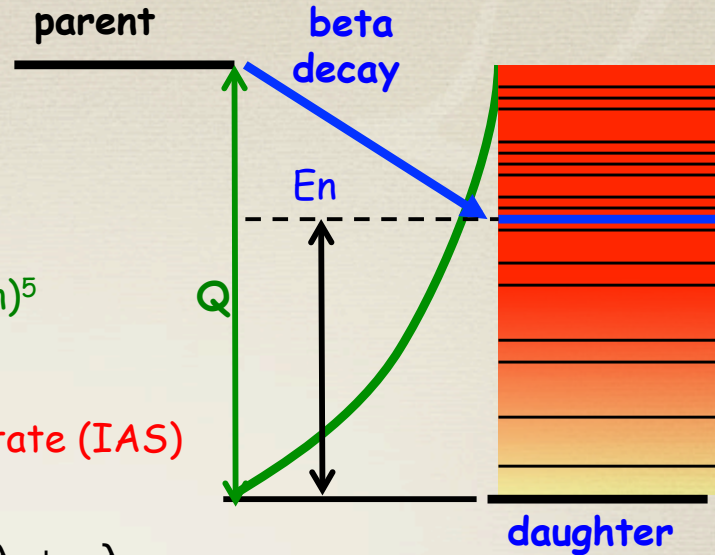
$$F(Z_f, Q - En) \sim (Q - En)^5$$

• Selection rules for allowed approx:

- Fermi: $\Delta T=0$; $\Delta J=0$; $\pi_f = \pi_i$ → Isobaric analog state (IAS)
- Gamow-Teller: $\Delta T=0\pm 1$; $\Delta J=0\pm 1$; $\pi_f = \pi_i$

Branching ratios and partial half-life

$$\begin{cases} \lambda_T = \lambda_1 + \lambda_2 + \dots + \lambda_N \\ T_T = \ln(2) / \lambda \rightarrow 1/T = 1/T_1 + 1/T_2 + \dots + 1/T_N \\ Br(i) = \lambda_i / \lambda_T = T_T / T_i \end{cases}$$



ft-value (comparative half-life)

$$\lambda = \ln(2) / T_{1/2}$$

$$ft = f * \frac{T_{1/2}}{Br} = \frac{K}{g_V^2 |M(F)|^2 + g_A^2 |M(GT)|^2}$$

$$ft = \frac{C}{B(F) + B(GT)}$$

B(F), B(GT): reduced transition probability

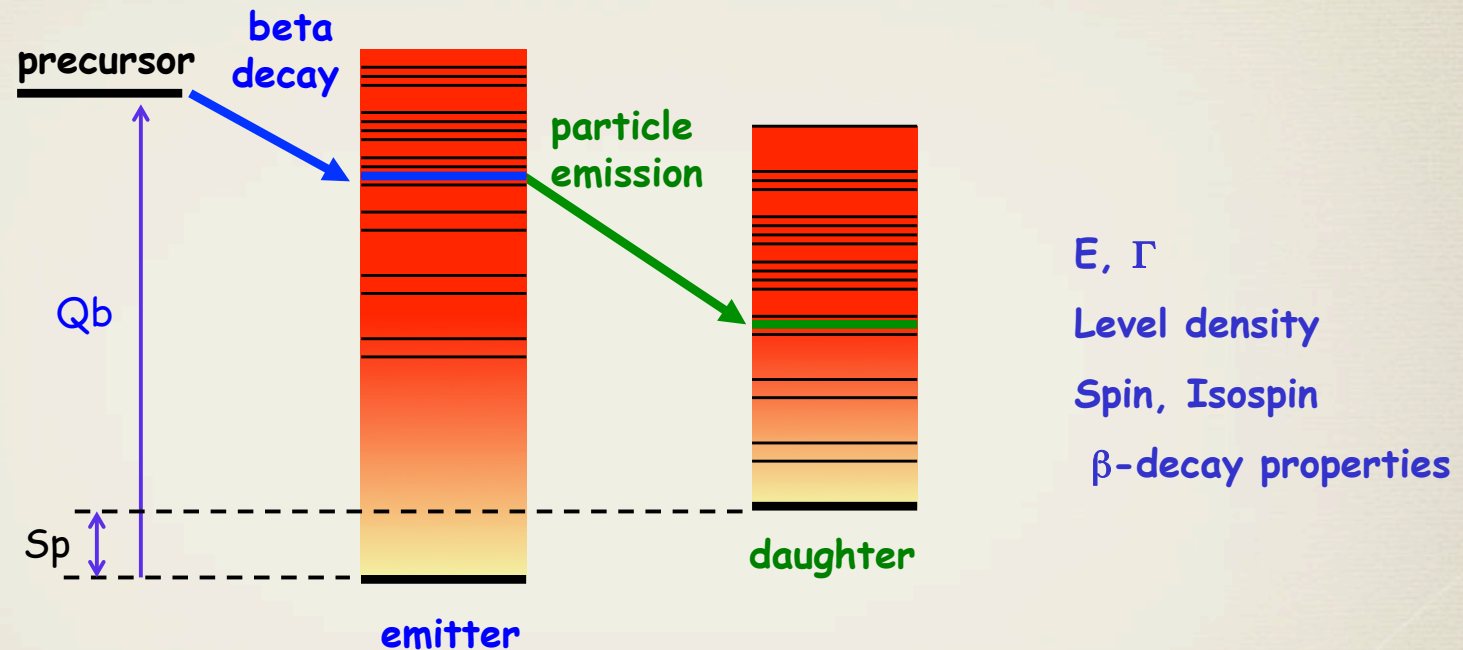
Large range ft ~ 10³ → 10²⁰ → Tabulate Log(ft)

Beta delayed particle emission

Emission of particles from nuclear (excited) states populated by the beta decay

Two processes:

- **Beta decay** from the parent nucleus (precursor)
- **Particle emission** from excited states of "emitter" nucleus



→ beta decay to excited levels of "emitter" nucleus; if the excited state is over separation energy S_p → emission of particles

→ The half-life of beta decay is much longer than the nuclear level of emitter, the half-life of the process is given by the beta decay → "**beta - delayed ...**"

Particle emission: transitions and decaying states

The wave functions obtained by solving the Schrödinger equation for time independent potentials have the property of being stationary states

$$\hat{H}\Psi_o(r,t) = E(0)\Psi_o(r,t) \quad \Psi_o(r,t) = \Psi_o(r)e^{-i\frac{E(0)}{\hbar}t}$$

States will remain in that energy eigenstate forever!

Under a **sudden change** of the potential (like **beta decay** $p \leftrightarrow n$), we get a new hamiltonian H_{new} and the "old" wavefunctions are no more eigenvalues \rightarrow **start evolution** with time:

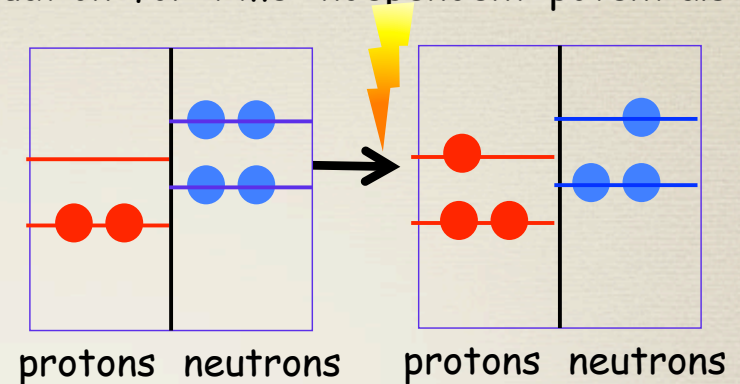
$$\hat{U}(t)\Psi_o(r,t) = e^{-i\frac{\hat{H}_{new}t}{\hbar}}\Psi_o(r,t) = \sum c_i(t)\phi_i(r)e^{-i\frac{E(i)_{new}t}{\hbar}}$$

$$\hat{H}_{new}\phi_i(r,t) = E(i)_{new}\phi_i(r,t)$$

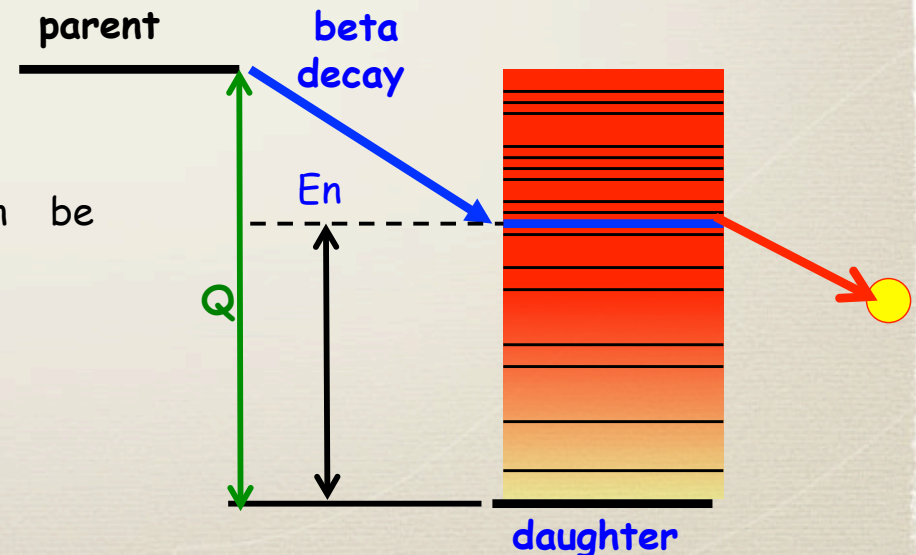
The transition process (**particle emission**) can be described by the Fermi Golden Rule

$$\lambda = \frac{2\pi}{\hbar} |V_{if}|^2 \rho(E_f)$$

$$V_{if} = \int \Psi_f^*(H_{new} - H_{old})\Psi_i \quad \rho(E_f) = \frac{dn_f}{dE_f}$$



Some of these new states are **continuum states** \rightarrow **particle emission**



If $E_f > S_p \rightarrow$ tunnel through Coulomb barrier
 $\rightarrow P(r,t)$ decreases with time.

\rightarrow use of a complex energy eigenvalue in the final system:

$$E_d + i\Gamma_d/2$$

$$\phi_d(r,t) = N\phi_d(r)e^{-\frac{i}{\hbar}(E_d+i\Gamma_d/2)t} = N\phi_d(r)e^{-\frac{i}{\hbar}E_d t} e^{-\frac{1}{\hbar}\Gamma_d t}$$

$$P(r,t) = N^2 |\phi_d(r,t)|^2 = N^2 |\phi_d(r)|^2 e^{-\frac{1}{\hbar}\Gamma_d t}$$

$$P(r,t) = N^2 |\phi_d(r)|^2 e^{-\lambda_d t}$$

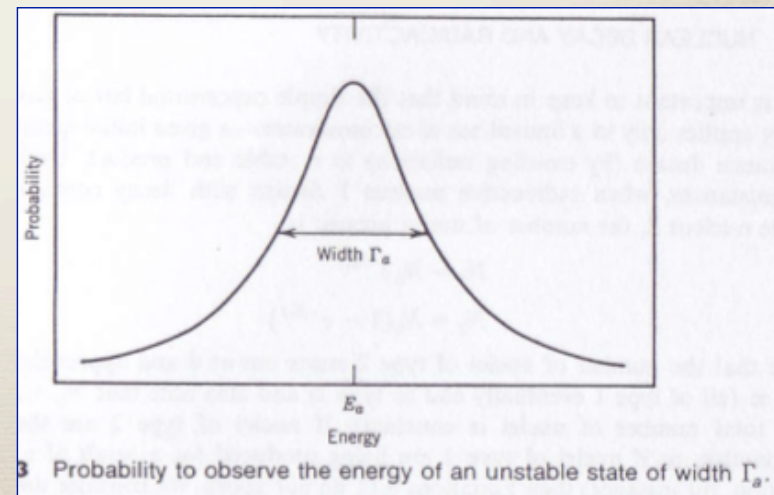
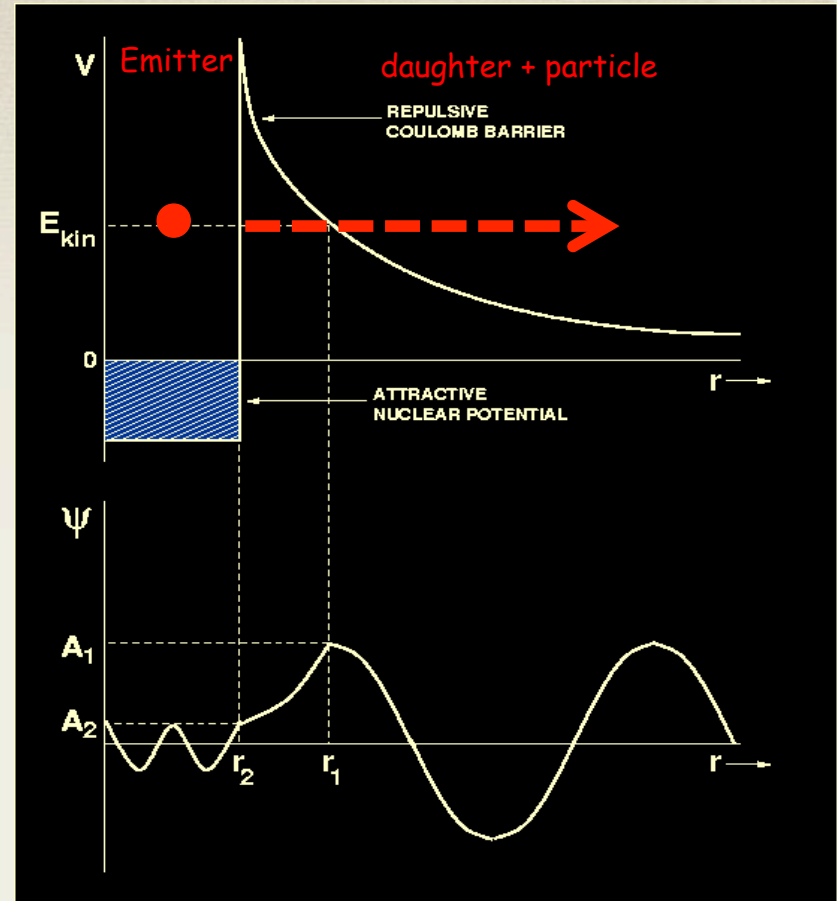
$$\lambda = \frac{1}{\tau} = \frac{\Gamma}{\hbar}$$

For the **energy distribution** (energy representation)
 \rightarrow Fourier transform

$$\phi_d(E) \approx \int e^{-\frac{i}{\hbar}Et} \phi_d(t) dt \approx \int e^{-\frac{i}{\hbar}Et} e^{-\frac{i}{\hbar}E_d t} e^{-\frac{1}{\hbar}\Gamma_d t} dt$$

$$\phi_d(E) \approx \frac{1}{(E - E_d) + i\frac{\Gamma}{2}}$$

$$P(E) \approx \frac{1}{(E - E_{dec})^2 + \left(\frac{\Gamma}{2}\right)^2}$$



Why complex eigenvalues? $\rightarrow E_d + i \Gamma_d/2 \rightarrow$ naturally arise from solving Schrödinger equation at $E > 0!$

Georg Gamow: simple model of alpha decay, G.A. Gamow, Zs f. Phys. 51 (1928) 204; 52 (1928) 510

\rightarrow Quantum tunneling through barrier

$$u''(r) = \left[\frac{l(l+1)}{r^2} + \frac{2\mu}{\hbar^2} V(r) - k^2 \right] u(r)$$

$$u(r) \sim C_0 r^{l+1}, \quad r \rightarrow 0$$

$$u(r) \sim C_+ H_{l,\eta}^+(kr), \quad r \rightarrow +\infty \text{ (bound, resonant)}$$

$$u(r) \sim C_+ H_{l,\eta}^+(kr) + C_- H_{l,\eta}^-(kr), \quad r \rightarrow +\infty \text{ (scattering)}$$

If keep same boundary condition $\rightarrow H^+(kr), r \rightarrow \infty$
 Bound and resonant states \rightarrow poles of the Scattering matrix $S(k)$ (matching with outgoing WF)

Bound states:

\rightarrow pure imaginary K values: $\sim -i K_i, E_r < 0$

Resonant states:

\rightarrow complex K values: $K_r - i K_i, E_r > 0, \Gamma > 0$

\rightarrow **GAMOW STATES**

$$\hat{1} = \sum_{i=b} |u_i\rangle \langle \tilde{u}_i| + \sum_{j=r} |u_j\rangle \langle \tilde{u}_j| + \int_{L^+} |\varphi(k)\rangle dk \langle \tilde{\varphi}(k^*)|$$

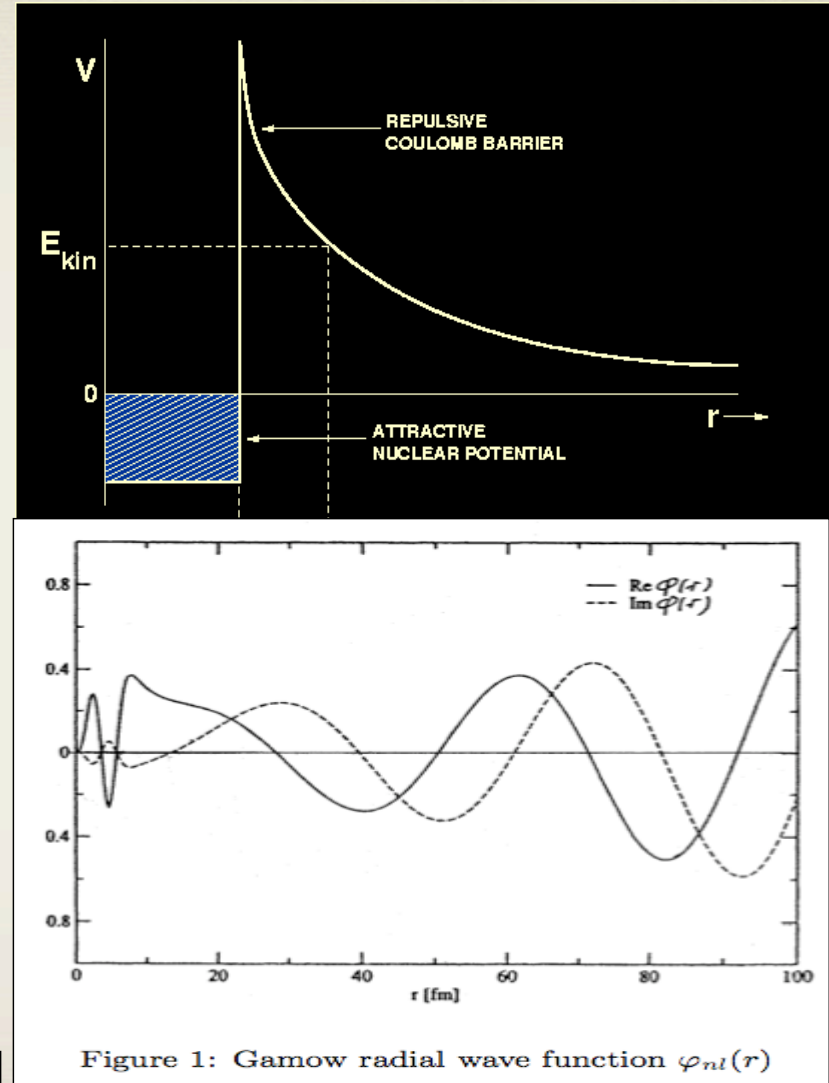


Figure 1: Gamow radial wave function $\varphi_{nl}(r)$

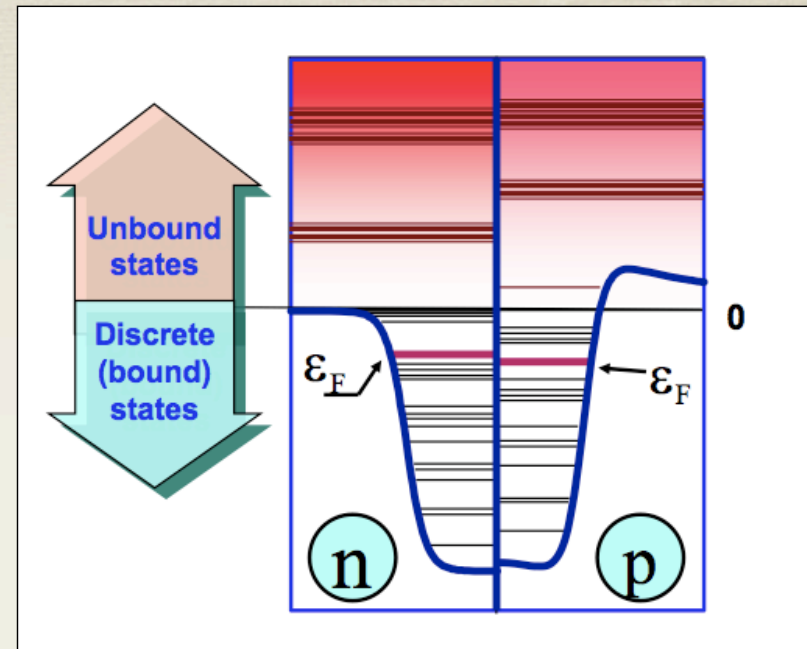
Consistent description of bound and scattering states: → a rigged Hilbert space (Gel'fand triple space): 1960s Gel'fand combined Hilbert space with the theory of distributions.

Spectacular applications: Shell model in the continuum// → Shell model in the complex energy plane; N. Michel, W. Nazarewicz, M. Oloszajzak and T. Vertse (J. Phys. G.: Nucl. Part. Phys. 36 (2009) 013101)

Difficult to overestimate the importance of Gamow theory!!.

Some references: Humblet and Rosenfeld, Nucl. Phys. 26, 529 (1961); T. Berggren, Nucl. Phys. A 109 (1968) 265. R. de la Madrid, Nucl. Phys. A812, 13 (2008)

Gamow states of a finite potential



R-MATRIX DESCRIPTION

Tradicional method → based on R-matrix theory for unbound nuclei → scattering, reactions, particle decay. (F. C. Barker, Aust. J. Phys., 1988, 41, 743-63, E.K. Warburton, PRC 33 (1986)303-313)

$$P(E) \propto \left| \sum_i \frac{G(i)^{1/2} \Gamma(i)^{1/2}}{(E(i) + \Delta(i) - E - i \frac{\Gamma}{2})} \right|^2$$

$G(i)$: feeding factor of the decaying state

$\Gamma(i)$: level width $\Gamma(i) : 2 P(E) * \gamma^2$

$\Delta(i)$: shift factor

(WF matching at pot. radius)

$E(i)$: level energy/resonance

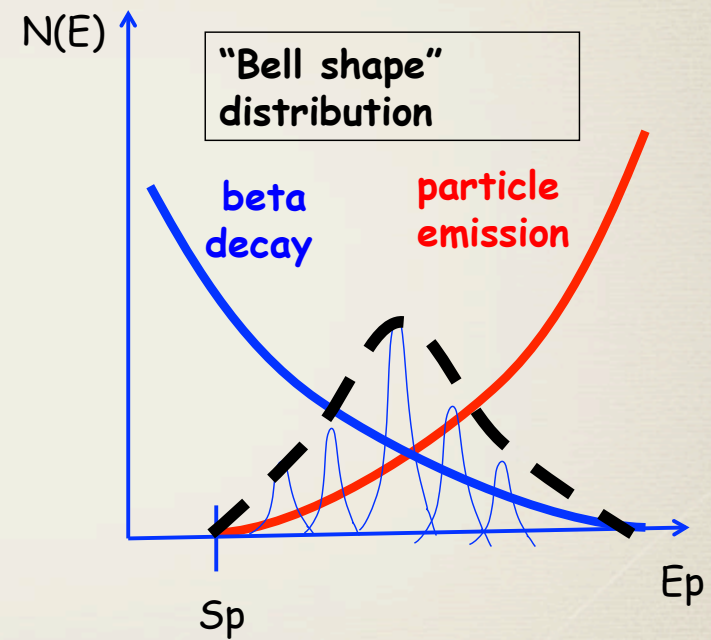
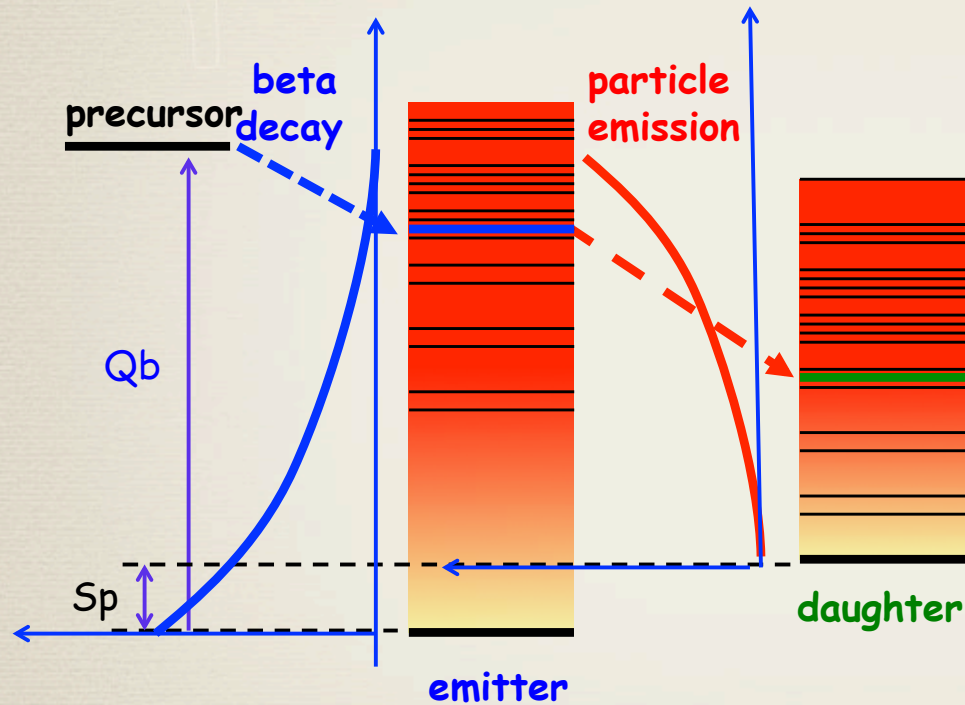
Penetration factor (barrier)

Reduced width (nuclear matrix element)

SUMMARY: what to expect for beta delayed particle emission

Two processes:

- Beta decay \rightarrow FERMI INTEGRAL (Matrix elements) $\rightarrow (Q-E_n)^5$
- Particle emission \rightarrow BARRIER PENETRABILITY $\sim P(E_k) \sim 1/(1 + \exp((E_B - E_k)/w_b))$ (parabolic)
- Breit - Wigner shapes on each level
- Density of states above S_p



history → observed since early stages of nuclear physics:

Beta delayed alphas ($\beta\alpha$): Rutherford (1916) [*Philos. Mag.* **31** (1916) 379]

→ "Long range alpha particles followed by beta decay of ^{212}Bi "

Beta delayed protons (βp): Marsden (1914) → $^{14}\text{N}(\alpha, p)^{17}\text{O}$ [*Philos. Mag.* **37** (1919) 537]; Álvarez (1950) bombarded ^{10}B and ^{20}Ne with 32 MeV protons → beta delayed ^8B , ^{20}Na α -emitters

The modern era begins in 1960's (βp , $\beta 2p$)/ Zeldovich, Karnaukhov, Goldansky...

[*Goldanskii NPA* **19** (1960) 482]

→ **Spectroscopic tool**: Information about level energies, spins and parities of participant nuclei (precursor, emitter, daughter)

At present days investigations on beta delayed radioactivity are very intense, particularly with the use of **radioactive beams**:

typical decay mechanism at drip lines

- **Large Q_b values** → access high energy states

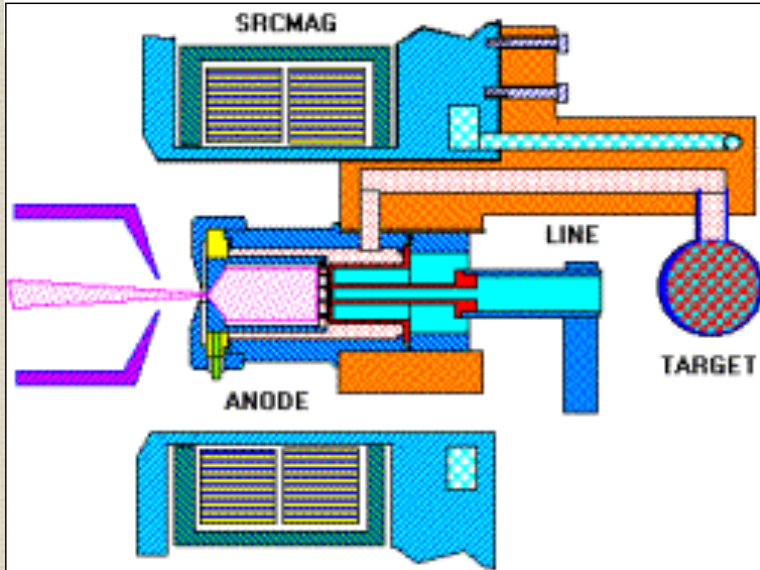
Good alternative to gamma spectroscopy and nuclear reactions limited by beam intensities $\sim 10^4$ pps

- Beta delayed particle emission → limited by selection rules of beta decay

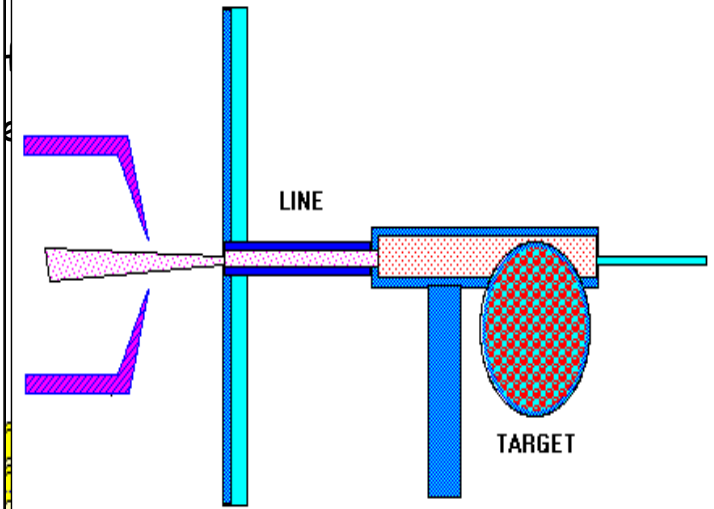
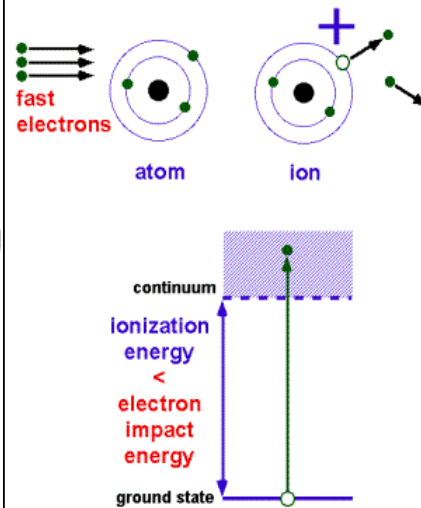
- Usually first type of studies close to drip lines → low isotope production → largest yields obtained directly after ion source and implanted on decay foil.

- Relatively "simple" experimental setups.

Plasma ion source

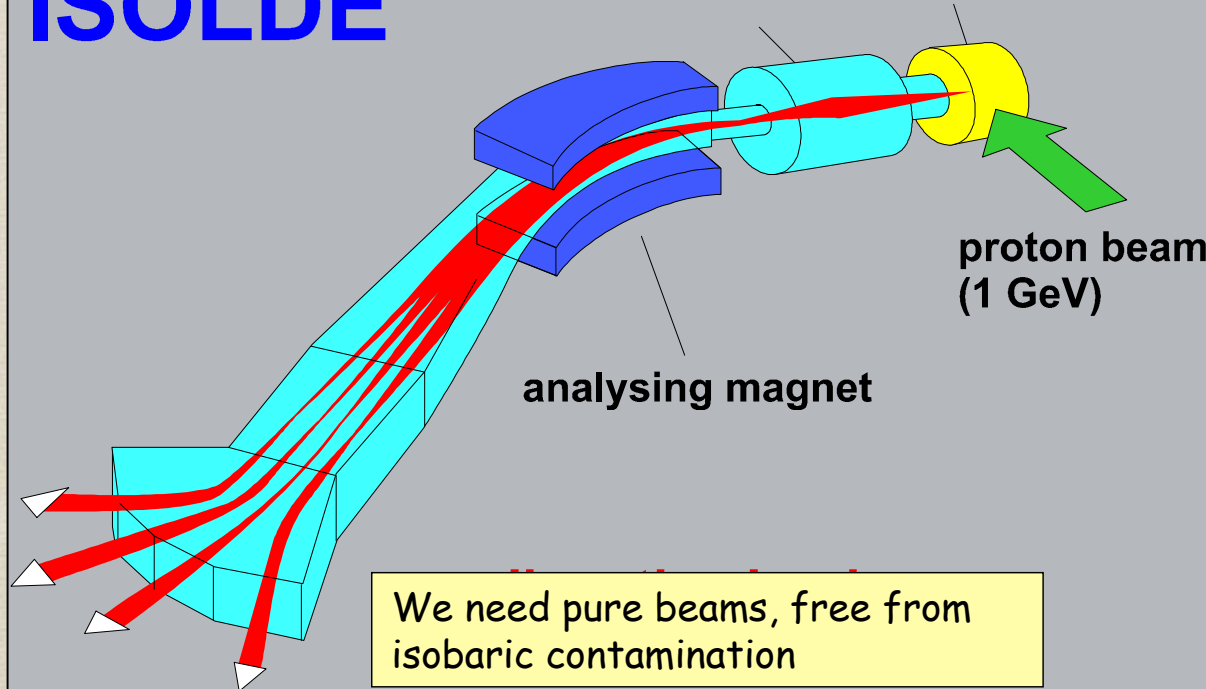


Ionization by electron impact

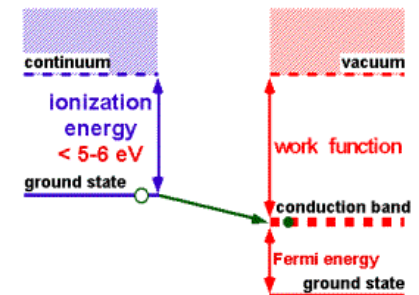
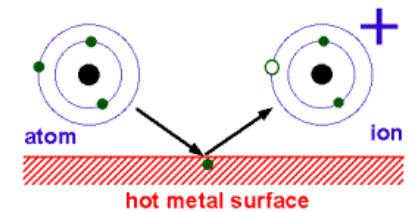


ISOLDE

target - ion source

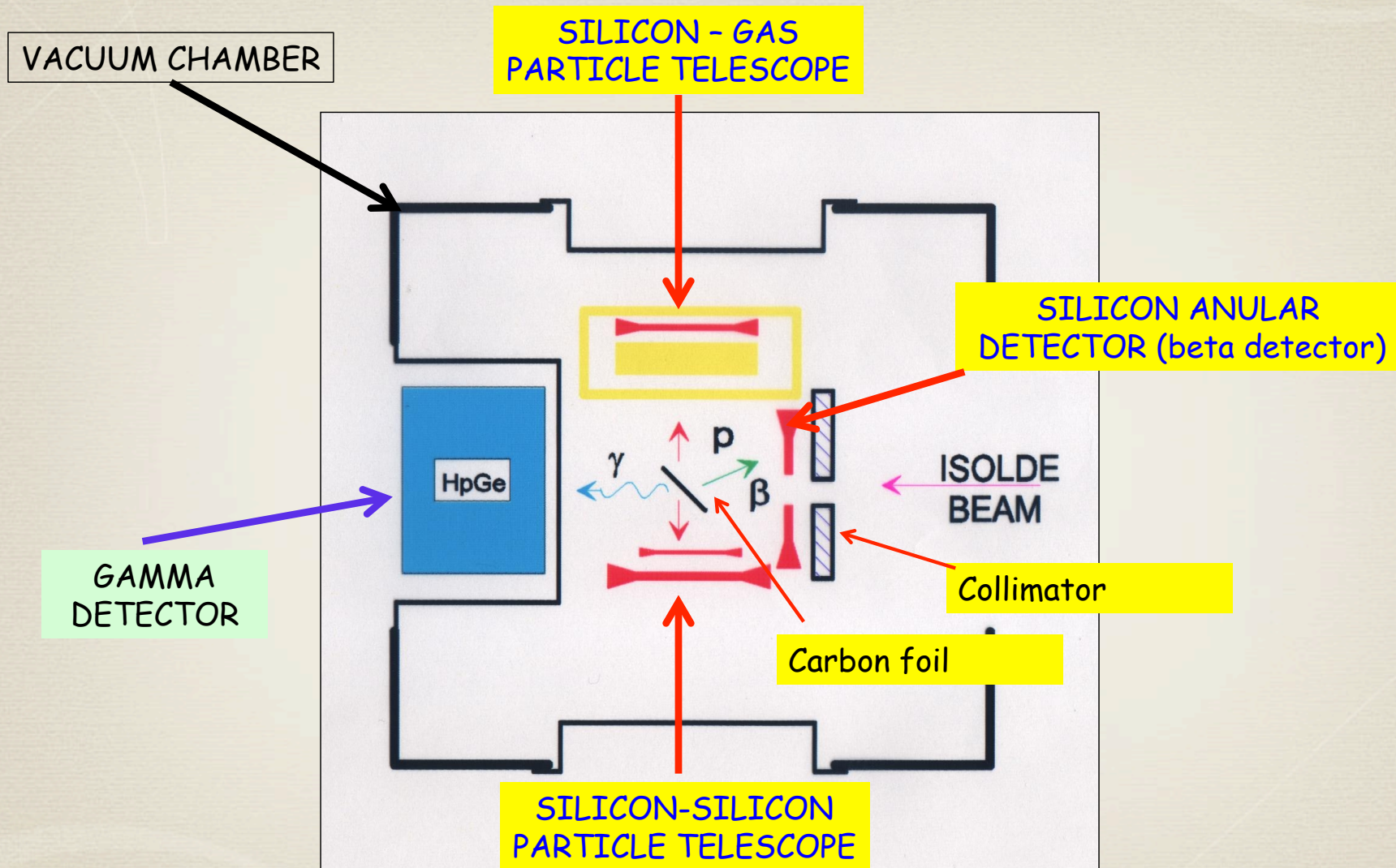


Surface Ionization



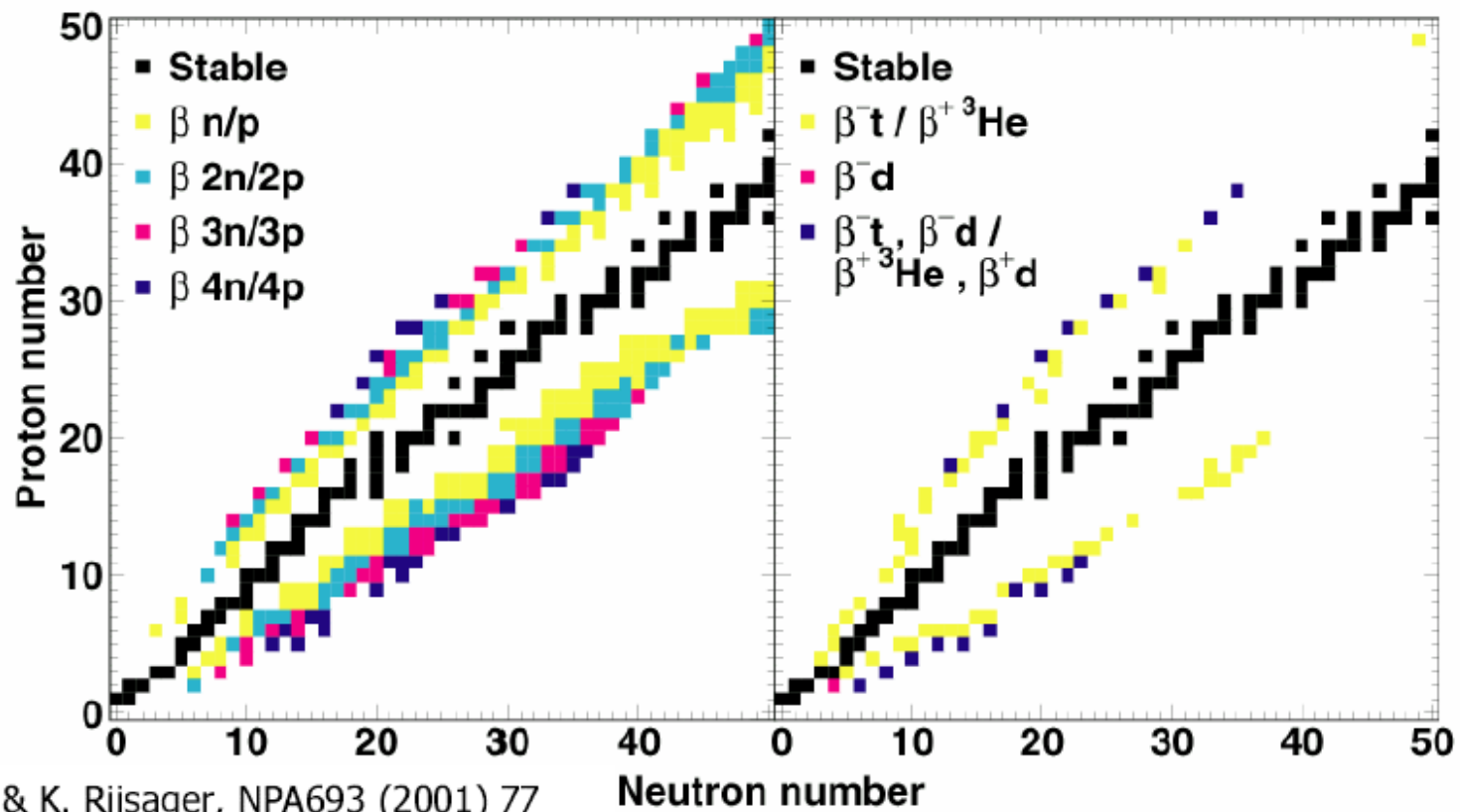
Example: ^{31}Ar produced at ISOLDE with a CaO -target and plasma ion-source (cooled transfer line) at a rate of 2 atom/s!!

TYPICAL EXPERIMENTAL SETUP



Low energy beam (~ 60 keV) \rightarrow point like sources \rightarrow good angular resolution \rightarrow angular correlations

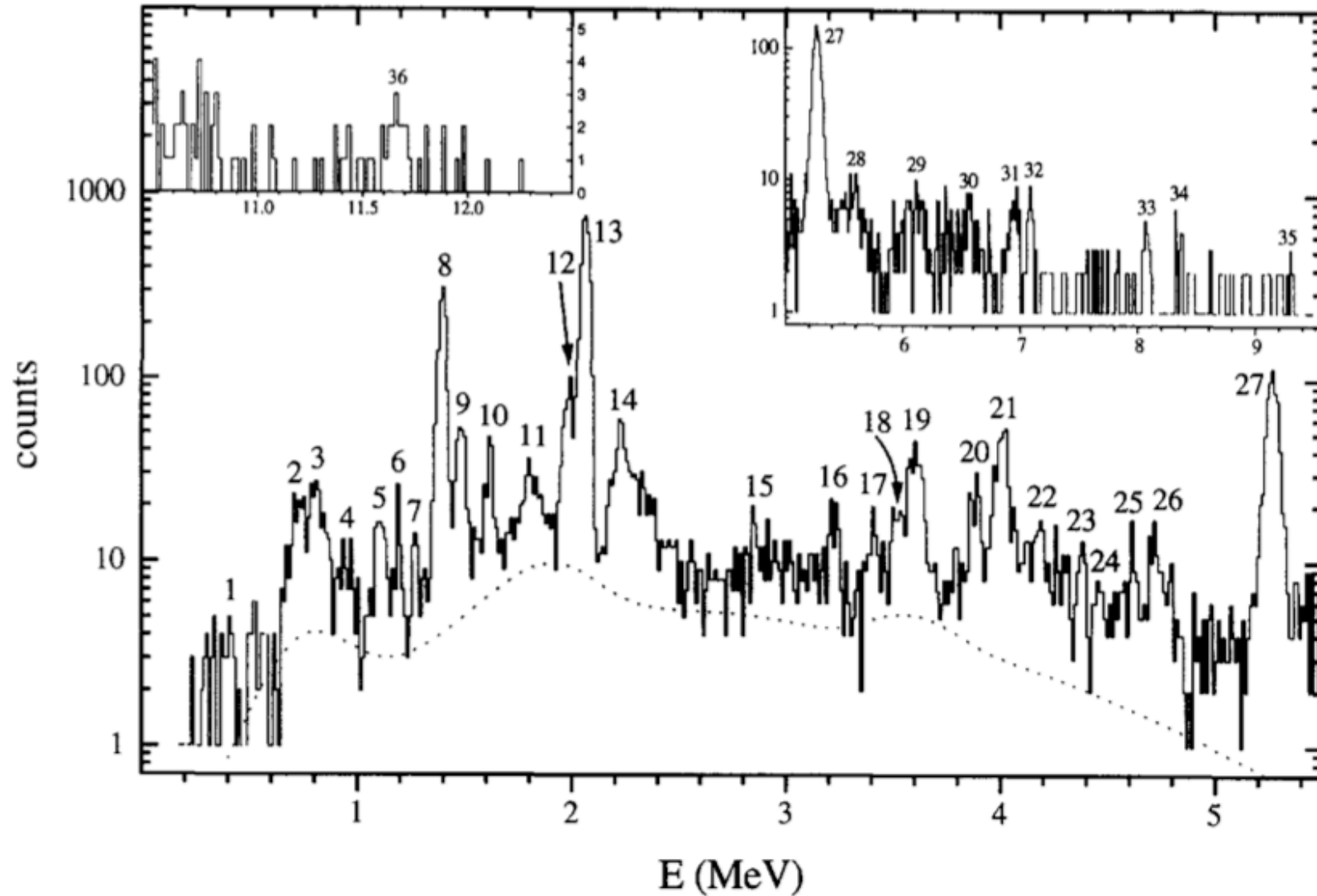
Beta delayed particle emitters



EXPERIMENTAL PROTON SPECTRUM

480

L. Axelsson et al./Nuclear Physics A 634 (1998) 475-496



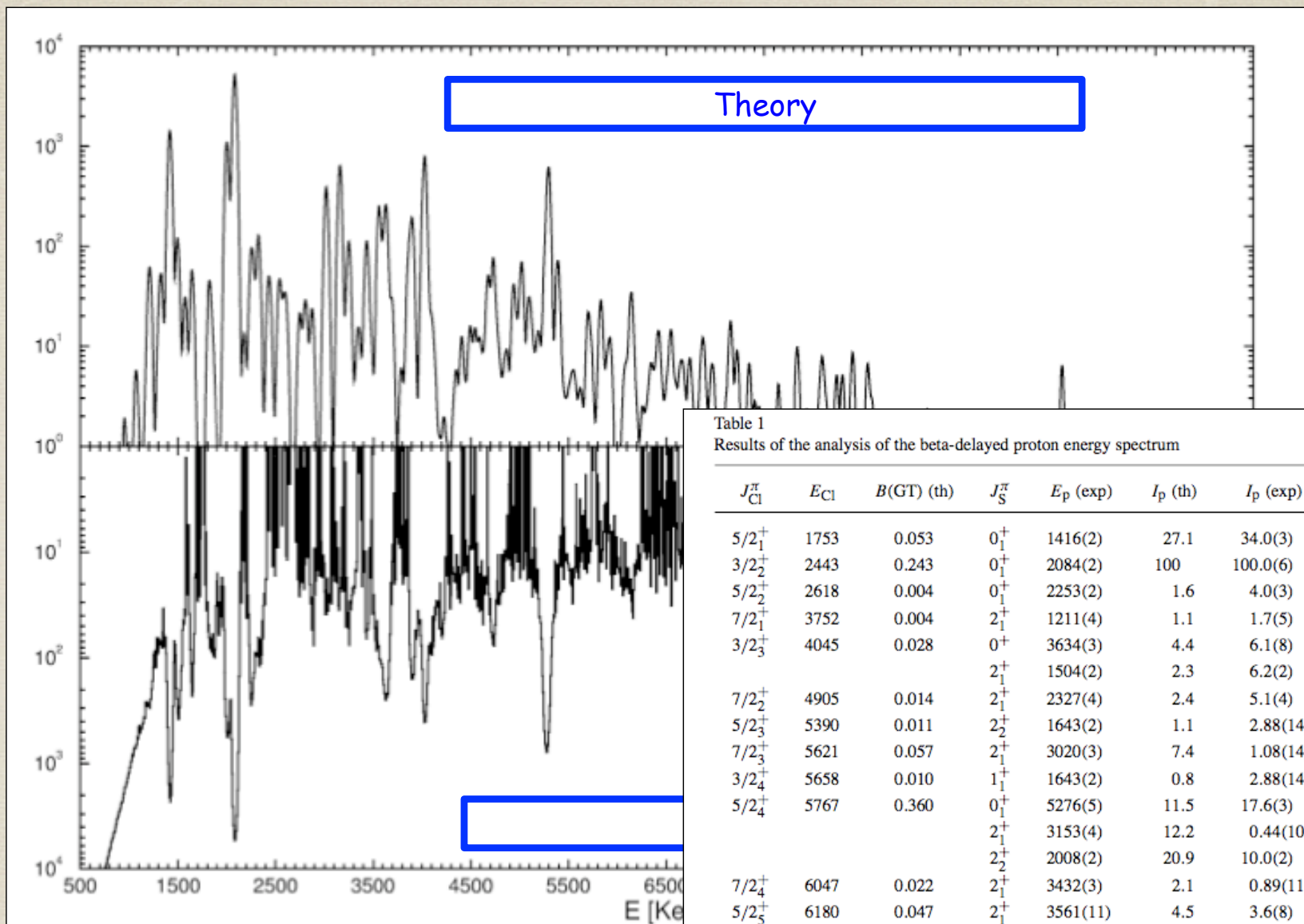


Fig. 3. Experimental energy spectrum of the beta-delayed proton energy spectrum (up) after adjustment of level energies. The width of 12 keV.

Table 1
Results of the analysis of the beta-delayed proton energy spectrum

J_{Cl}^{π}	E_{Cl}	$B(GT)$ (th)	J_S^{π}	E_p (exp)	I_p (th)	I_p (exp)	Γ (th)
$5/2_1^+$	1753	0.053	0_1^+	1416(2)	27.1	34.0(3)	0.02
$3/2_2^+$	2443	0.243	0_1^+	2084(2)	100	100.0(6)	0.4
$5/2_2^+$	2618	0.004	0_1^+	2253(2)	1.6	4.0(3)	0.8
$7/2_1^+$	3752	0.004	2_1^+	1211(4)	1.1	1.7(5)	0.11
$3/2_3^+$	4045	0.028	0^+	3634(3)	4.4	6.1(8)	9.8
			2_1^+	1504(2)	2.3	6.2(2)	5.1
$7/2_2^+$	4905	0.014	2_1^+	2327(4)	2.4	5.1(4)	1.7
$5/2_3^+$	5390	0.011	2_2^+	1643(2)	1.1	2.88(14)	3.4
$7/2_3^+$	5621	0.057	2_1^+	3020(3)	7.4	1.08(14)	3.1
$3/2_4^+$	5658	0.010	1_1^+	1643(2)	0.8	2.88(14)	3.8
$5/2_4^+$	5767	0.360	0_1^+	5276(5)	11.5	17.6(3)	7.4
			2_1^+	3153(4)	12.2	0.44(10)	7.8
			2_2^+	2008(2)	20.9	10.0(2)	13.4
$7/2_4^+$	6047	0.022	2_1^+	3432(3)	2.1	0.89(11)	9.8
$5/2_5^+$	6180	0.047	2_1^+	3561(11)	4.5	3.6(8)	30.8
$7/2_5^+$	6533	0.044	2_1^+	3902(3)	3.4	2.22(14)	11.3
$3/2_5^+$	6640	0.023	0_1^+	6145(7)	0.5	0.51(12)	5.4
$7/2_6^+$	6665	0.186	2_1^+	4030(3)	14.7	7.0(2)	4.6
			2_2^+	2881(3)	0.4	0.99(13)	0.13
$7/2_7^+$	7050	0.050	2_2^+	3249(4)	1.9	1.17(15)	2.6
			3_2^+	1300(13)	0.9	0.7(11)	1.3

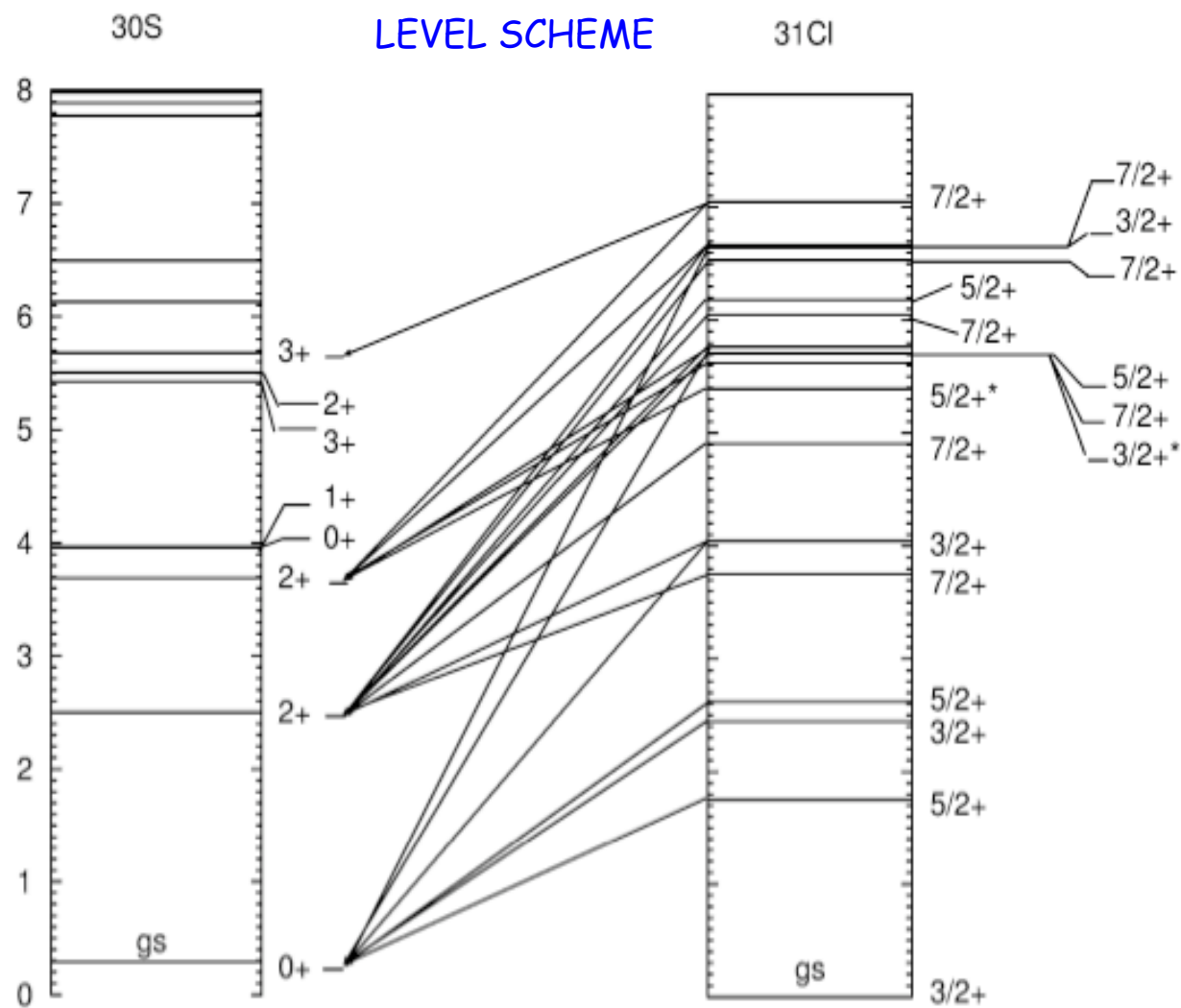


Fig. 4. The proposed level scheme and decay mechanism for ^{31}Cl . Energy levels are given in MeV, relative to the ground state of ^{31}Cl . (*) means ambiguous assignment; see text for details.

SUMMARY

We have revised the physics concepts behind the beta delayed particle emission process:

- Basic ideas about the exotic decay process
- Exotic decays are an important source of spectroscopic information: level energies, spins, $B(F)$ and $B(GT)$ values, etc
- Technical aspects to measure these decay modes
- Status of beta delayed nucleon emission
- Basic ideas for beta decay and isospin
- Simple models for particle emission (Gamow states, R-Matrix,...)

THANKS FOR YOUR ATTENTION...