

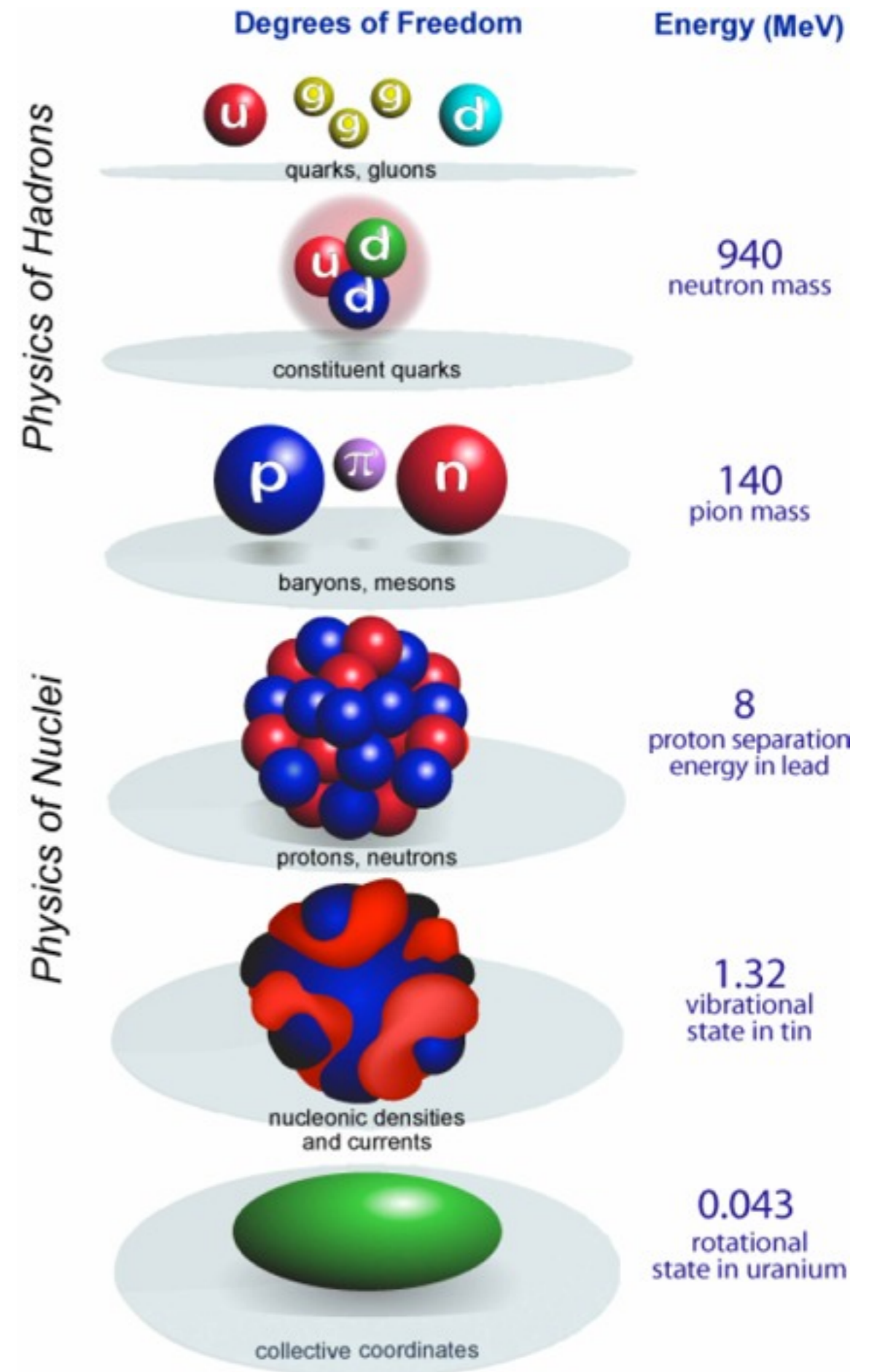
# A microscopic approach to nuclear dynamics

Cédric Simenel  
CEA/Saclay, France

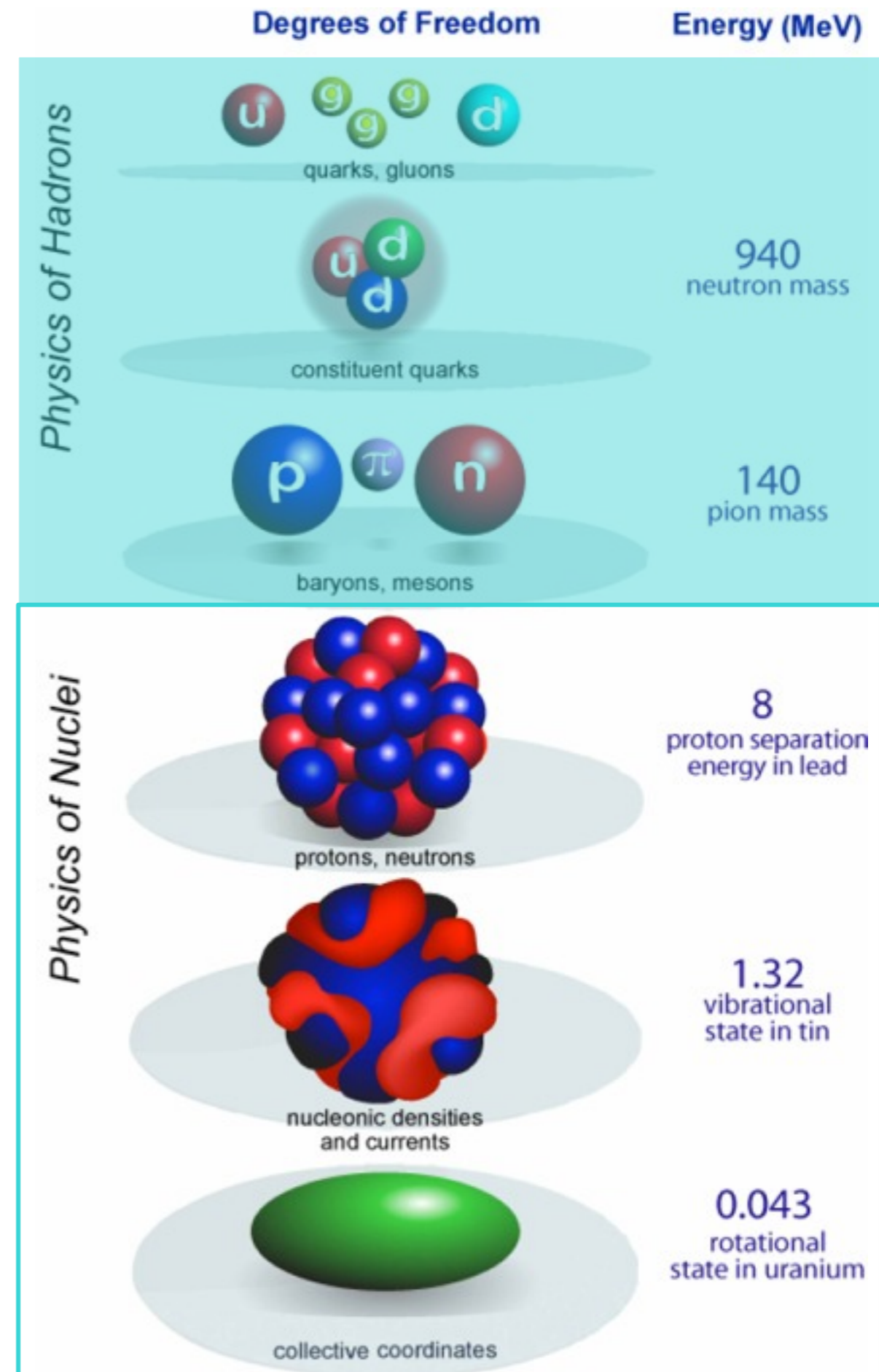
# Introduction

- Quantum dynamics of complex systems (nuclei, molecules, BEC, atomic clusters...)
- Collectivity: from vibrations to collisions
- Interplay with single-particle d.o.f. (Giant Resonance decay, Competition fusion/transfer...)
- Quantum many-body problem

# Microscopic, up to which scale?



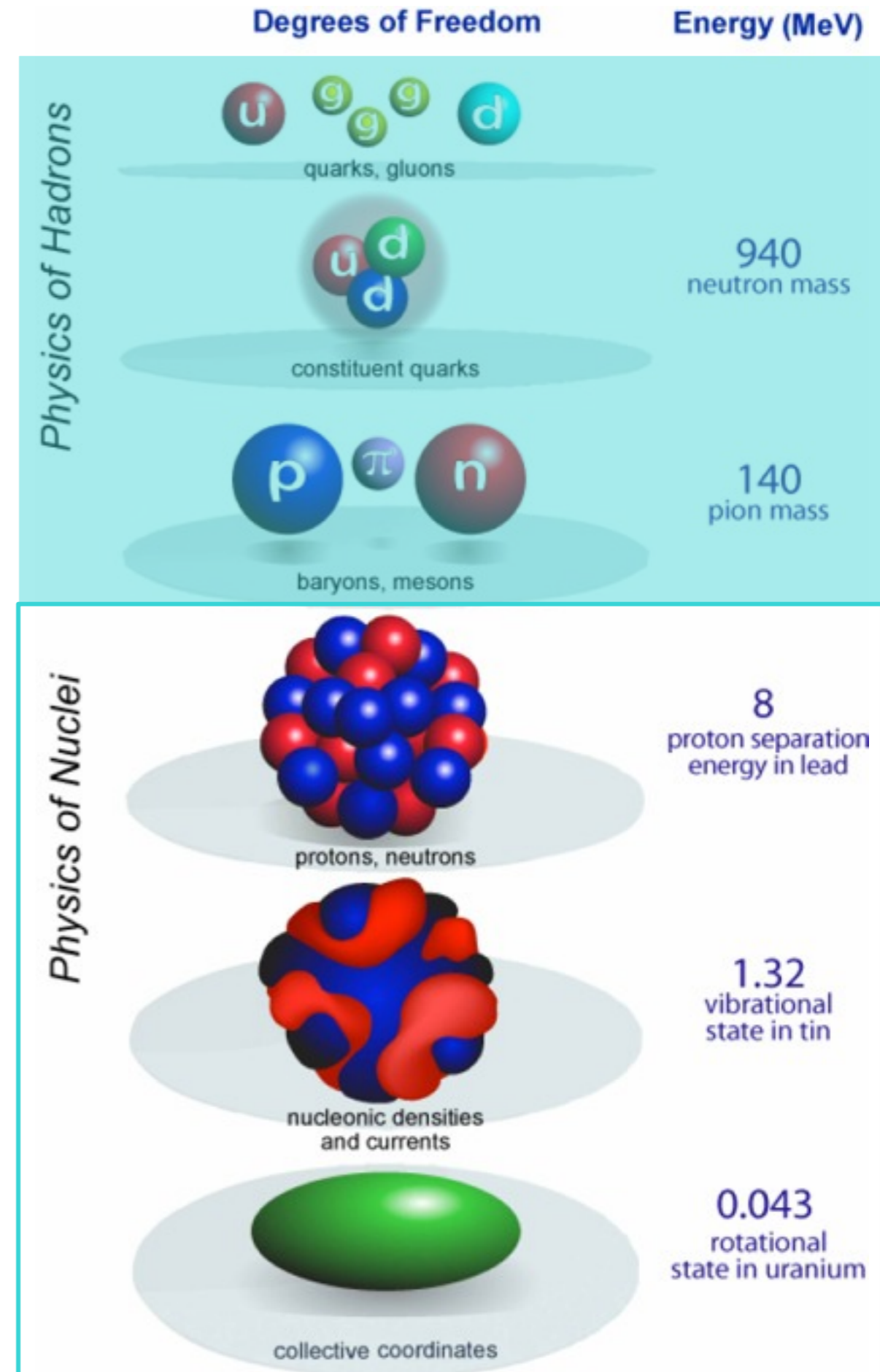
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## Ingredients of the model:

- nucleon-nucleon interaction
- approximations of the many-body problem



# Hamiltonian

$$\hat{H} = \hat{T} + \hat{V}$$

$$\hat{T} = \sum_{i=1}^N \frac{\hat{p}_{(i)}^2}{2m}$$

$$\hat{V} = \frac{1}{2} \sum_{i \neq j=1}^N \hat{v}(i, j)$$

# Hamiltonian

Separation into a mean field  $U$  and a residual interaction

$$\hat{V} = \hat{U} + \hat{V}_{res} \quad \text{with} \quad \hat{U} = \sum_{i=1}^N \hat{u}(i)$$

Mean-Field approximation: neglect  $V_{res}$

Each nucleon is assumed to evolve independently in the MF generated by the other nucleons. The interactions are «averaged» into a MF.

# Hamiltonian

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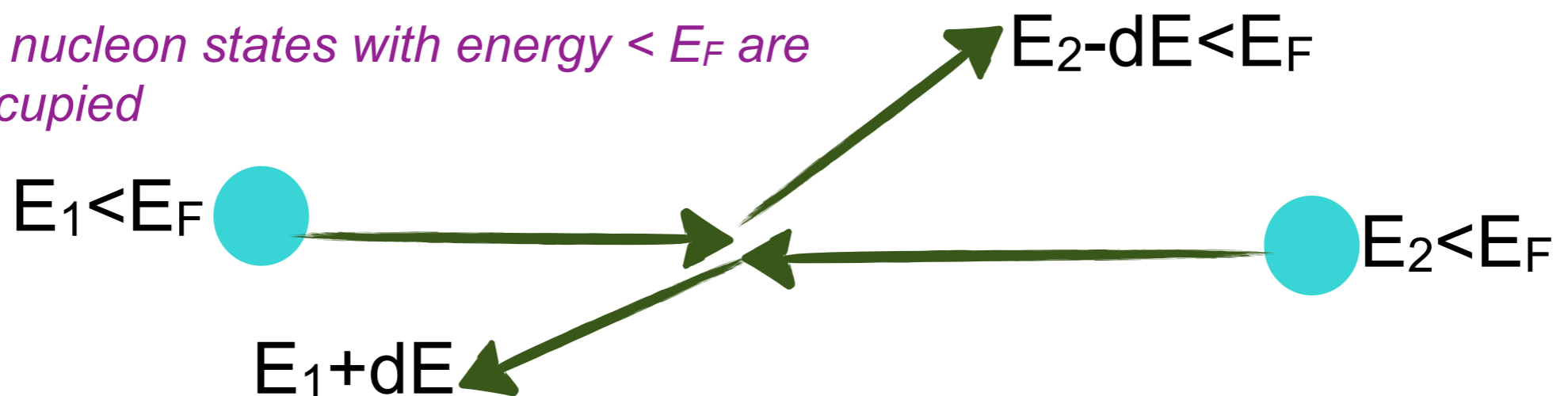
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Justified by the mean free path of a nucleon in the nucleus of the order of the size of the nucleus, thanks to the Pauli principle:

*All nucleon states with energy  $< E_F$  are occupied*





# Hamiltonian

Separation into a mean field  $U$  and a residual interaction

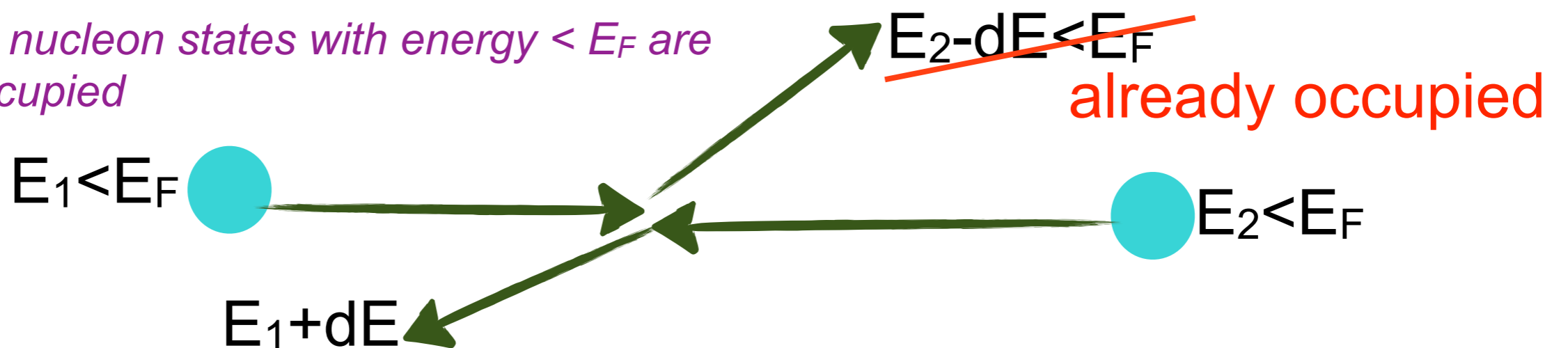
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# Hartree-Fock

- Mean-field determined from the interaction

$$\hat{u}(i) = \sum_{j=1}^N \langle \varphi_j | \hat{v}(i, j) | \varphi_j \rangle$$

- Self-consistent mean-field

$$\hat{u} \equiv \hat{u}[\varphi_1, \varphi_2, \dots, \varphi_N] \equiv \hat{u}[\rho]$$

- HF equation

$$\left( \frac{\hat{p}^2}{2m} + \hat{u}[\rho] \right) |\varphi_i\rangle \equiv \hat{h}[\rho] |\varphi_i\rangle = e_i |\varphi_i\rangle \quad \text{for } i = 1, 2, \dots, N$$

# Practical aspects of HF calculations

## Imaginary time method

*for a ground state with no self-consistency (  $\hat{h} \neq \hat{h}[\rho]$  ):*

$$|\varphi_{init}\rangle = \sum_i C_i |\varphi_i\rangle$$

$$e^{-i\hat{h}t} \rightarrow e^{-\beta\hat{h}} |\varphi_{init}\rangle = \sum_i C_i e^{-\beta E_i} |\varphi_i\rangle \xrightarrow{\beta \rightarrow \infty} C_0 e^{-\beta E_0} |\varphi_0\rangle$$

# Practical aspects of HF calculations

## HF calculations

Start with Nilsson or harmonic oscillator w.f.

Imaginary time method

*for a  $N$  lowest states of  $h[\rho]$ : iterative process (evolution with  $\Delta\beta$ )*

$$\{|\varphi_1^{[n]}\rangle \cdots |\varphi_N^{[n]}\rangle\} \Rightarrow \rho^{[n]} \Rightarrow \hat{h}^{[n+1]} = \hat{h}[\rho^{[n]}]$$

$\uparrow$   $\downarrow$

$$|\varphi_i^{[n+1]}\rangle = \frac{1}{\mathcal{N}_i} \left( |\varphi'_i\rangle - \sum_{j=0}^{i-1} \langle \varphi_j^{[n+1]} | \varphi'_i \rangle |\varphi_j^{[n+1]}\rangle \right) \Leftarrow |\varphi'_i\rangle = (1 - \Delta\beta \hat{h}^{[n+1]}) |\varphi_i^{[n]}\rangle$$

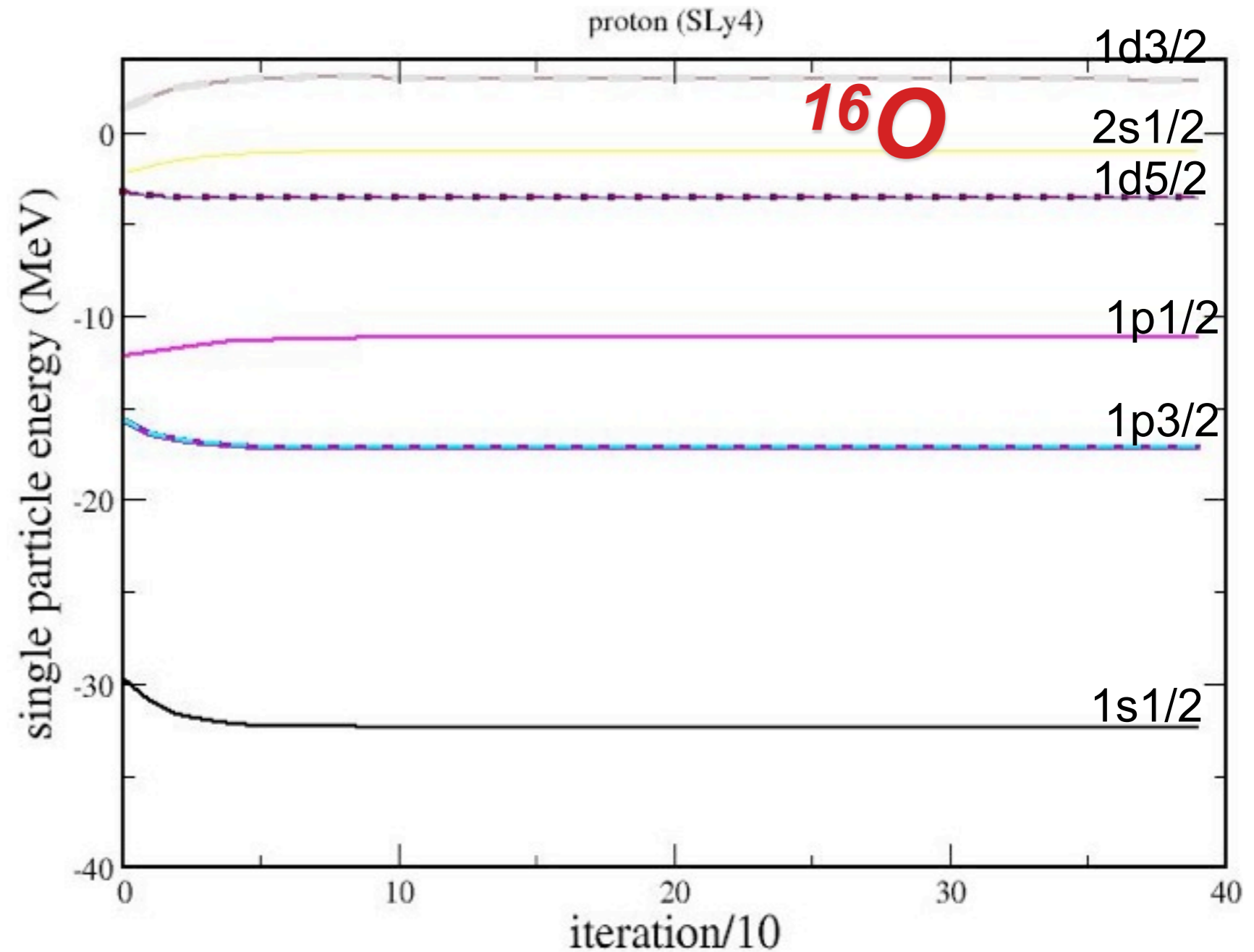
**(Graham-Schmidt orthonormalization)**

# Practical aspects of HF calculations

## HF calculations

Start with Nilsson w.f.

Imaginary time method



ev8, P. Bonche et al., Comp. Phys. Com. 171, 49 (2005)

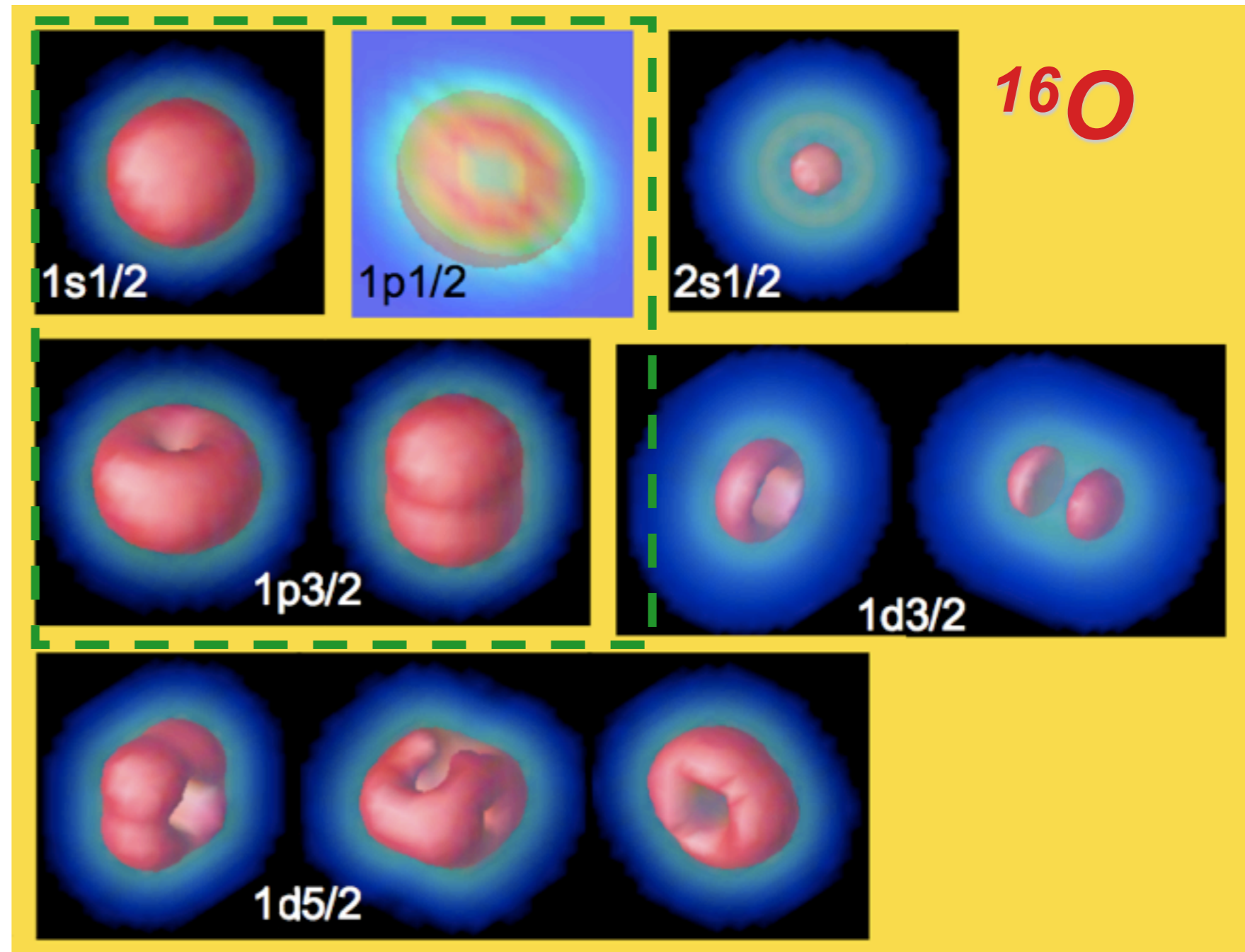
# Practical aspects of HF calculations

## HF calculations

Start with Nilsson w.f.

Imaginary time method

Occupied neutron w.f.

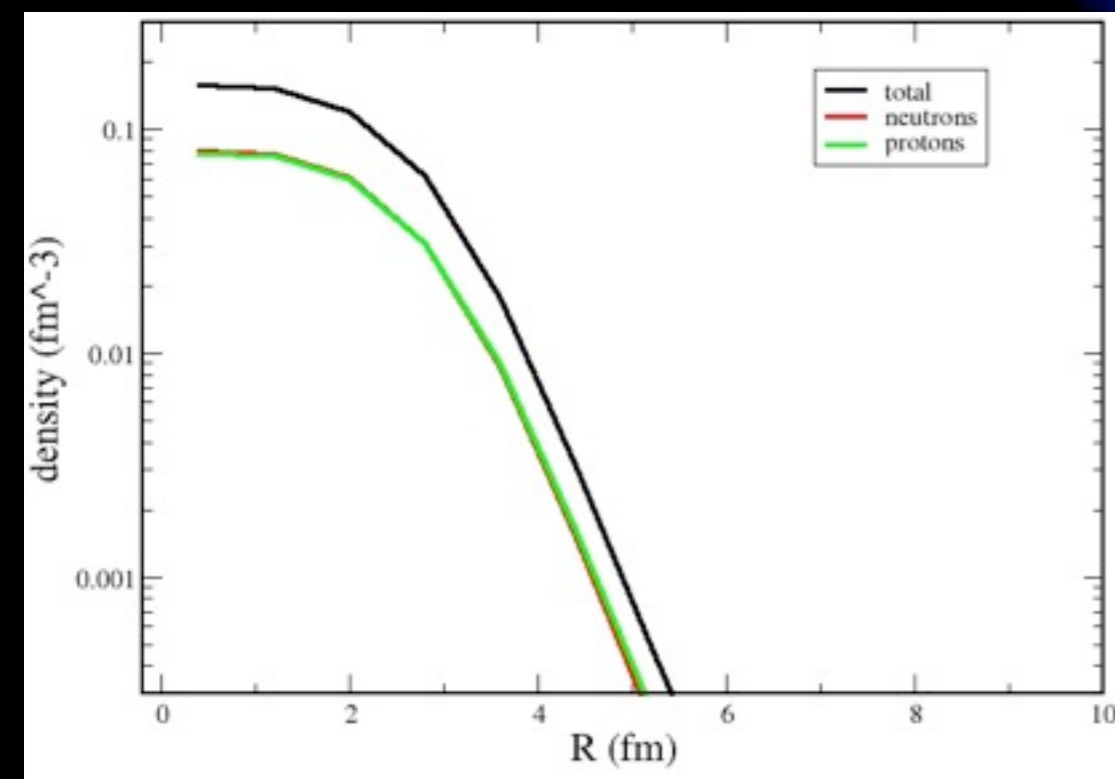
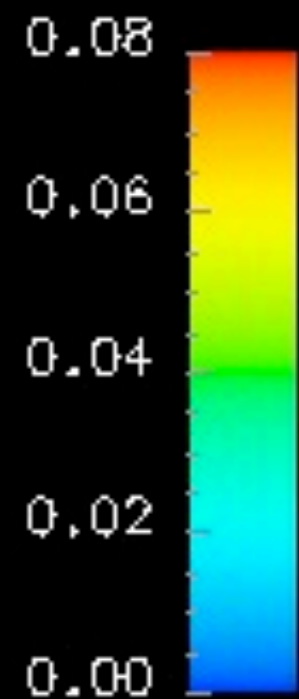
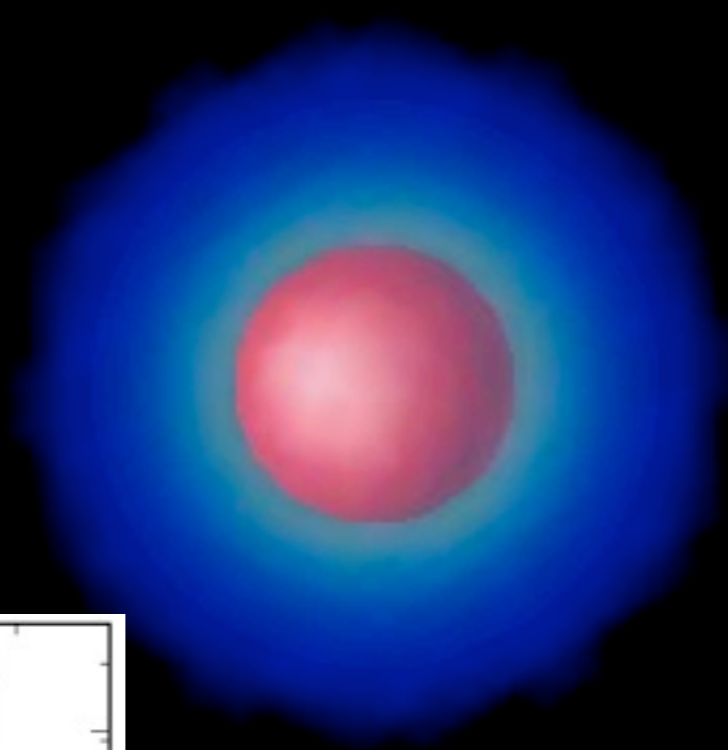


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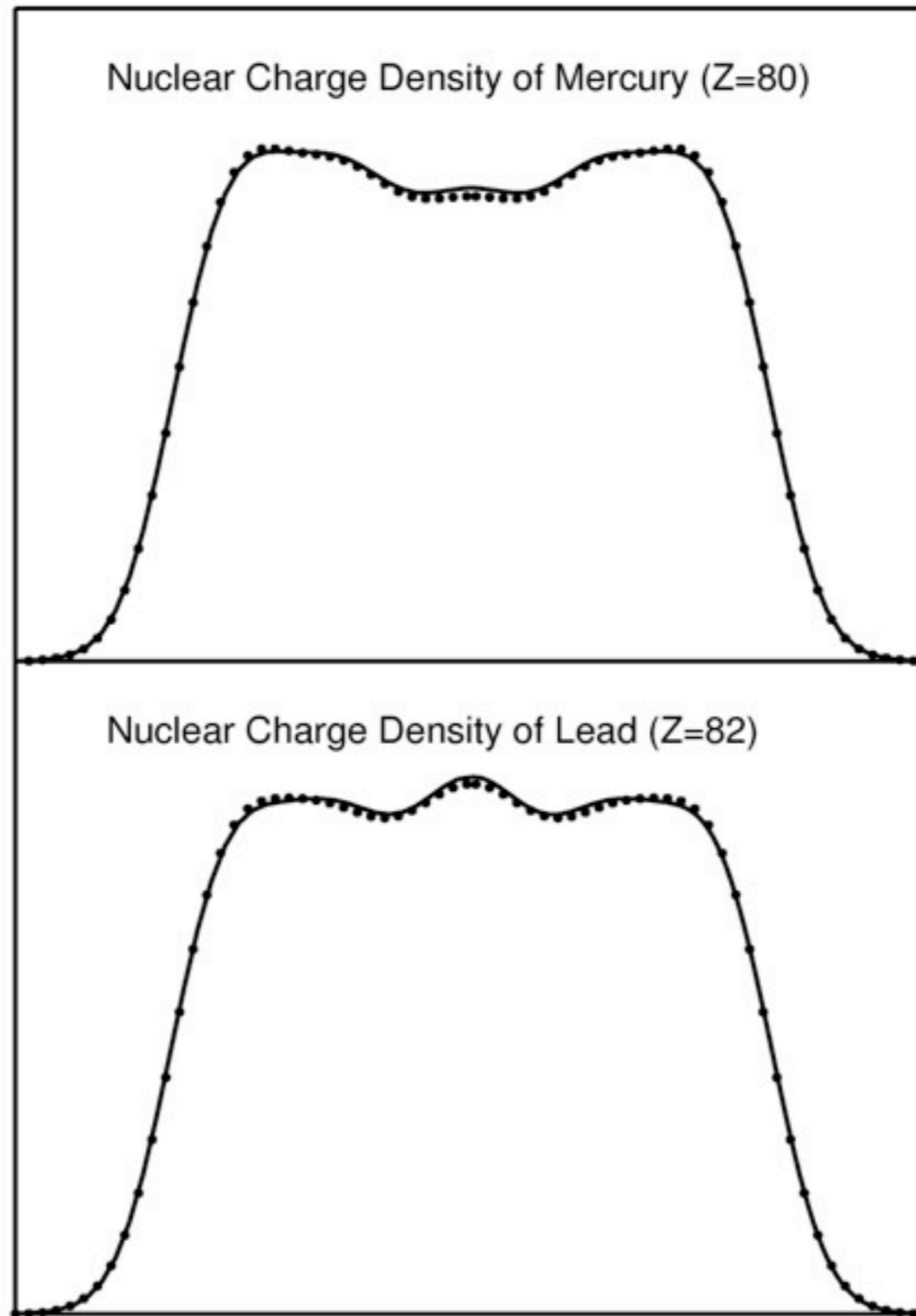
# Ground-state density from HF calculations

**$^{16}\text{O}$**

density ( $\text{fm}^{-3}$ )

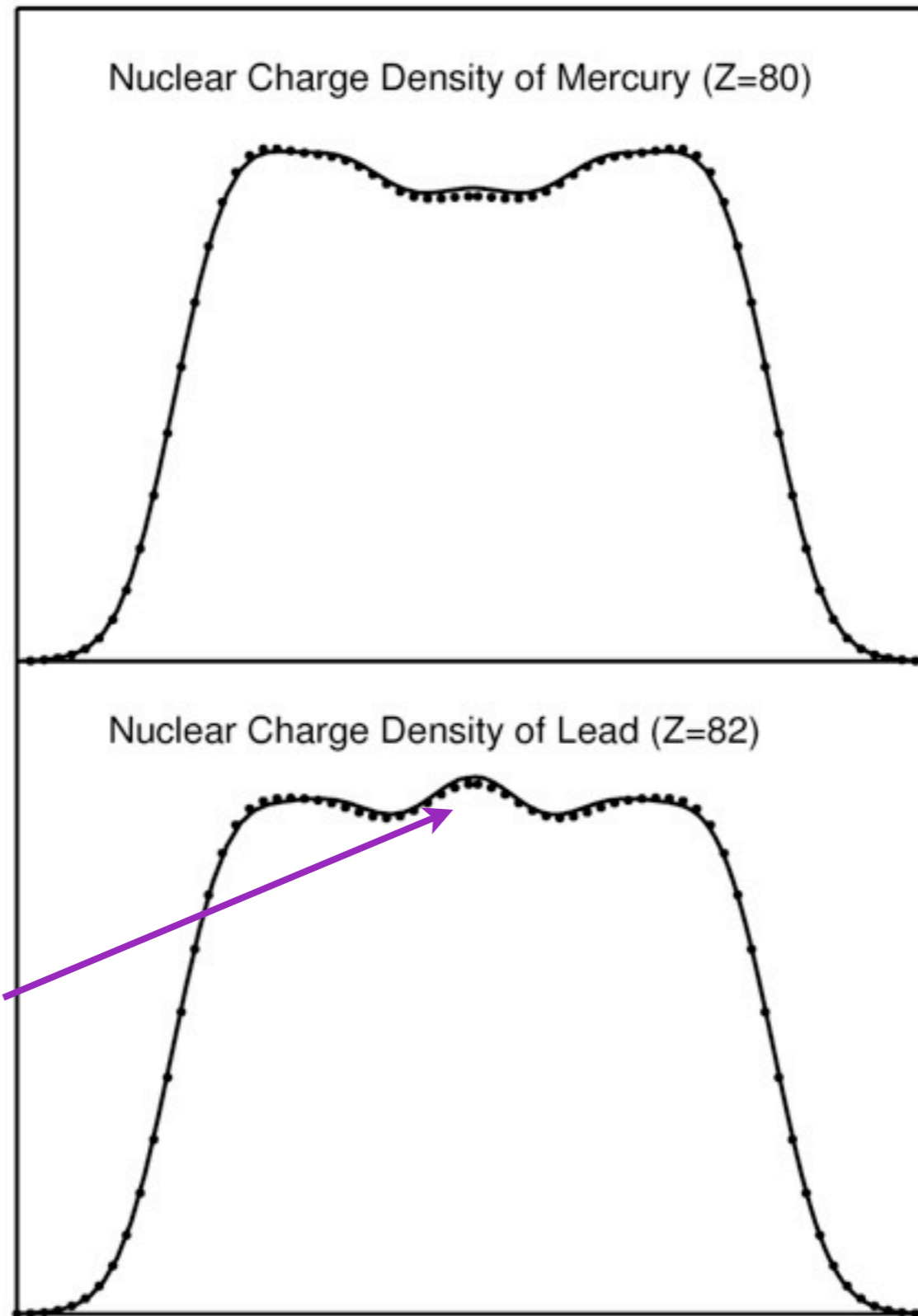


# *Ground-state density from HF calculations*





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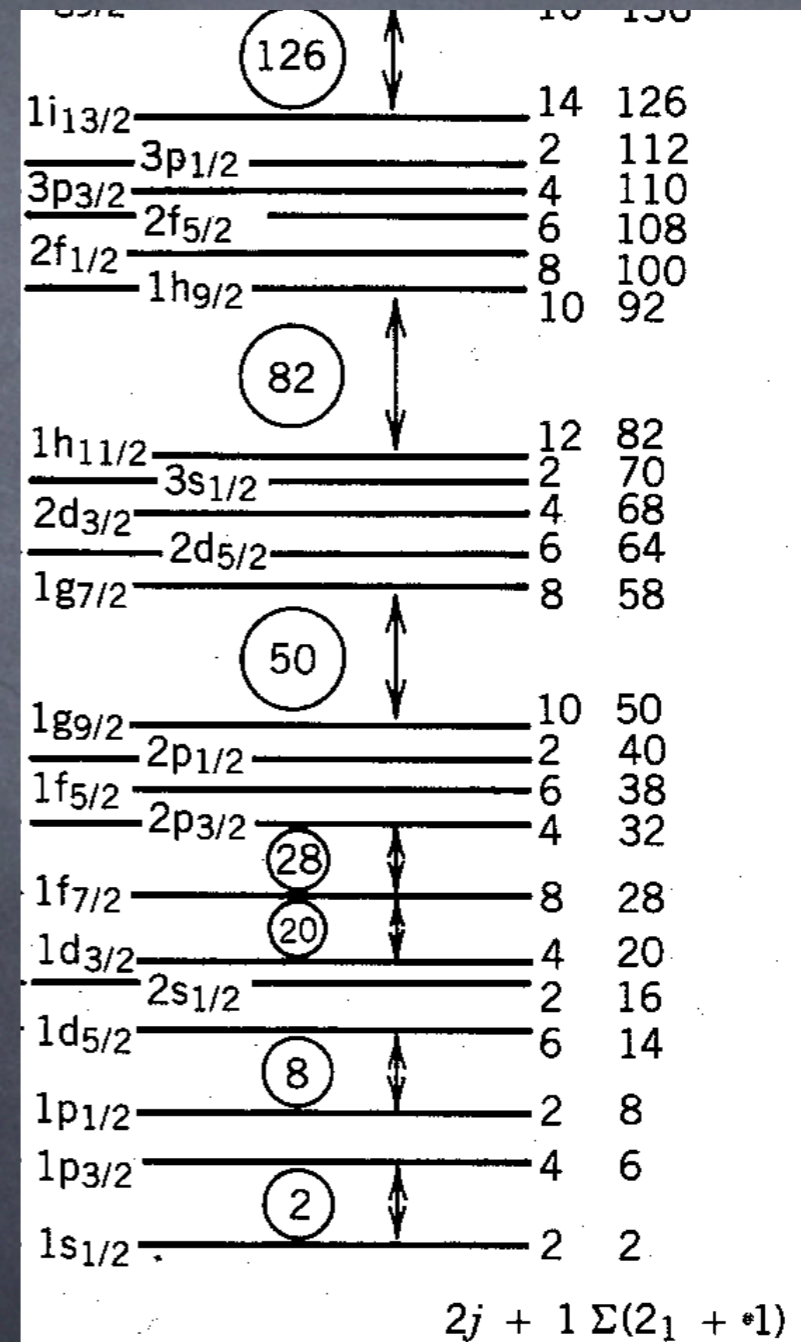
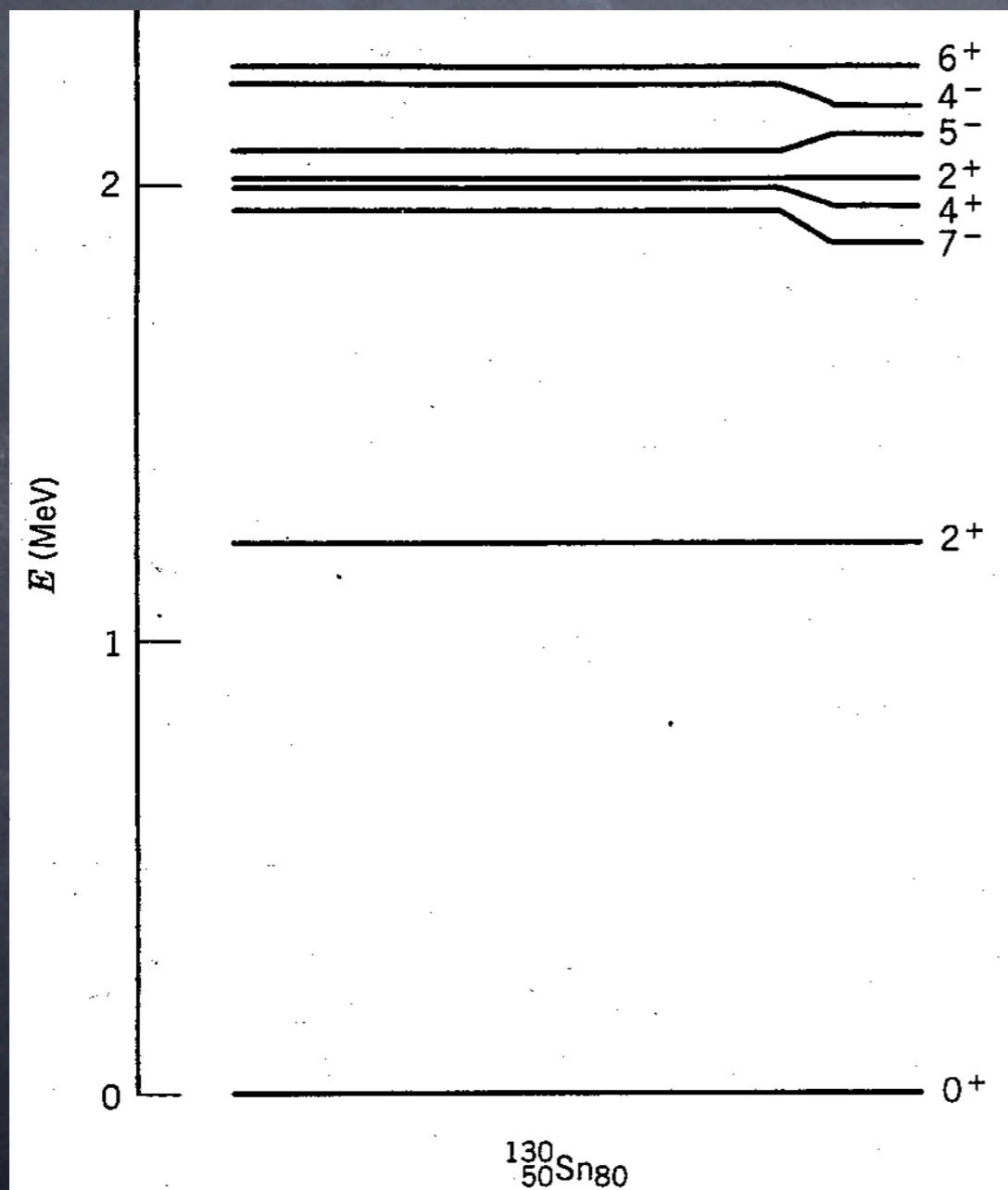


filling of 3s<sub>1/2</sub> shell

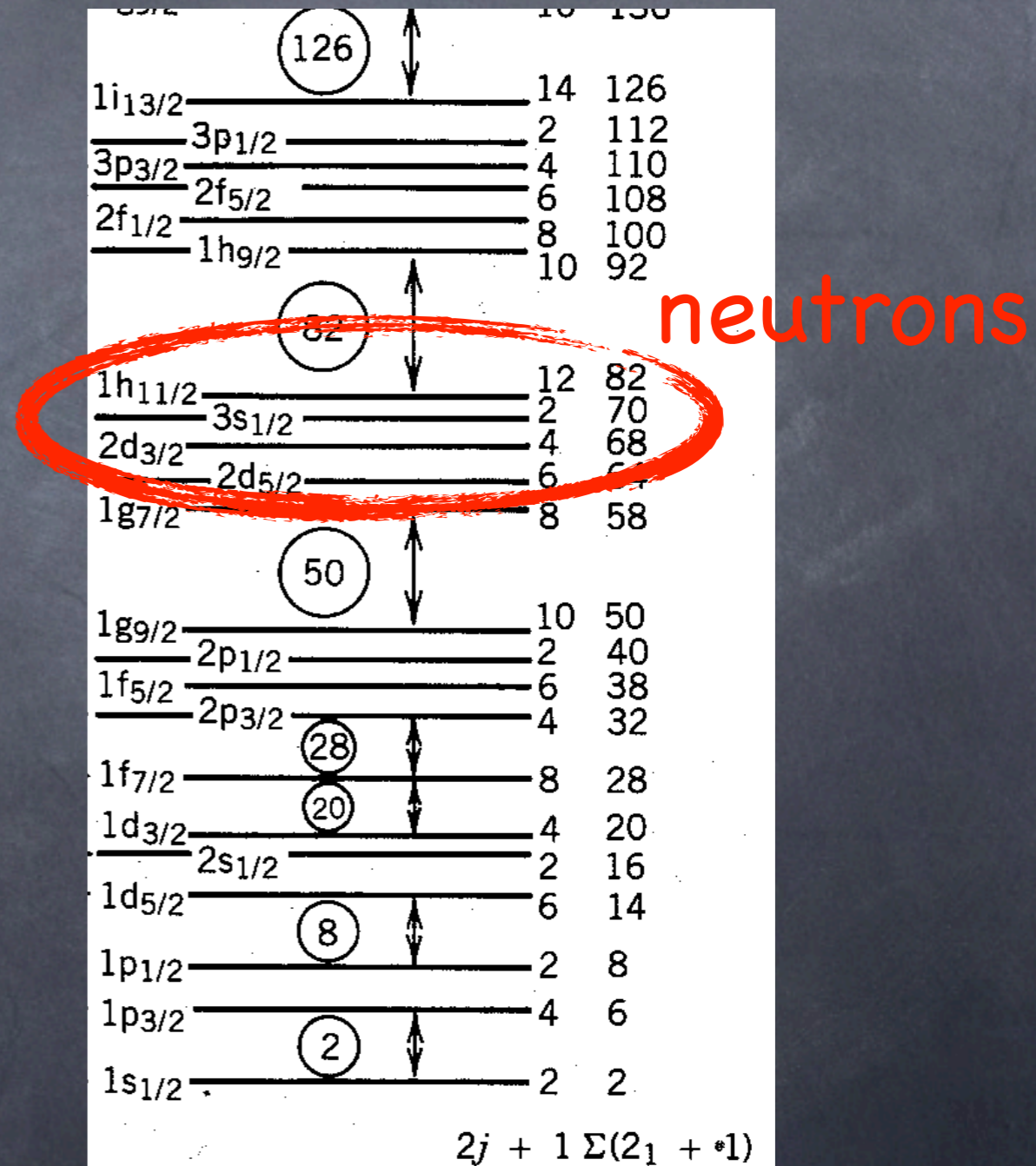
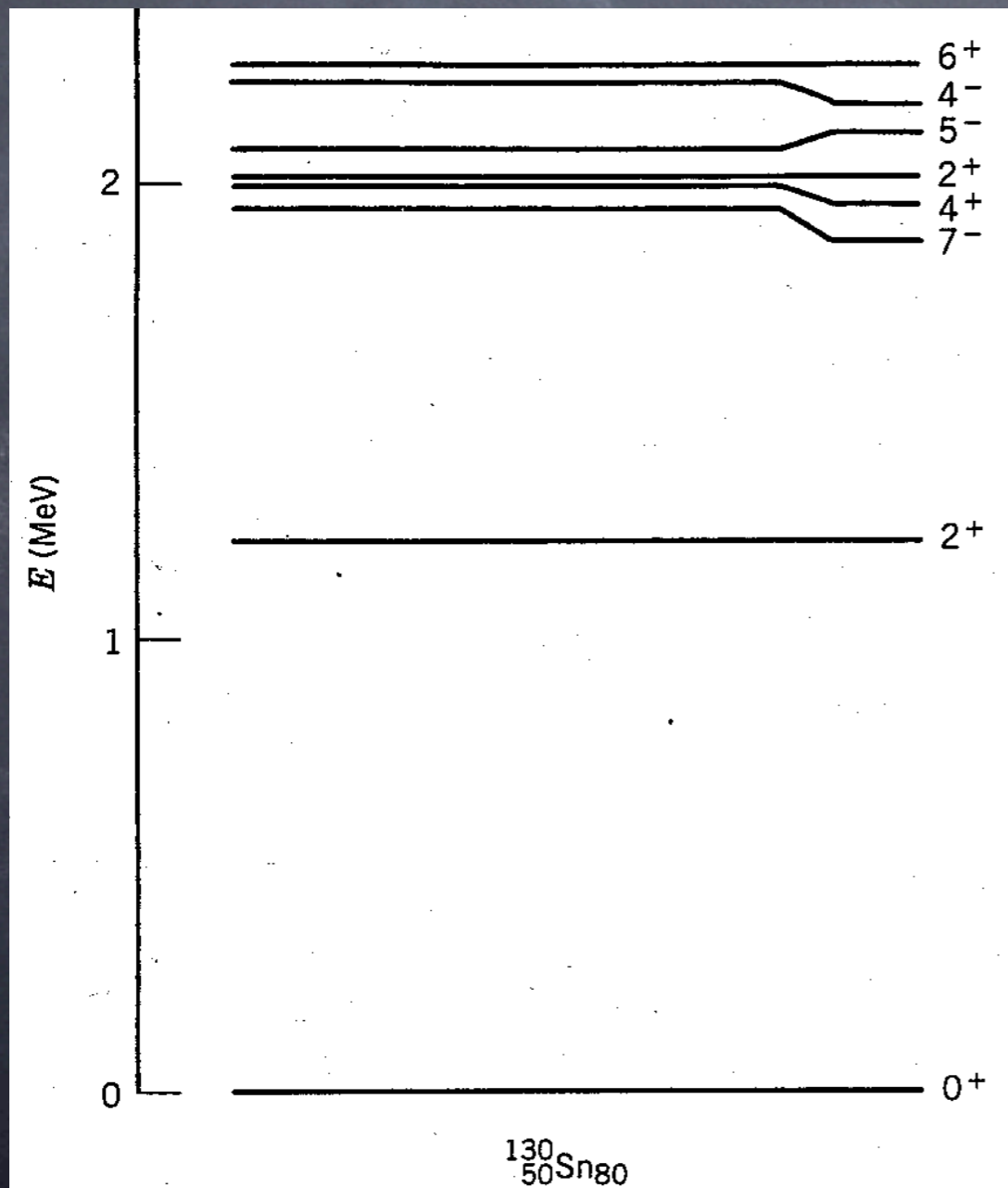
# Excited states

- single-particle excitations
- Low-lying collective vibrations
- Giant resonances

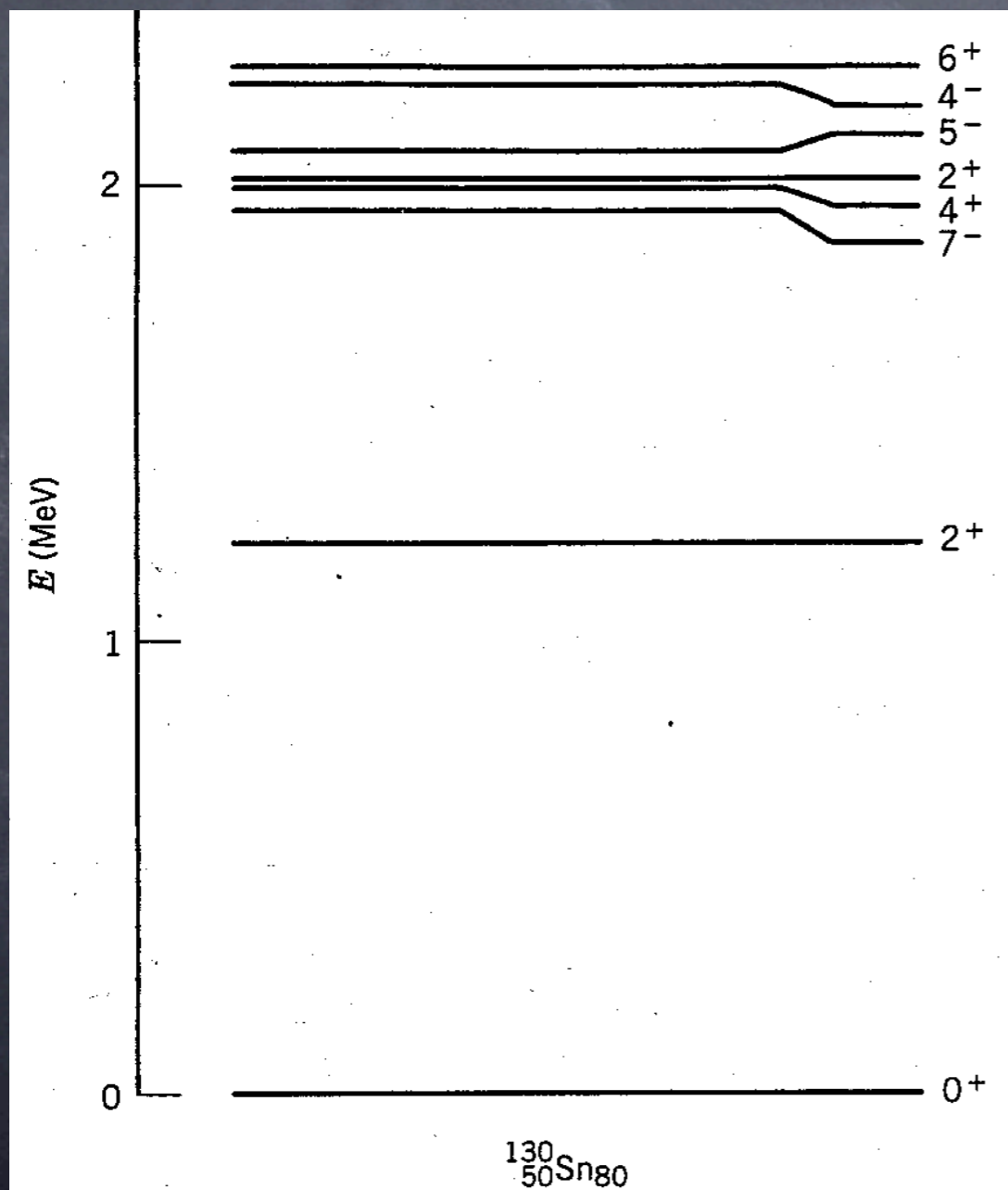
# Interpretation of $^{130}\text{Sn}$ ( $Z=50, N=80$ ) spectrum



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# Interpretation of $^{130}\text{Sn}$ spectrum



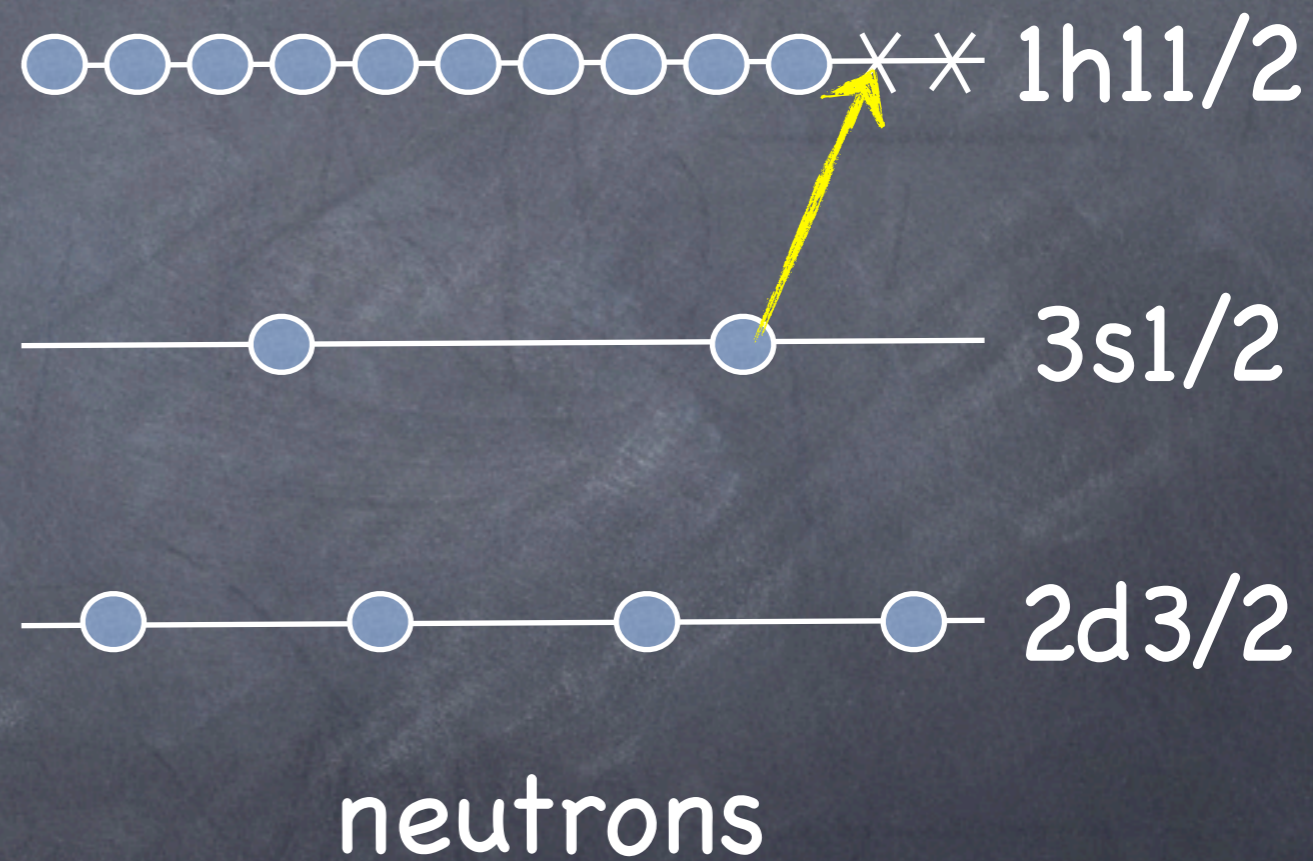
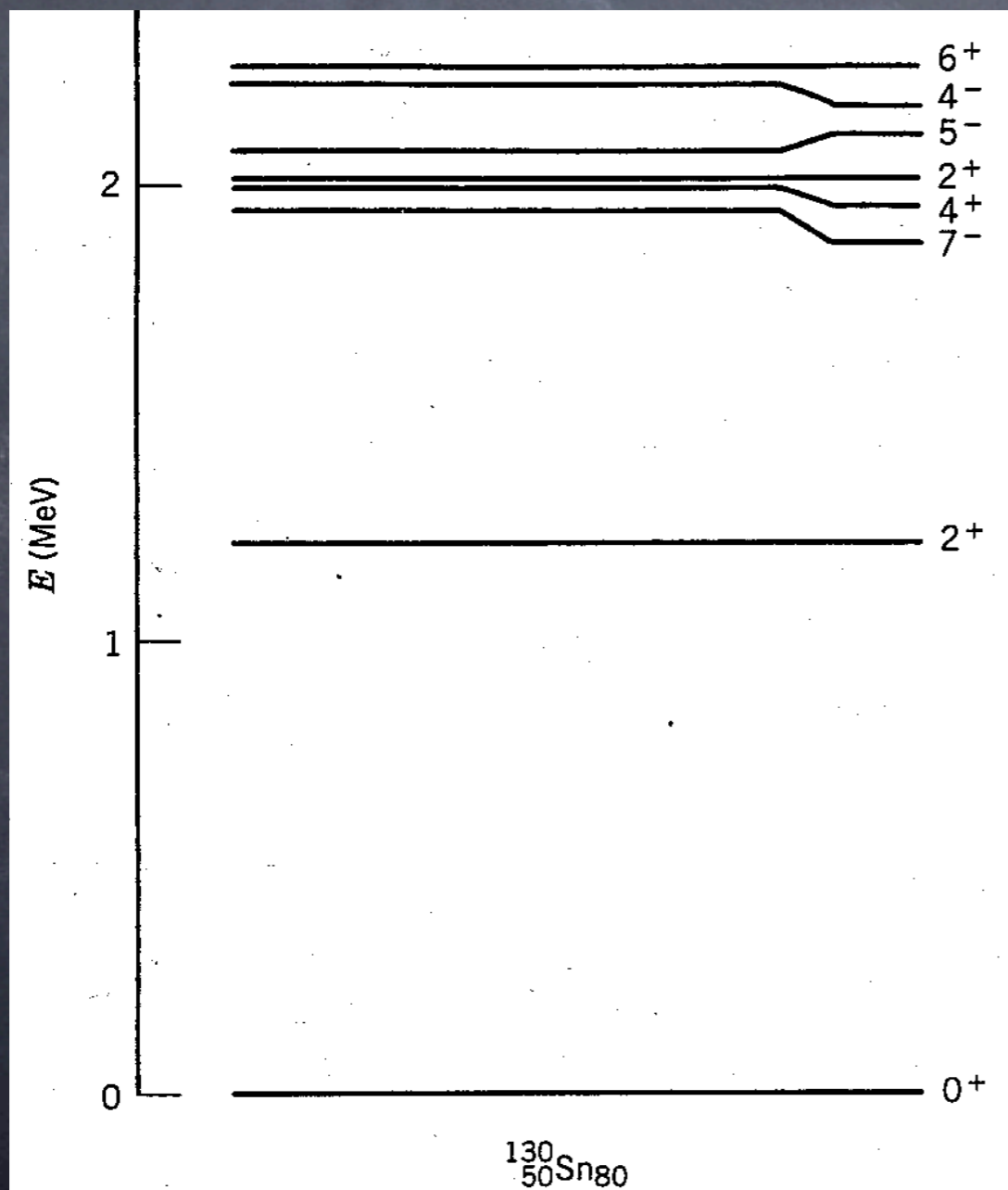
$\circ \circ \circ \circ \circ \circ \circ \circ \circ \circ \times \times$   $1h_{11/2}$

—  $\circ$  —  $\circ$  —  $3s_{1/2}$

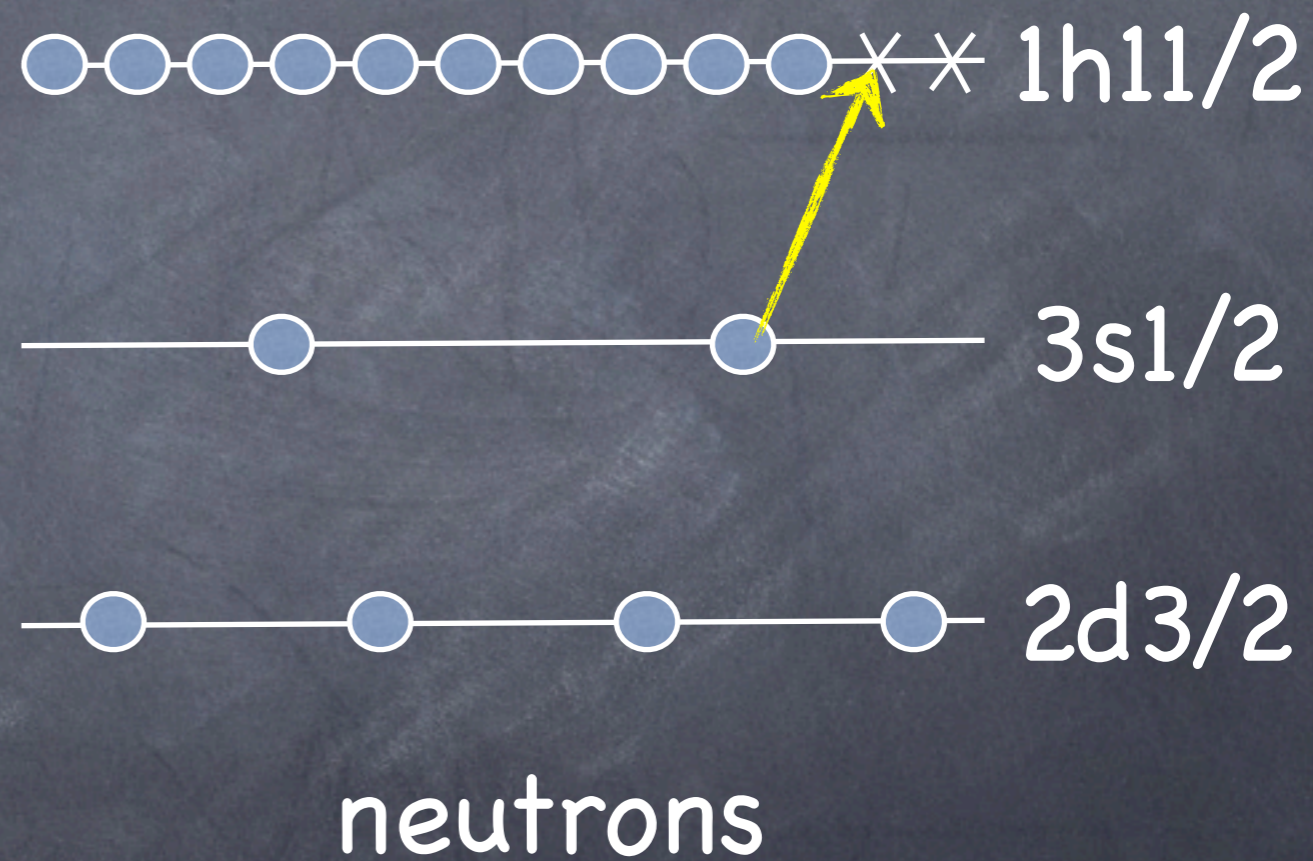
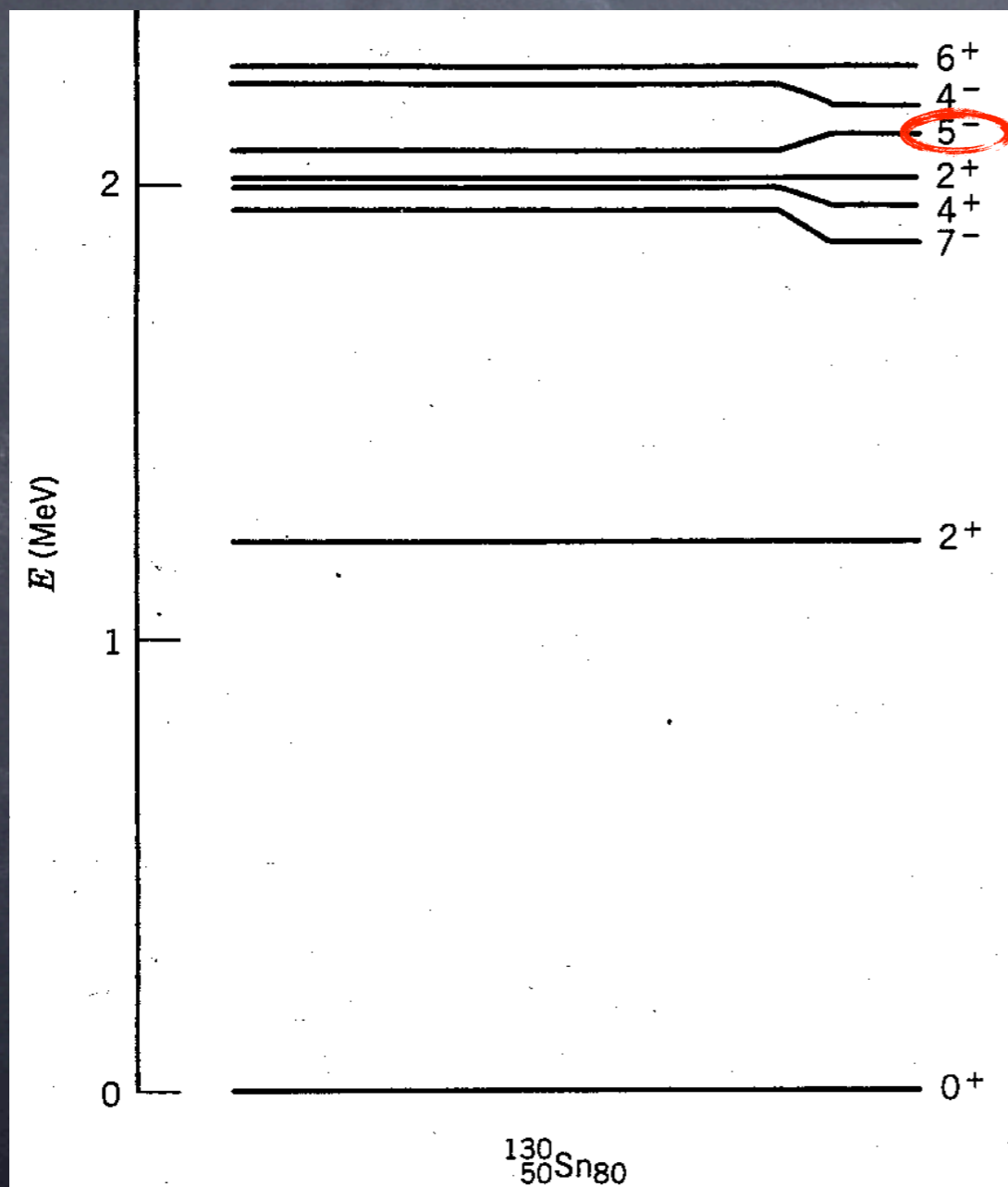
—  $\circ$  —  $\circ$  —  $\circ$  —  $\circ$  —  $2d_{3/2}$

neutrons

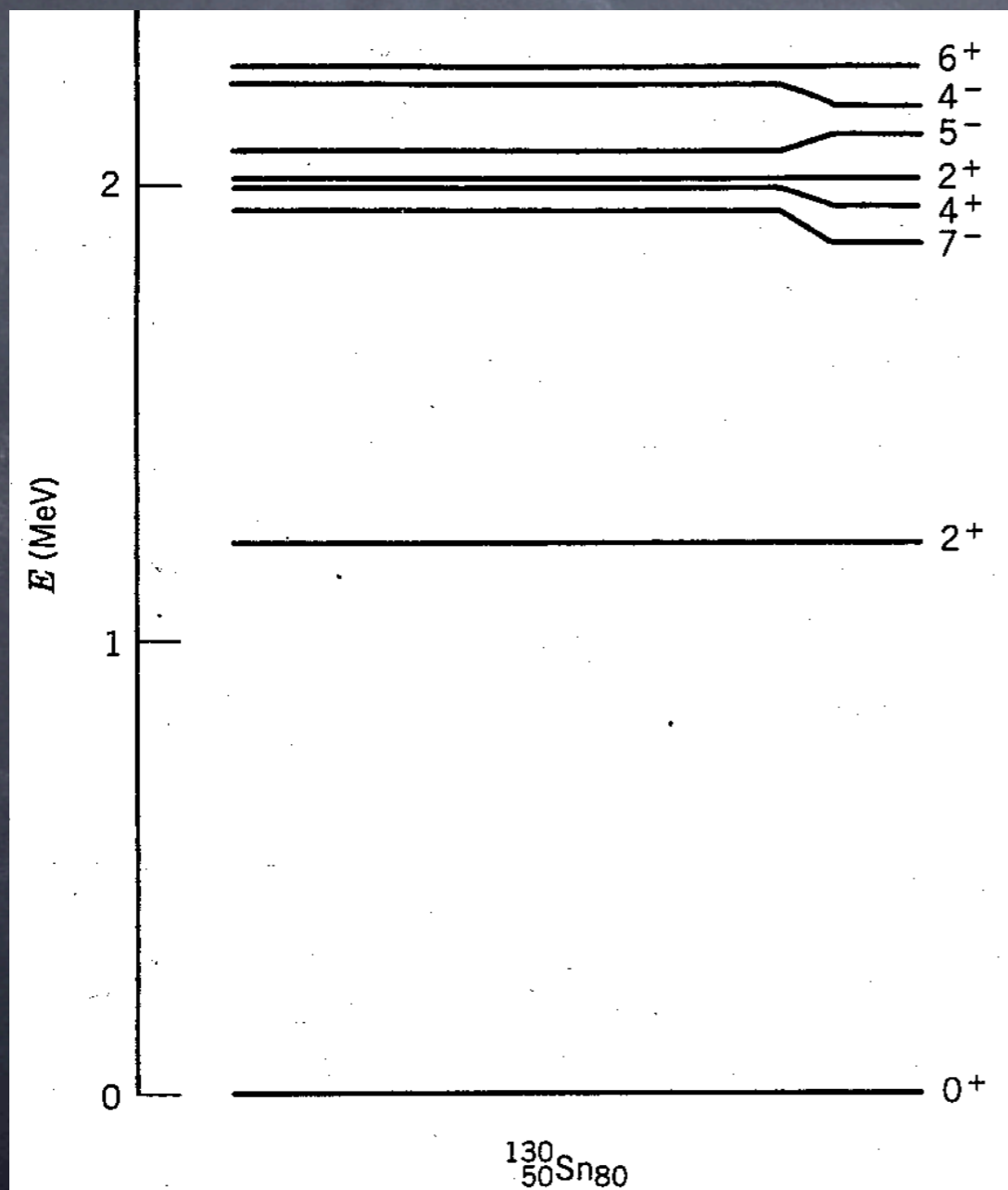
# Interpretation of $^{130}\text{Sn}$ spectrum



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$\circ \circ \circ \circ \circ \circ \circ \circ \circ \circ \times \times$   $1h_{11/2}$

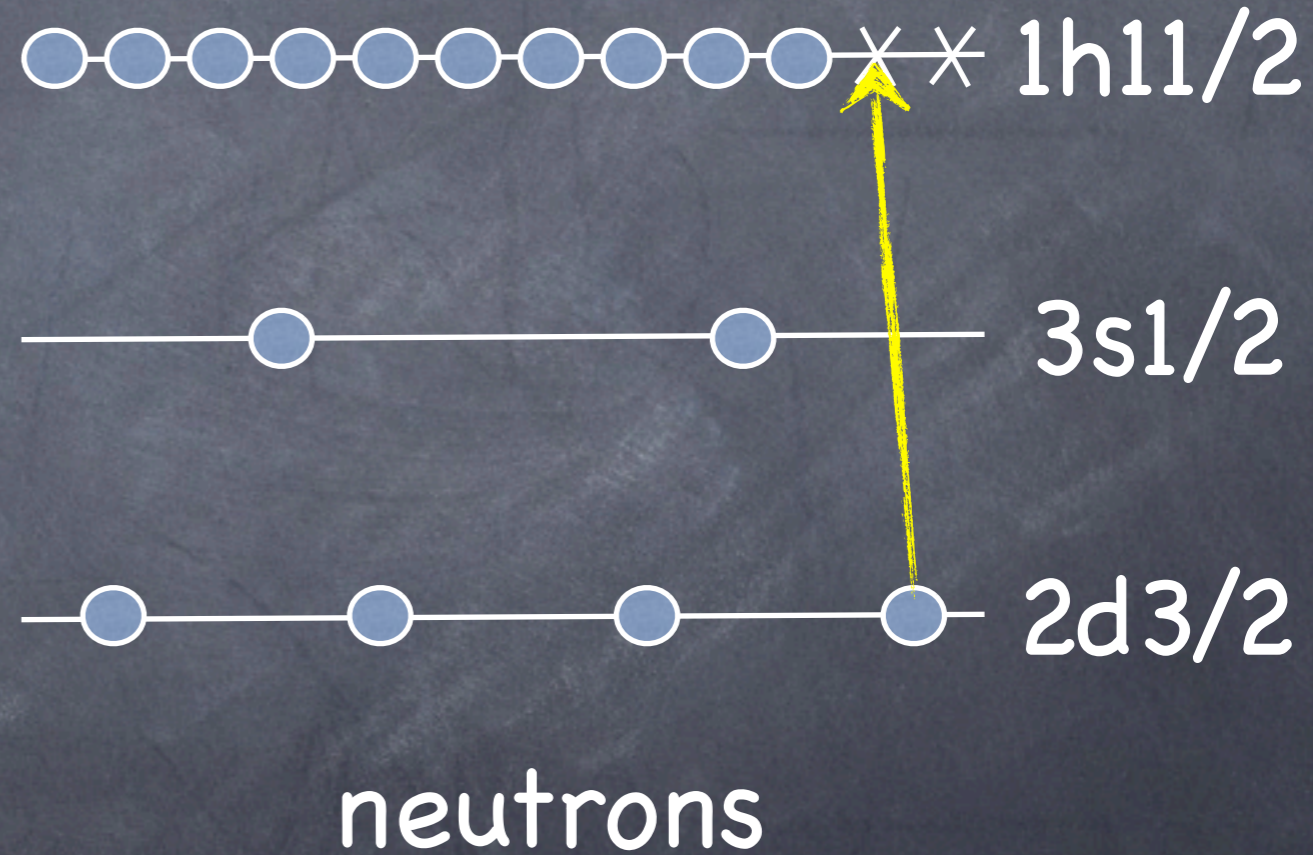
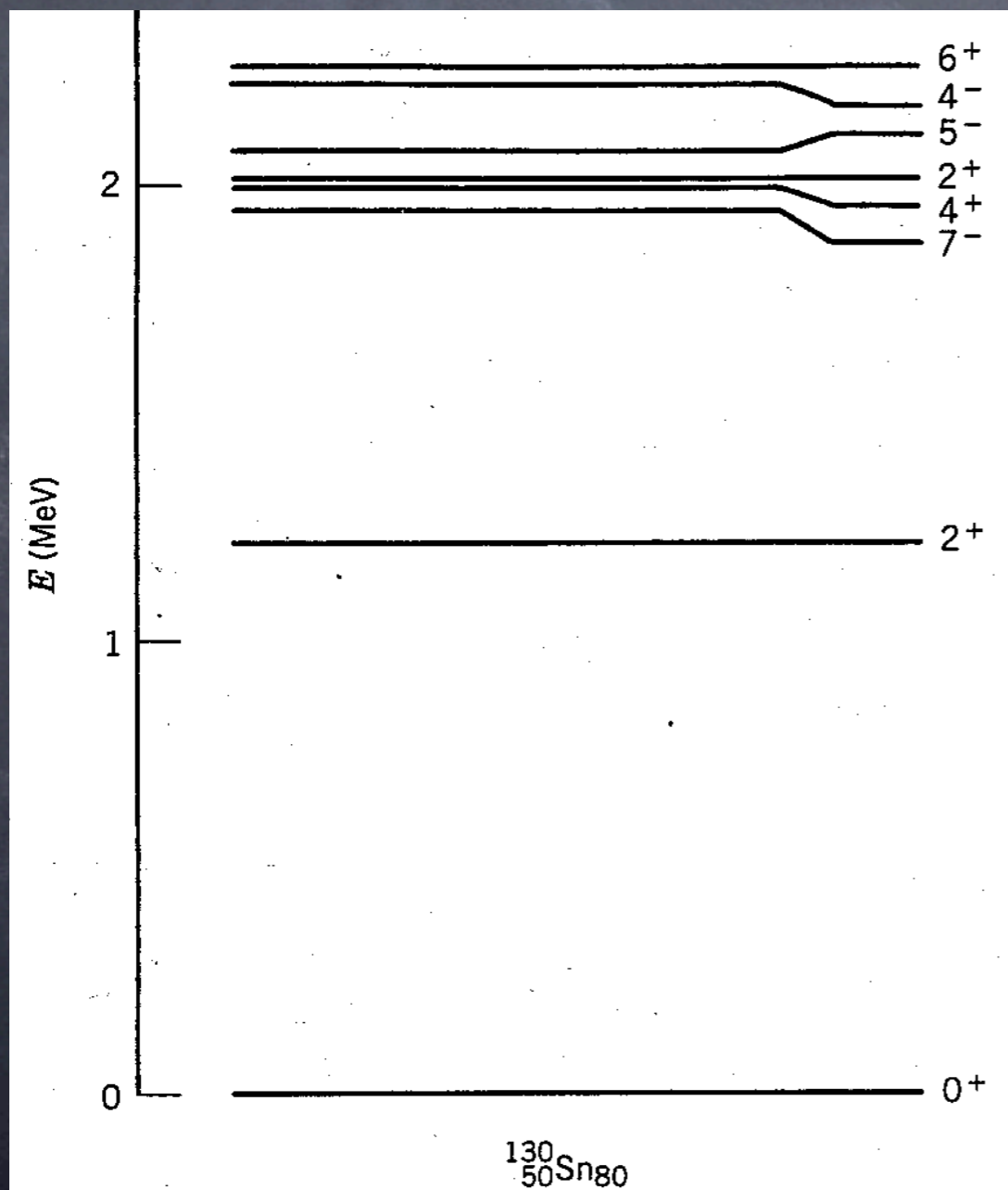
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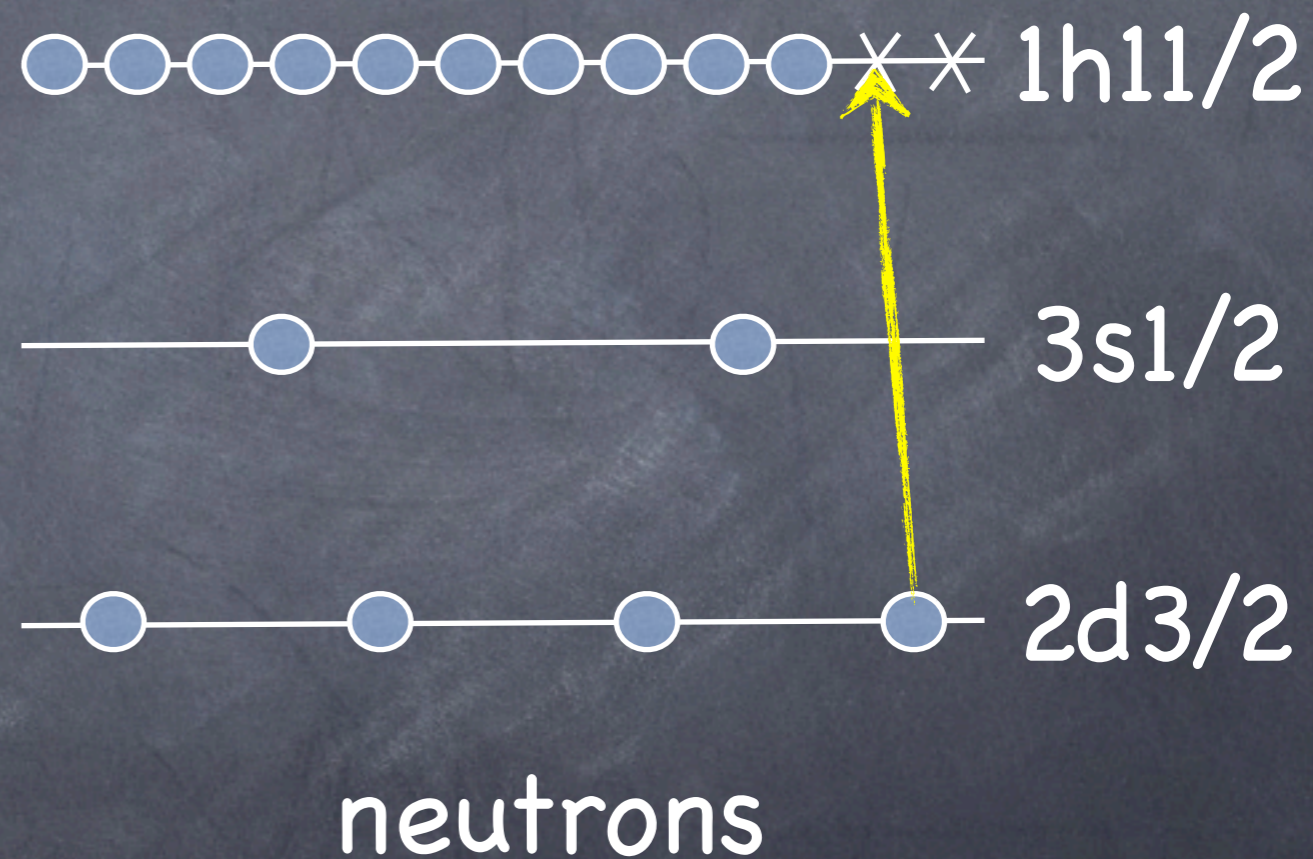
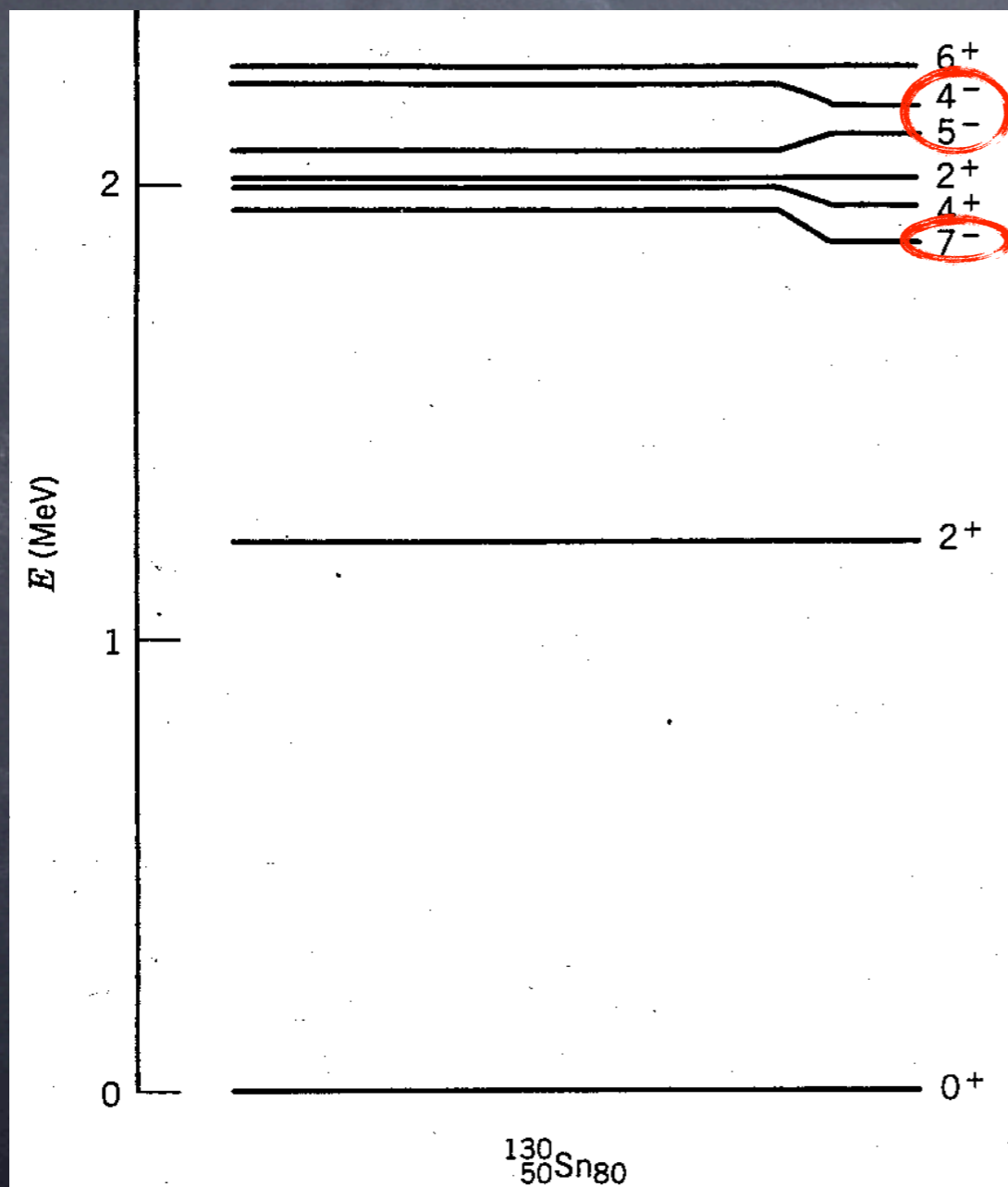
neutrons



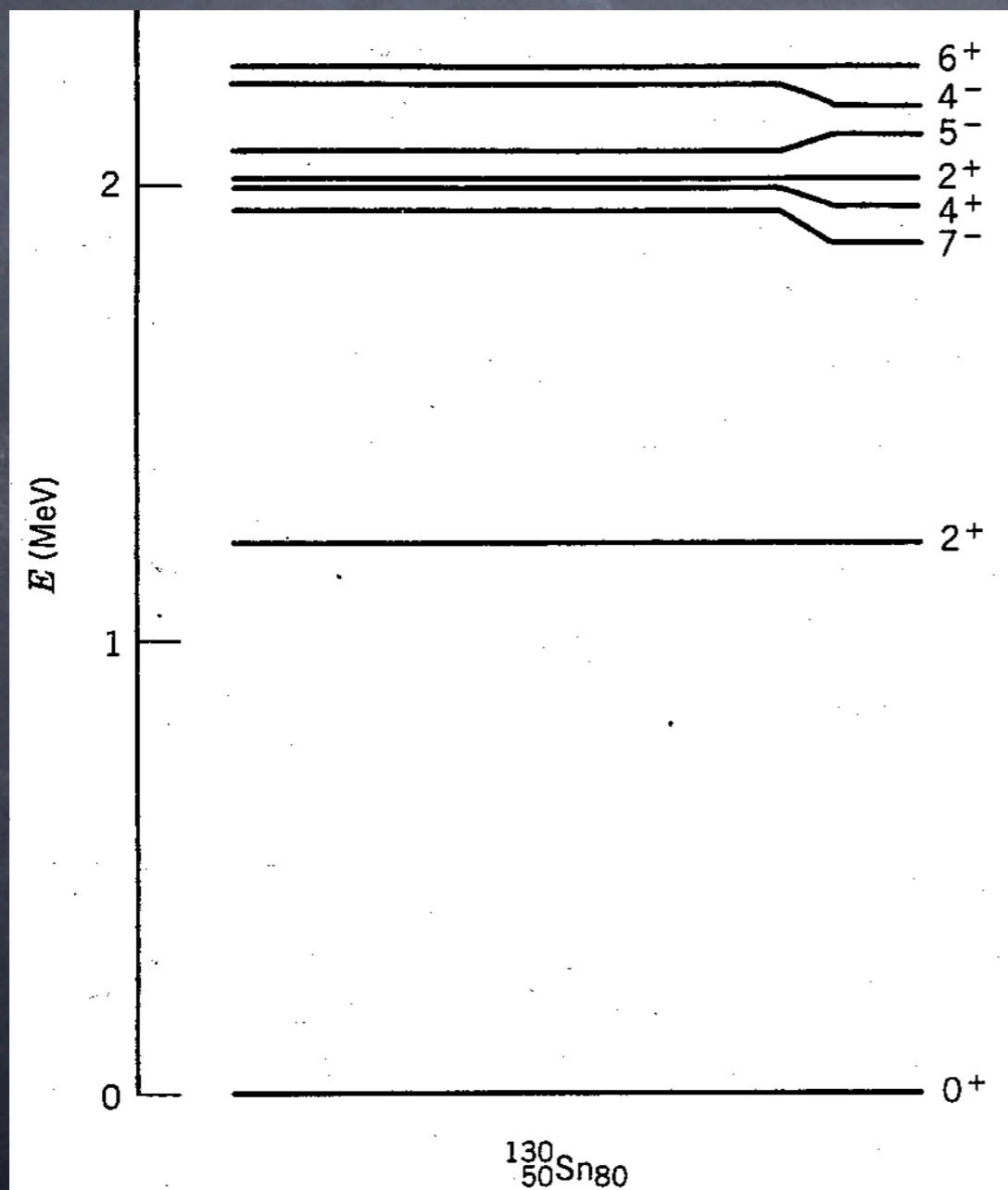
# Interpretation of $^{130}\text{Sn}$ spectrum



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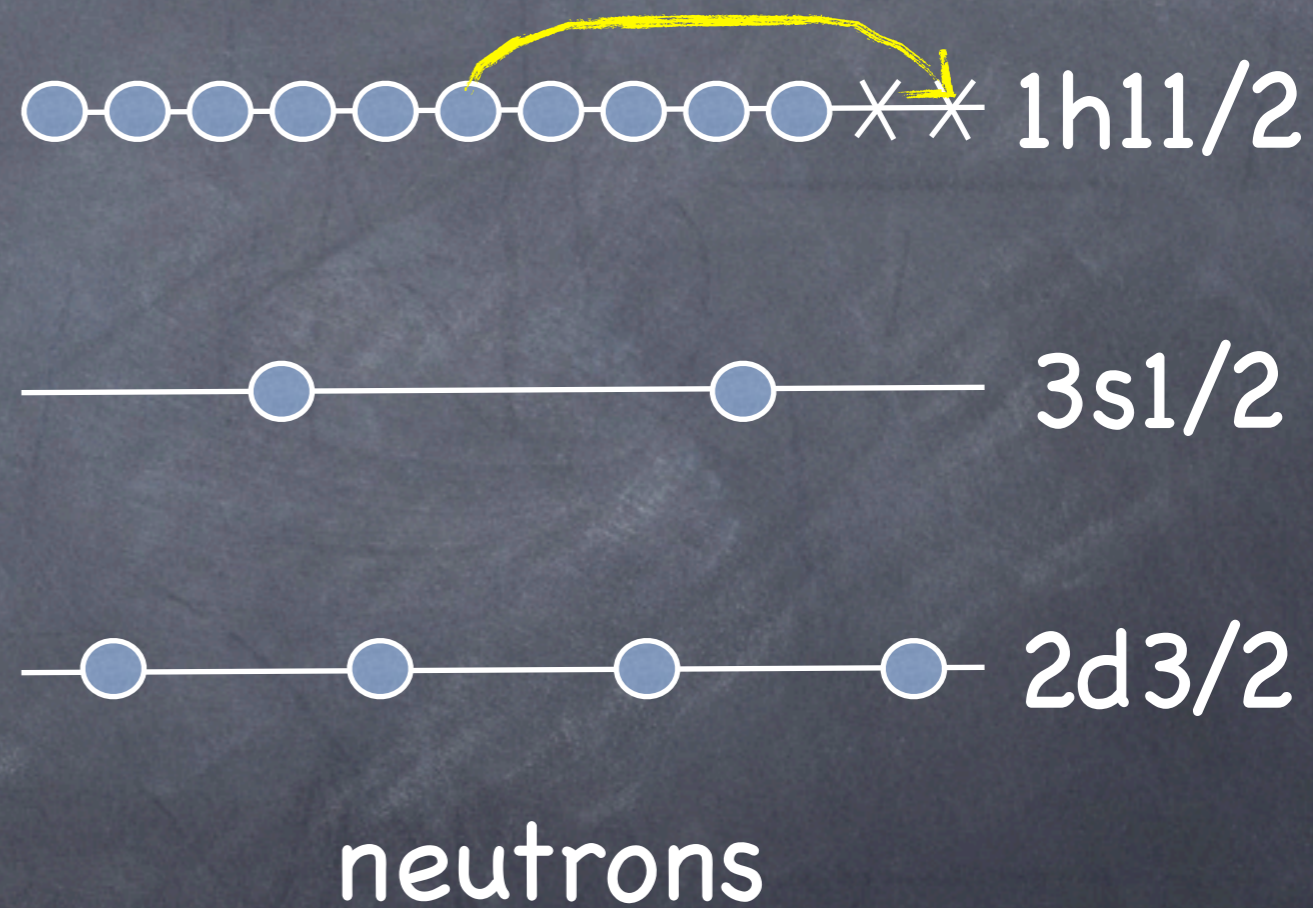
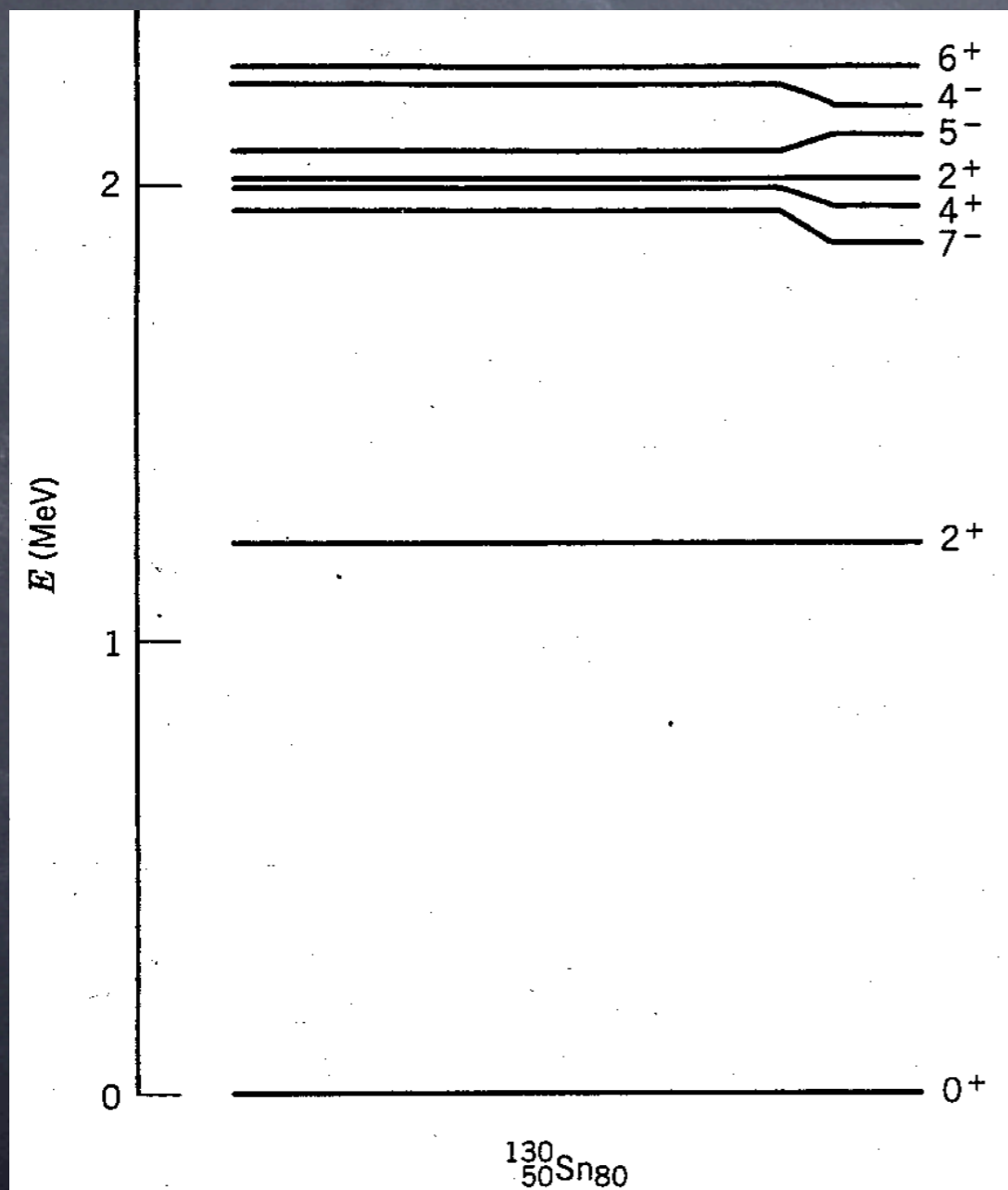
$1h_{11/2}$

$3s_{1/2}$

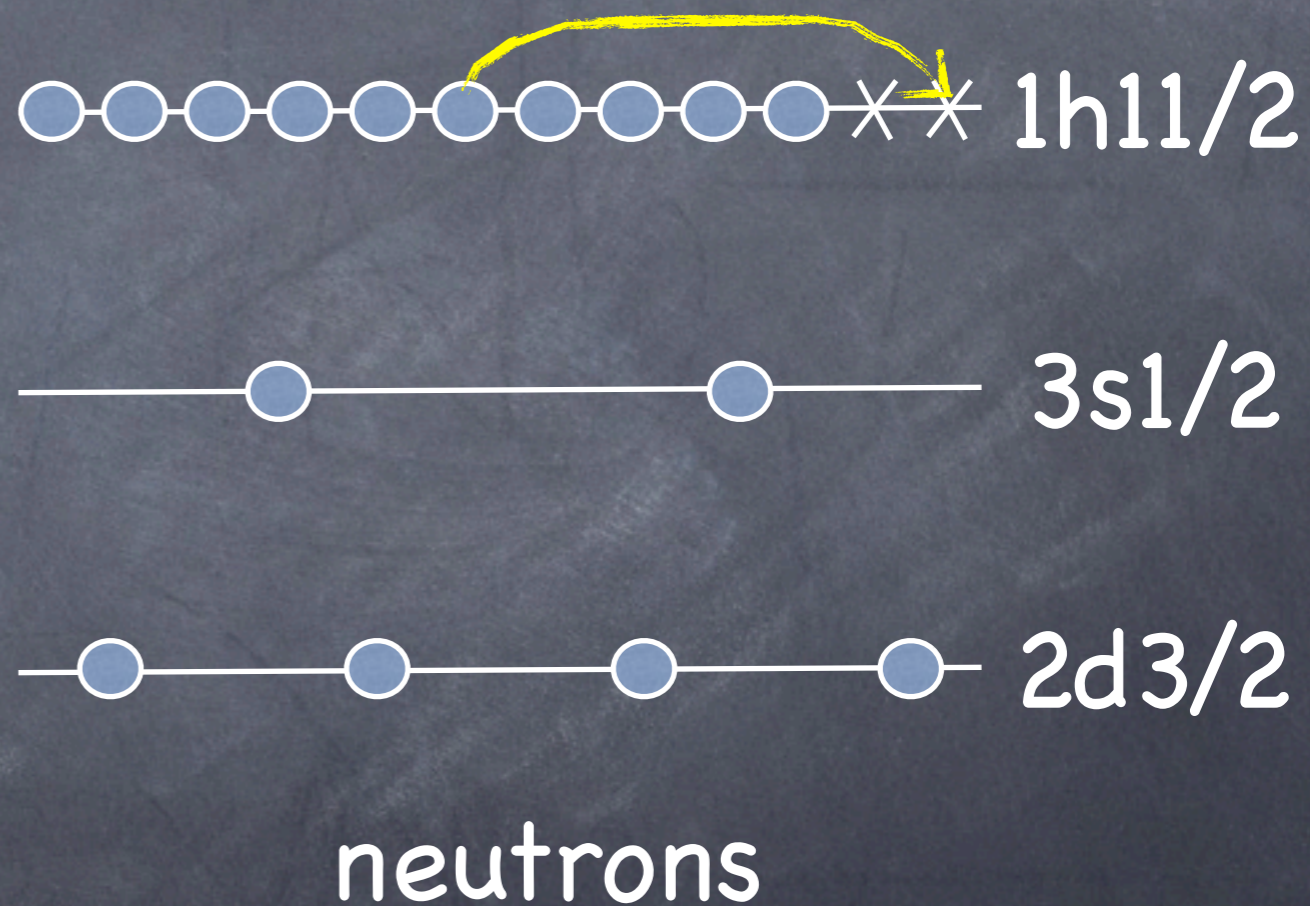
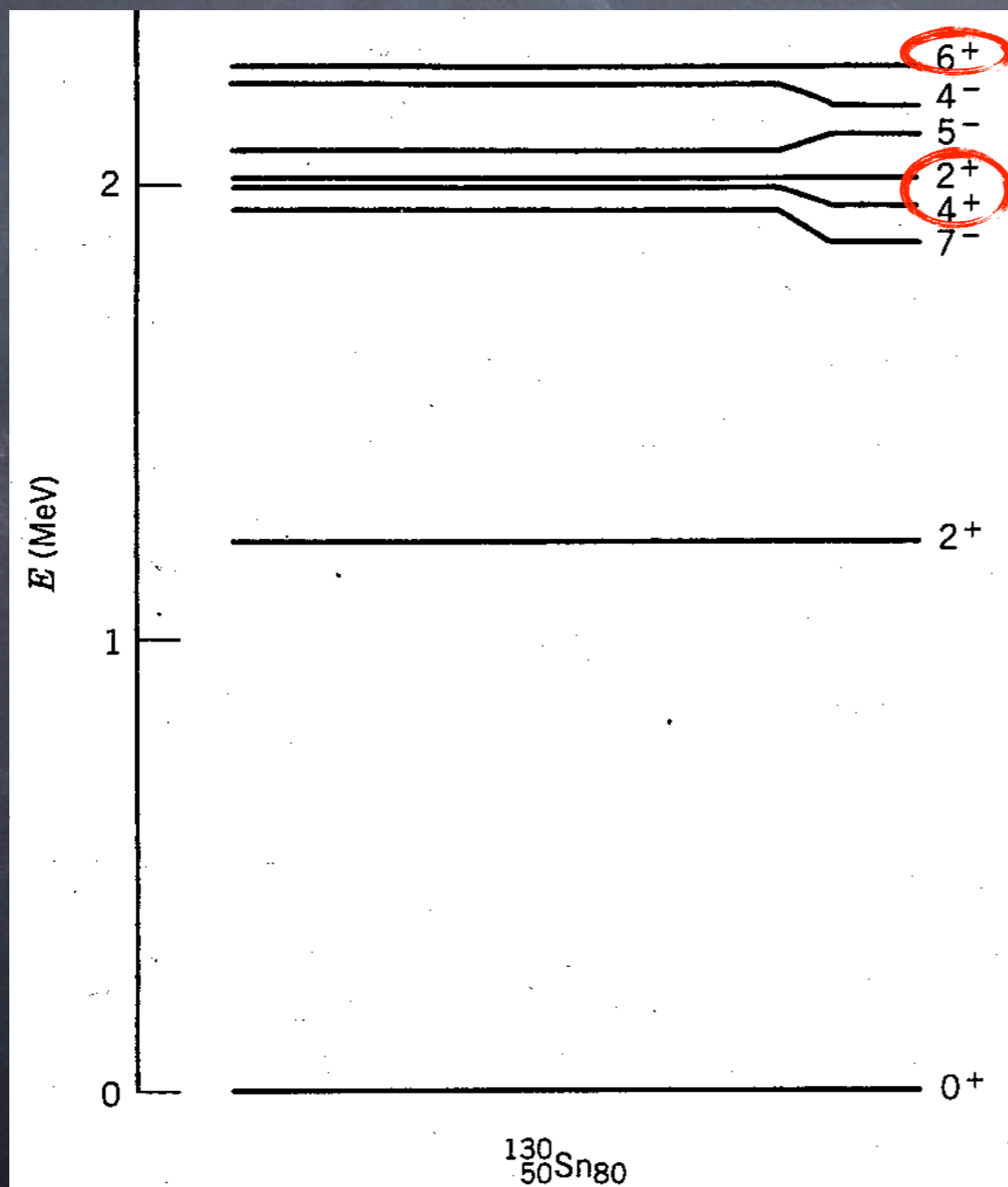
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neutrons

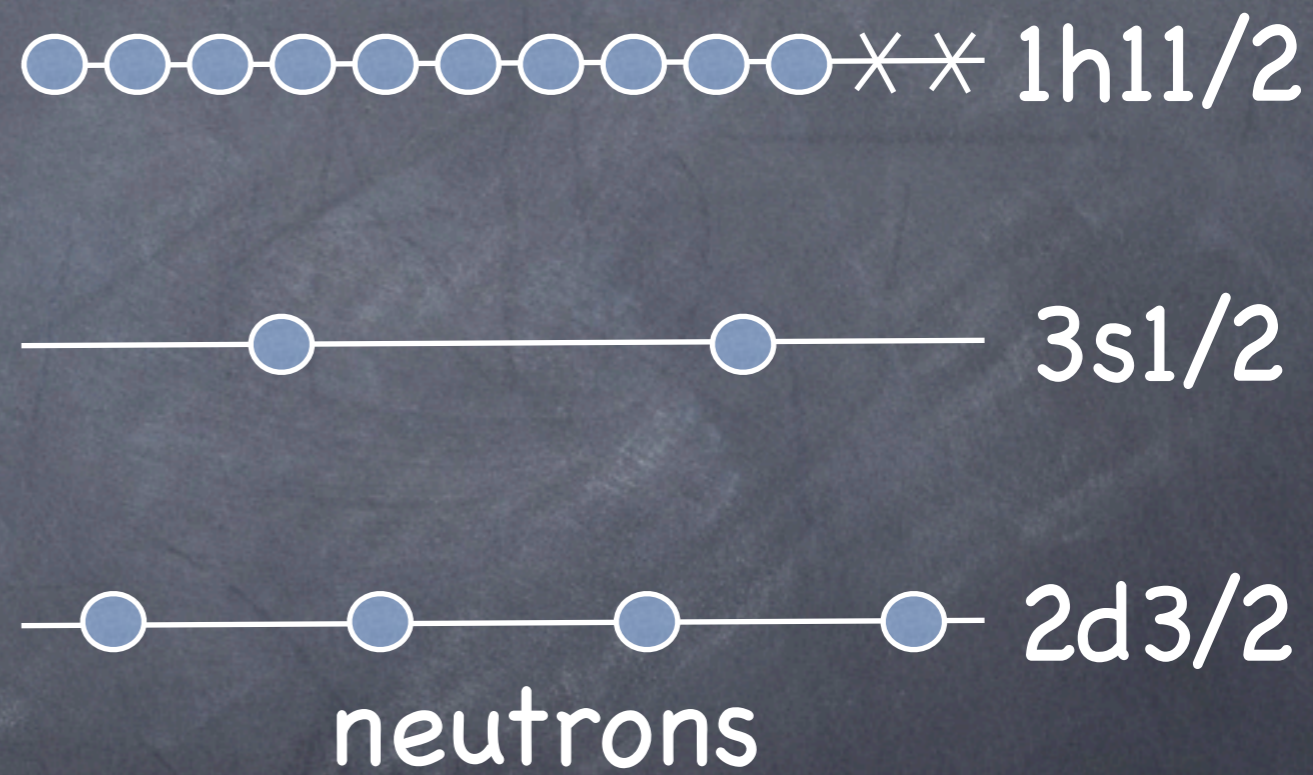
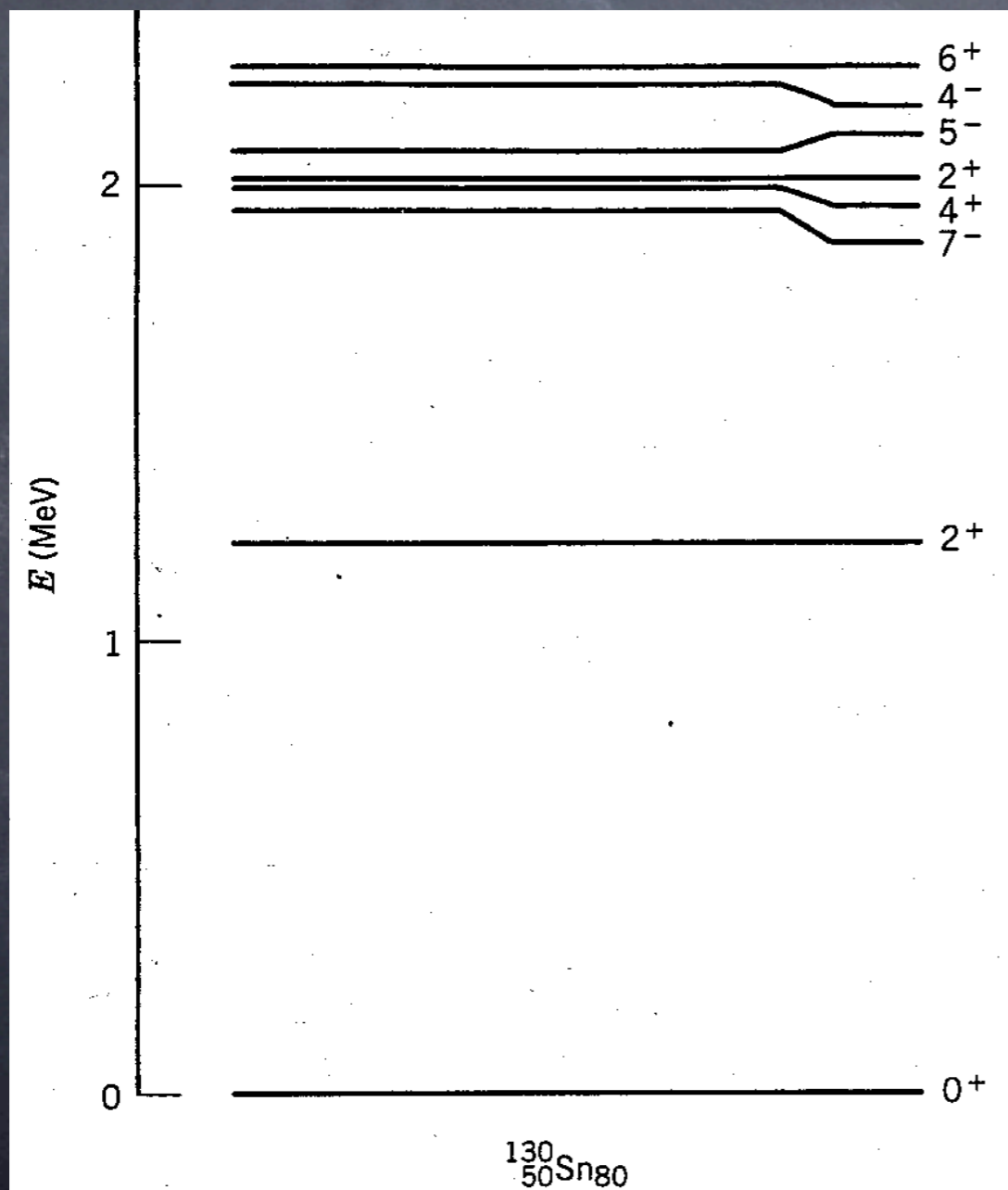
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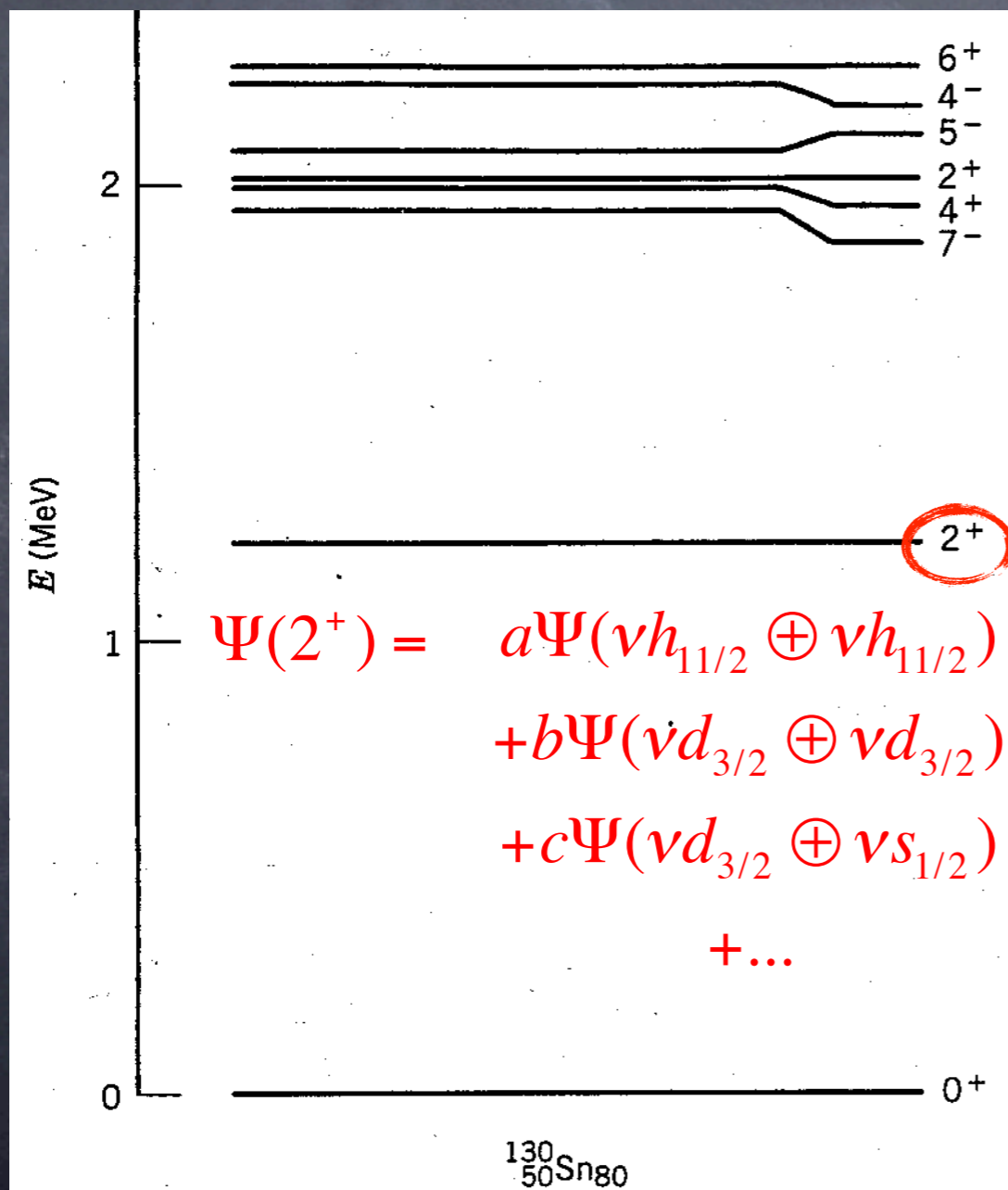
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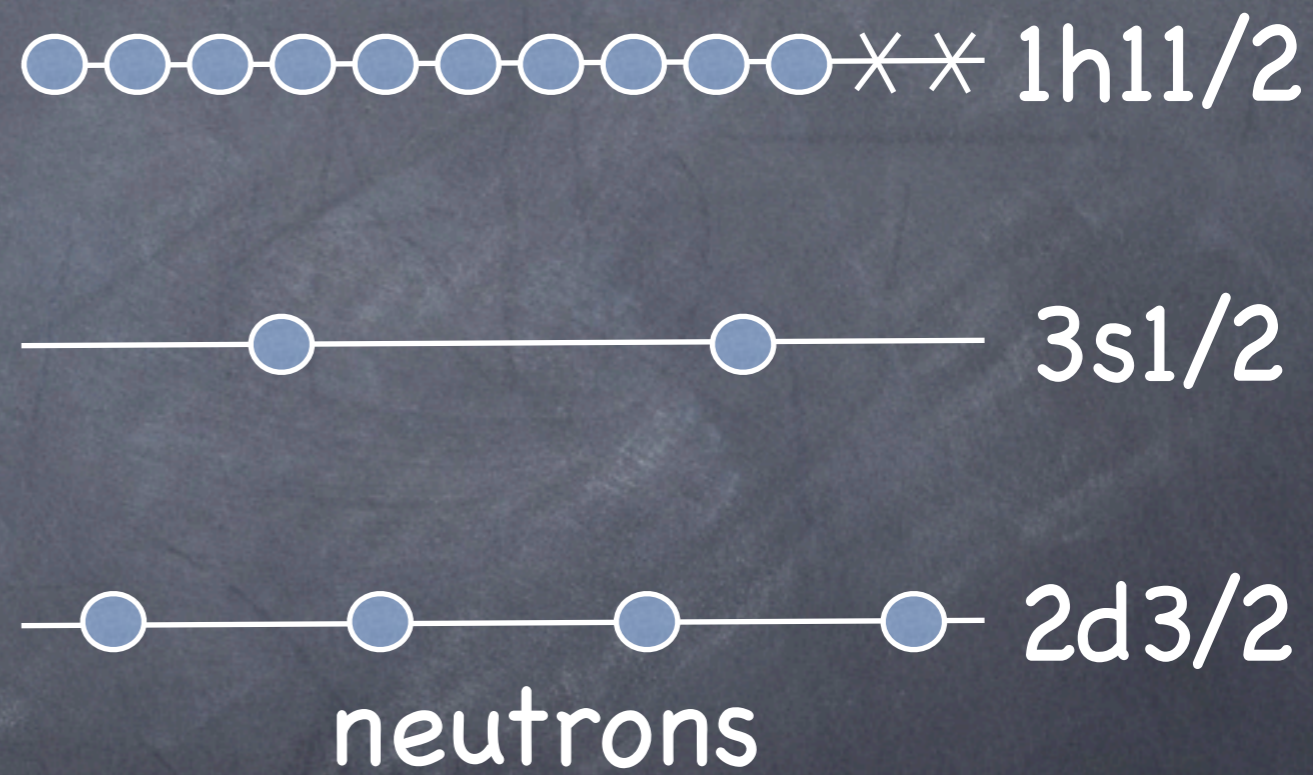
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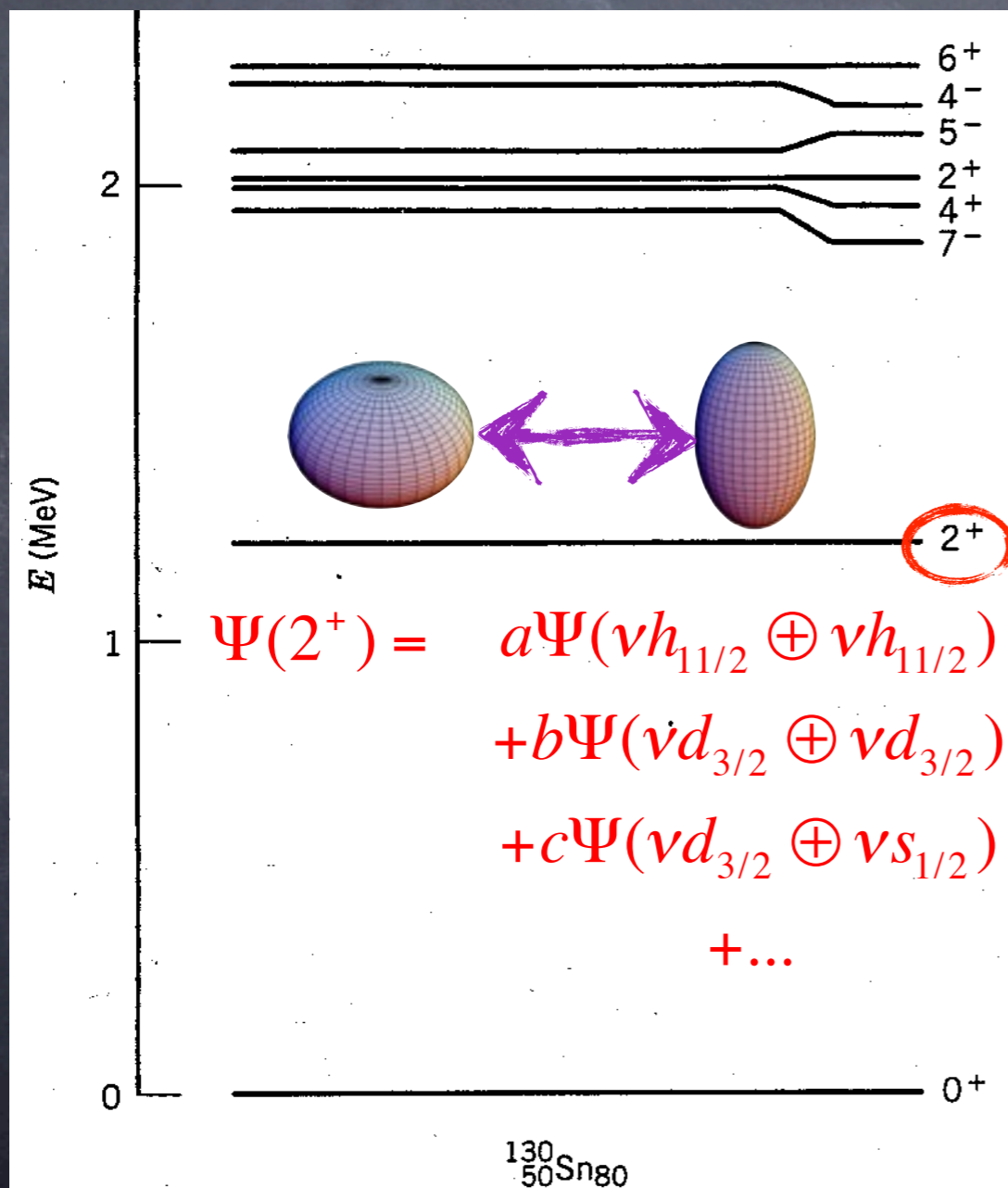
# Interpretation of $^{130}\text{Sn}$ spectrum



$$\Psi(2^+) = a\Psi(\nu h_{11/2} \oplus \nu h_{11/2}) + b\Psi(\nu d_{3/2} \oplus \nu d_{3/2}) + c\Psi(\nu d_{3/2} \oplus \nu s_{1/2}) + \dots$$



# Interpretation of $^{130}\text{Sn}$ spectrum



$\circ \circ \circ \circ \circ \circ \circ \circ \circ \circ \times \times$   $1h_{11/2}$

$\circ \circ$   $3s_{1/2}$

$\circ \circ \circ \circ$   $2d_{3/2}$

neutrons

$\Rightarrow$  low-lying collective vibration  $\neq$  HF eigenstate